

The TARANIS mission

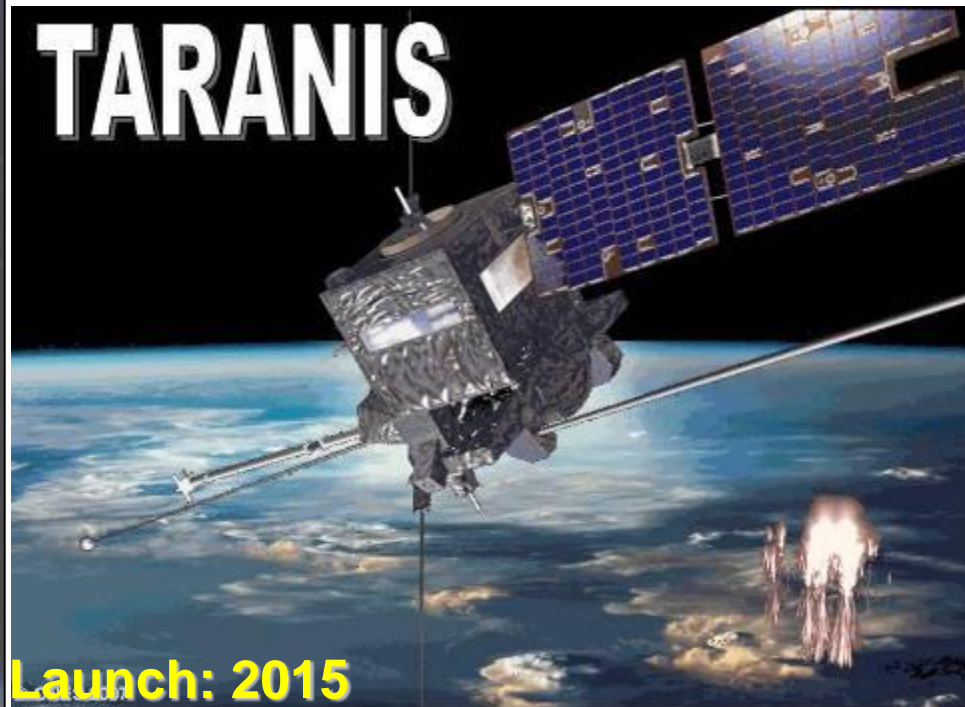
J-L. Pinçon¹; E. Blanc²; **P-L. Blelly**³; F. Lebrun⁵; M. Parrot¹; J-L. Rauch¹; J-A. Sauvaud⁴; E. Seran⁴

1. LPC2E, CNRS/University of Orléans, France - 2. DASE/LDG, CEA, France - 4. LATMOS, CNRS, France
3. **IRAP, CNRS/University of Toulouse**, France - 5. APC, CNRS/University of Paris 7, France



Tool for the Analysis of RAdiation from lightNing and Sprites

Motivation: Atmosphere-ionosphere-magnetosphere couplings



Launch: 2015

- Dimension: $\sim 1\text{m}^3$
- Mass of the Scientific Payload: $\sim 35\text{ Kg}$

TARANIS is a satellite dedicated to the study of impulsive energy transfers between the atmosphere of the Earth and the space environment.

Orbit:

- Sun-synchronous

local time 22:30 \rightarrow 02:00

- Inclination: 98°
- Altitude: 700 km

Time stamping accuracy: $\pm 1\text{ ms}$

Pointing accuracy
- localization: 5 km

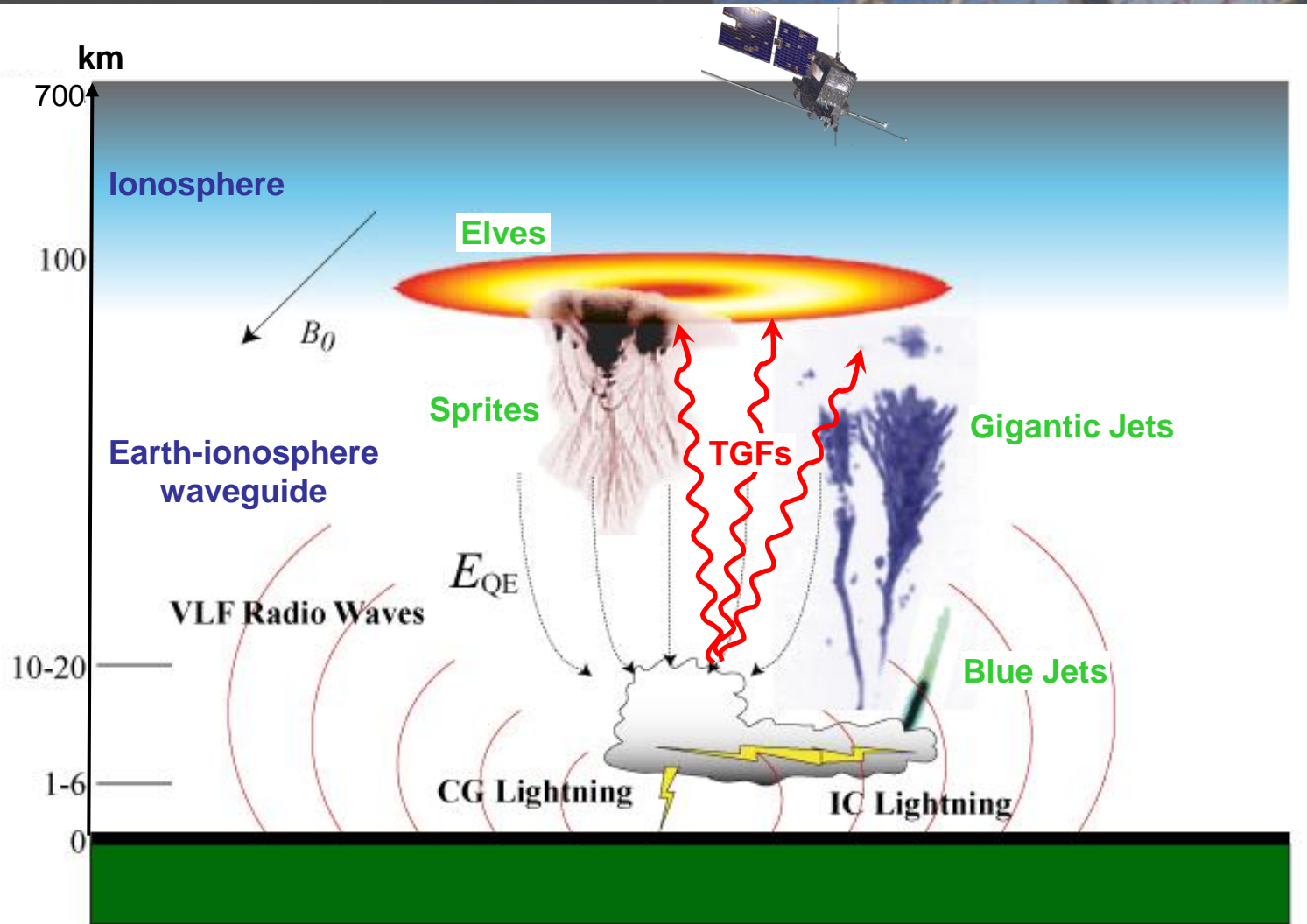
Subsystems:

- mass memory: 16 Gbits
- X band telemetry: 16.8 Mbits/s

Data: 4 GB/day

- Combined Nadir observations of TLEs and TGFs.
- High resolution measurement of energetic electrons.
- Wave field measurements over the frequency range [DC - 35 MHz].

TLEs and TGFs above thunderclouds



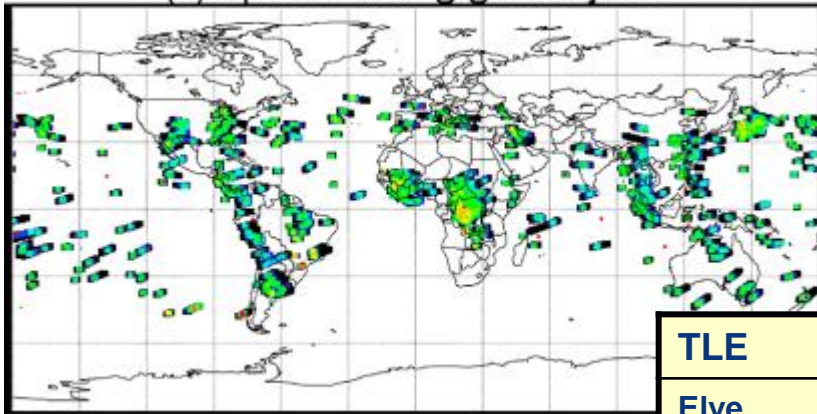
Elves: 1 ms
E&M pulse

Halo: 1-5 ms:
glow discharge
Sprite: 1-10 ms;
streamer discharge

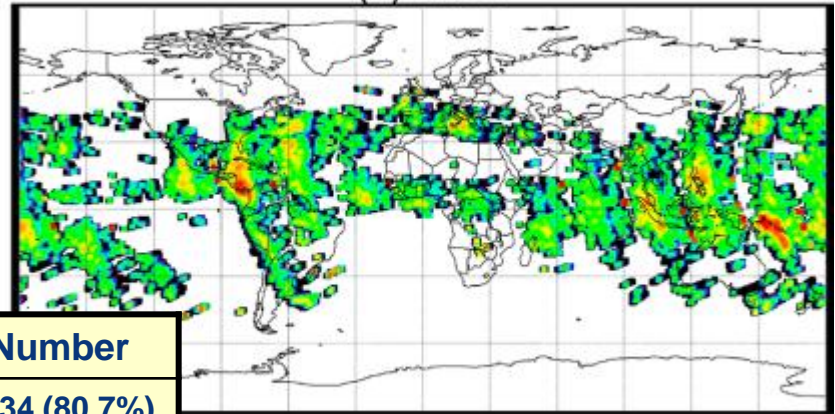
Blue jets:
5-1000ms
process ?

Transient Luminous Events (TLEs)

(a) sprites and gigantic jets

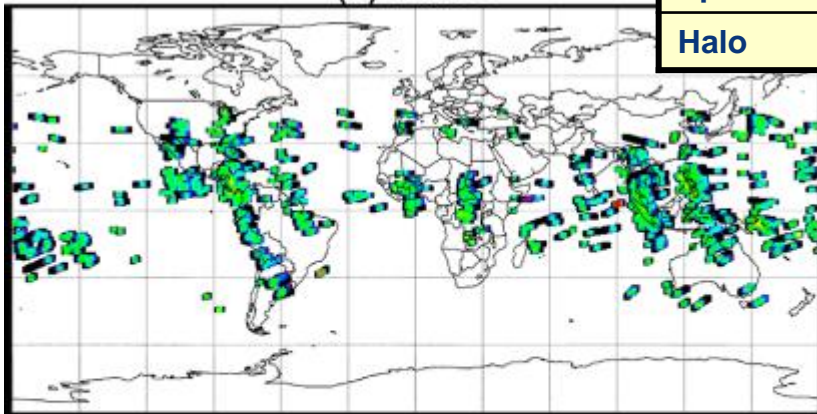


(b) elves

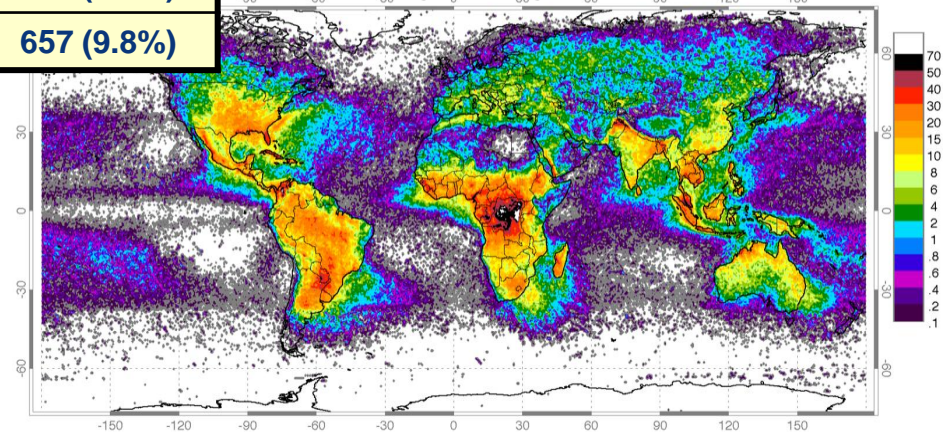


TLE	Number
Elve	5434 (80.7%)
Sprite	633 (9.4%)
Halo	657 (9.8%)

(c) halos



(d) lightning



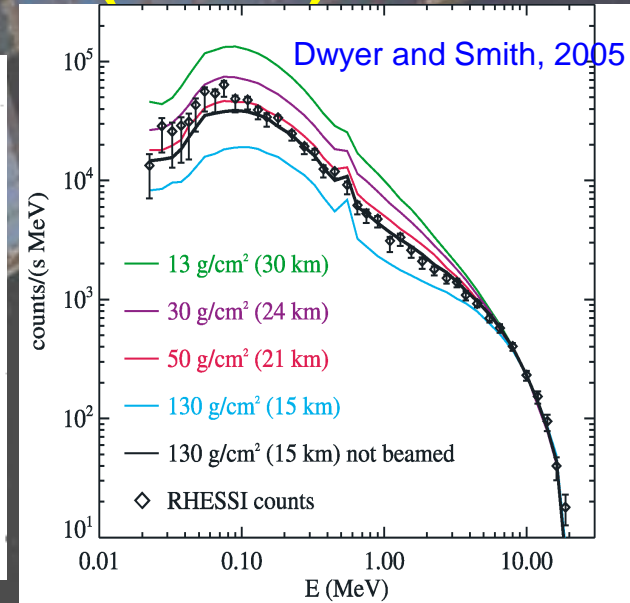
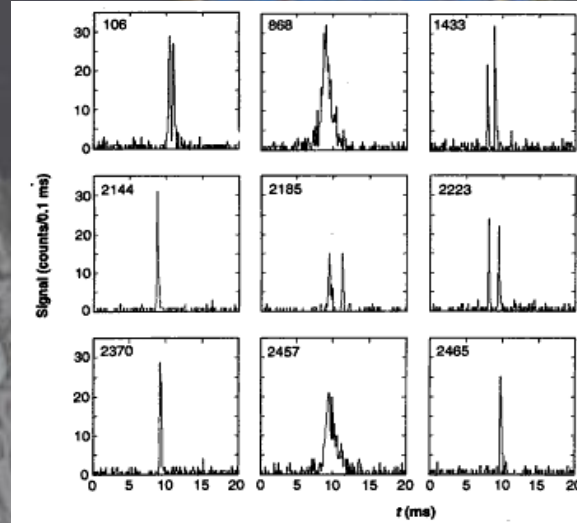
$10^{-4.5}$ $10^{-4.0}$ $10^{-3.5}$ $10^{-3.0}$ $10^{-2.5}$ $10^{-2.0}$ $10^{-1.5}$ $10^{-1.0}$ $10^{-0.5}$ (#/yr/km²)

a), b) c) from Chen et al. (2008). d) : from Christian et al., (2003).

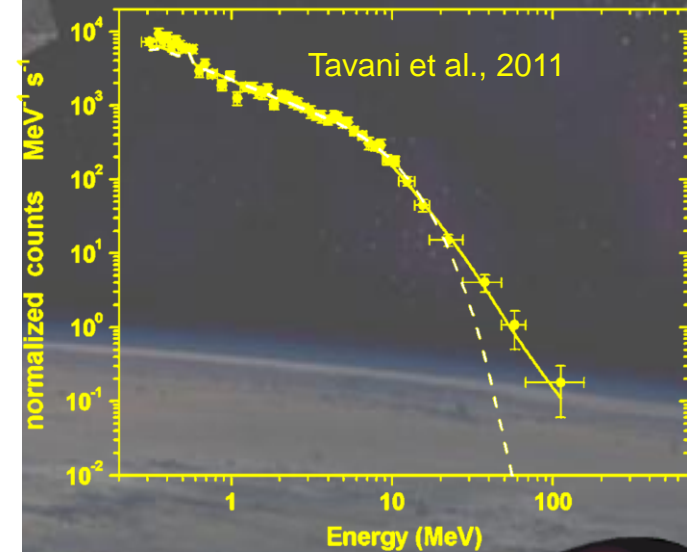
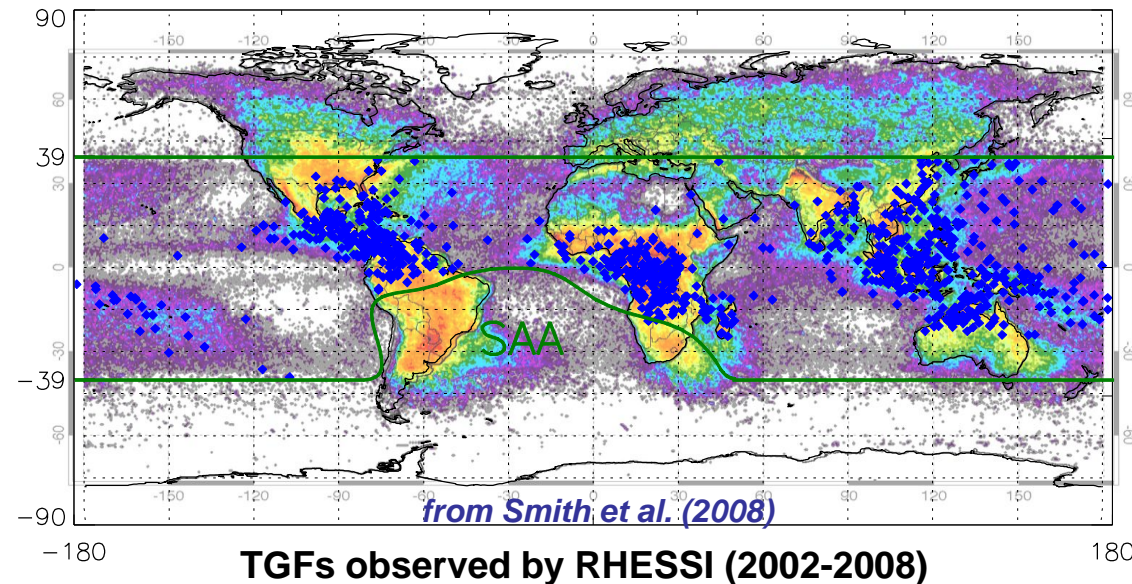
TLEs observed by ISUAL (July 2004- June 2007)



Terrestrial Gamma ray Flashes (TGFs)

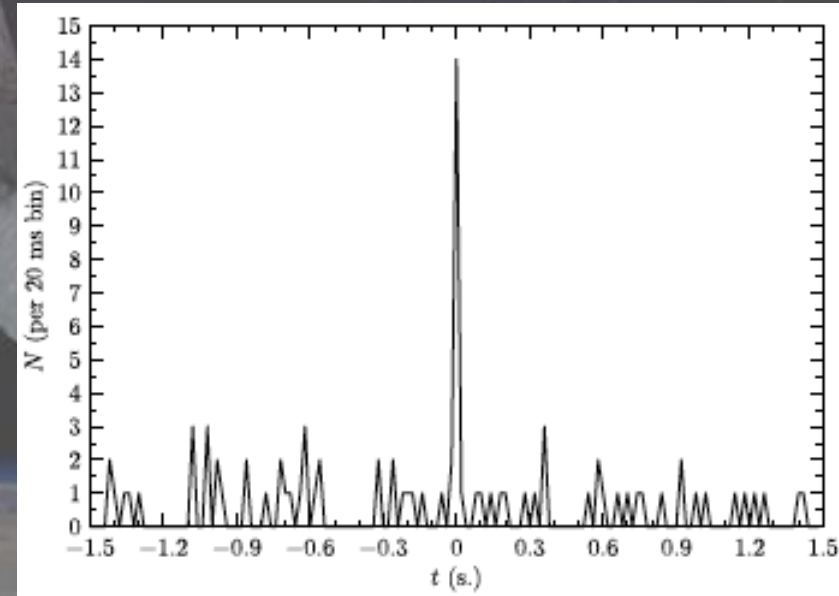


On average 27 events/week expected on board TARANIS



TGF electrons

- Brief electron pulses: 2-10 ms
- About 1 e/cm^2 spread over 50 km wide
- +/- along the magnetic field, coming from the Earth, and possibly reflected from the conjugate hemisphere
- Only 1 spacecraft with adequate performances
 - ◆ SAMPEX: $60 \text{ cm}^2 \cdot \text{ster}$, 20 ms



- Expected counts during 10 ms: 50 max
- High energy limit, at least 4 MeV (10 MeV)

November 22, 2002
Lat/Long:-2.65 N/166.95, L=1.1

Scientific objectives of TARANIS

- **To advance the physical understanding of the links between TLEs, TGFs and environmental conditions** (*lightning activity, geomagnetic activity, atmosphere/ionosphere coupling, occurrence of Extensive Atmospheric Showers, etc*).
- **To identify the signatures associated with these phenomena** (*electron beams, associated electromagnetic or/and electrostatic fields*) **and to provide inputs to test generation mechanisms.**
- **To provide inputs for the modelling of the effects of TLEs, TGFs and bursts of precipitated and accelerated electrons** (*lightning induced electron precipitation, runaway electron beams*) **on the Earth's atmosphere.**

TARANIS scientific payload

■ MEXIC (Multi Experiment Interface Controller)

- electronics equipment to power and to manage the whole scientific payload
- Alert management

■ MCP including MC (MicroCameras)

- a set of two cameras (a sprite camera and a lightning camera), and their associated analyzer, and PH (Photometers);
- a set of four photometers and their associated analyzer;

■ XGRE (X-ray, Gamma-ray and Relativistic Electron experiment)

- a set of three X and γ detectors and their associated analyzer;

■ IDEE (Instrument Détecteurs d'Electrons Energétiques)

- two electron detectors with one analyzer per detector;

■ IMM (Instrument de Mesure du champ Magnétique)

- a compound triaxial system of search-coil magnetometers to measure the alternative magnetic field in the low and medium frequency ranges and the medium frequency wave analyzer (electric and magnetic);
- IMM also includes a Sferics Detector (SD) to make on-board characterization of whistlers;

■ IME-BF (Instrument de Mesure du champ Electrique Basse Fréquence)

- a sensor to measure the electric field in the low frequency range and the low frequency wave analyzer (electric and magnetic);
- IME-BF also includes an Ion Probe (SI) to determine fluctuations of thermal plasma;

■ IME-HF (Instrument de Mesure du champ Electrique Haute Fréquence)

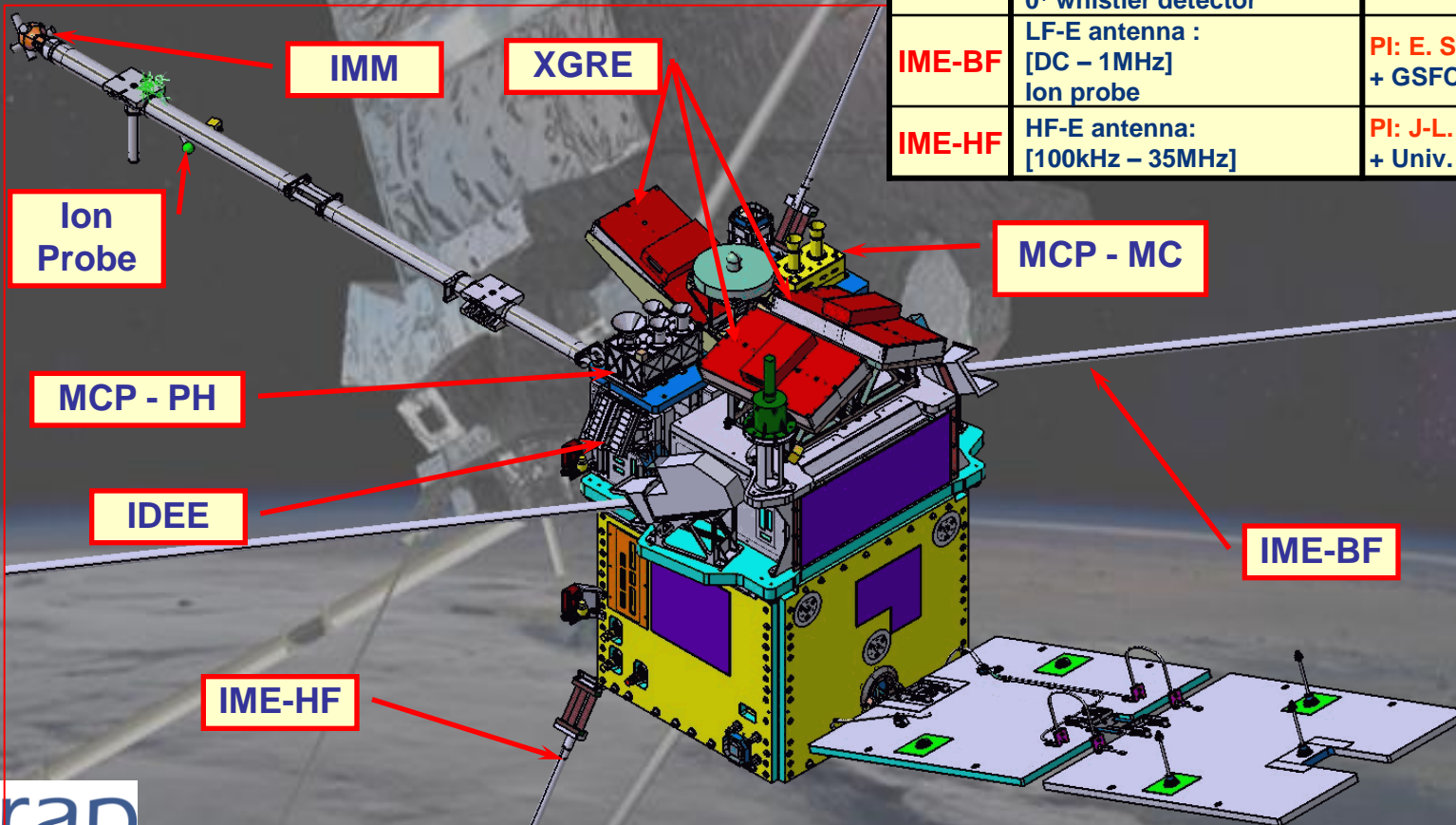
- a sensor to measure the electric field in the high frequency range and the associated high frequency wave analyzer;

Scientific payload accommodation

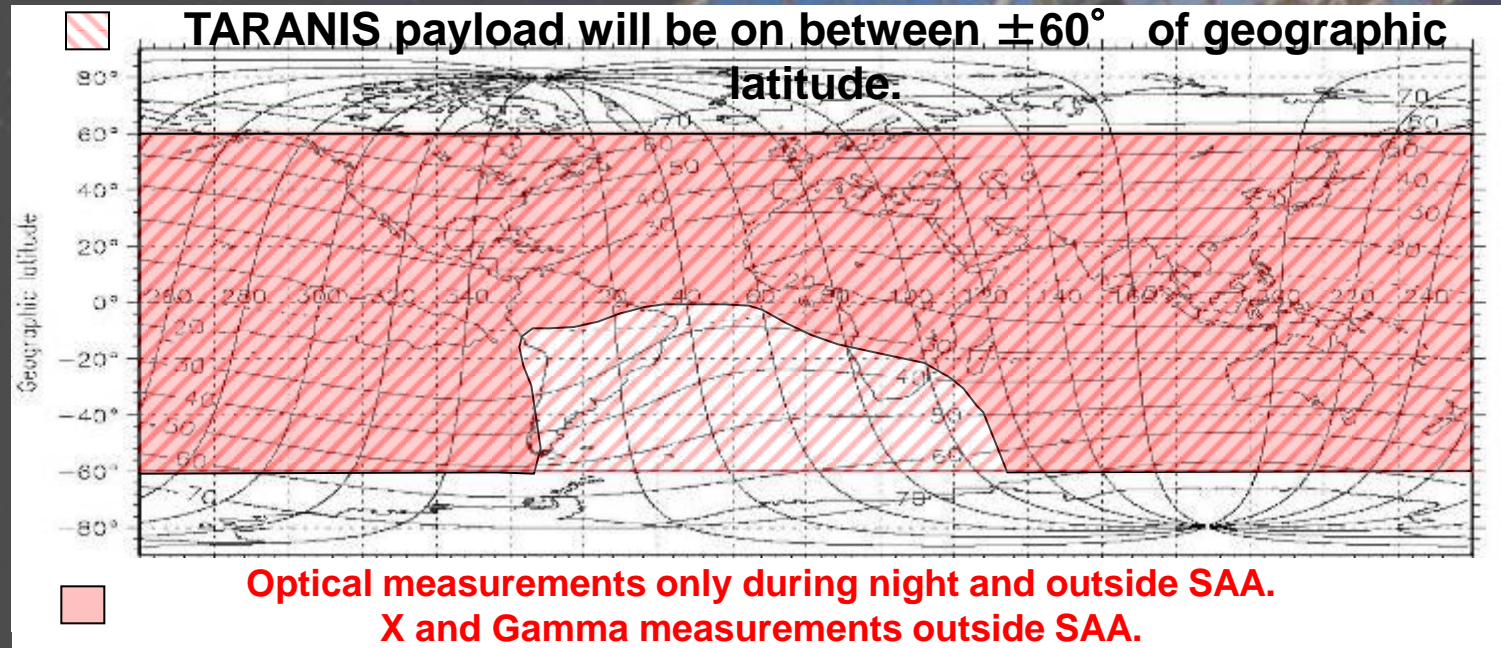
TOWARDS
THE EARTH



MCP	Lightning micro-camera TLE micro-camera Photometers	PI: E. Blanc, CEA (F) + Univ. Hokkaido (J)
XGRE	X and γ detectors: Photons : [20keV – 10MeV] e^- : [1 MeV – 10 MeV]	PI: P-L. Blelly, IRAP (F) + APC (F)
IDEE	Two e^- detectors: [70keV – 4MeV]	PI: J-A. Sauvaud, IRA (F) + Univ. Prague (Cz)
IMM	Triaxial search coil : [5Hz – 1MHz] 0^+ whistler detector	PI: J-L. Pinçon, LPC2E (F) + Univ. Stanford (USA)
IME-BF	LF-E antenna : [DC – 1MHz] Ion probe	PI: E. Seran, LATMOS (F) + GSFC (USA)
IME-HF	HF-E antenna: [100kHz – 35MHz]	PI: J-L. Rauch, LPC2E (F) + Univ. Prague, IAP (Cz)



TARANIS: Event and Survey modes



Survey data:

Continuous monitoring of the background conditions.

2 GB of low resolution data per day!

Event data:

Triggered when a priority event is detected (TLE, TGF, electron beam, burst of electromagnetic/electrostatic waves), then all instruments record and transmit high resolution data.

2 GB of high resolution data per day!

TARANIS event data (1/2)

TARANIS

Mass memory: 16 Gbits

X-band telemetry: 16.8 Mbits/s

2 GBytes of event data per day

- On average 12 events per half-orbit ($T=100\text{mn}$)
- A maximum of 36 events per half-orbit

4 triggering instruments

MCP-PH

TLE alert

XGRE

TGF alert

IDEE

Electron alert

IME-HF

Wave alert

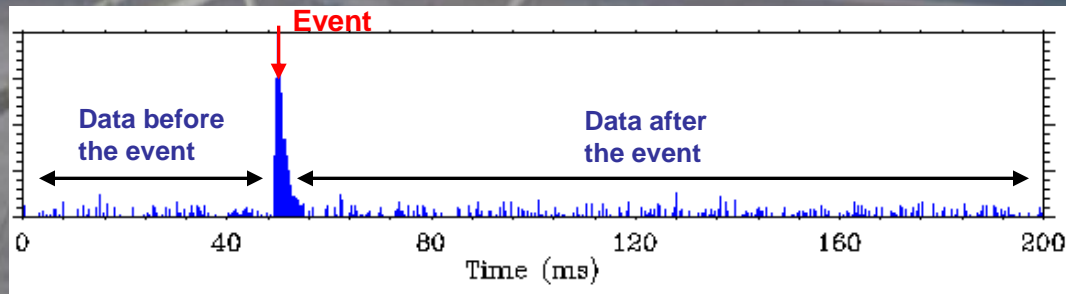
MEXIC

Event alert

ALL PAYLOAD INSTRUMENTS

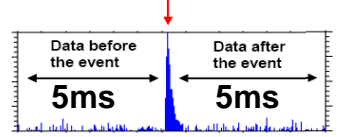
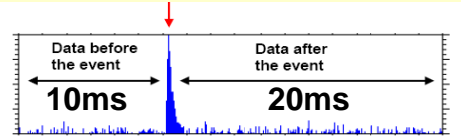
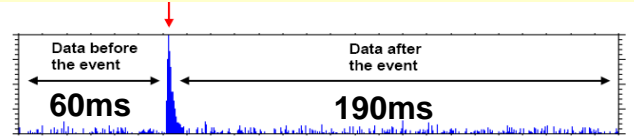
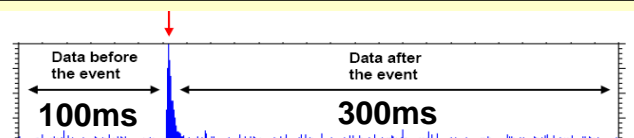
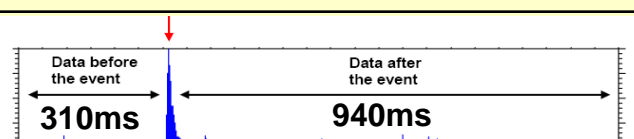
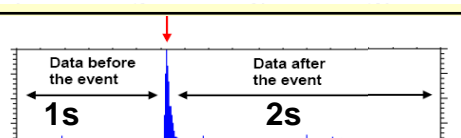
Multi EXperiment Interface Controller
to power and to manage the whole scientific payload.

On-board analyzers will include event buffer memory sized to record high resolution data both before and after the trigger



TARANIS event data (2/2)

**TGF event
triggering**

XGRE		photons
IME-HF (HF)		Fe = 80MHz
IME-IMM (LF-MF)		Fe = 2MHz
MCP-PH		Fe = 20kHz
MCP-MC	<div>Nadir Image TLEs, lightning</div> <div>Nadir Image TLEs, lightning</div> <div>Nadir Image TLEs, lightning</div> <div>Nadir Image TLEs, lightning</div>	500x500 $\Delta t \approx 100\text{ms}$
IME-IMM (ELF-VLF)		Fe = 50kHz
IDEE		electrons

Time stamping accuracy

Absolute accuracy:

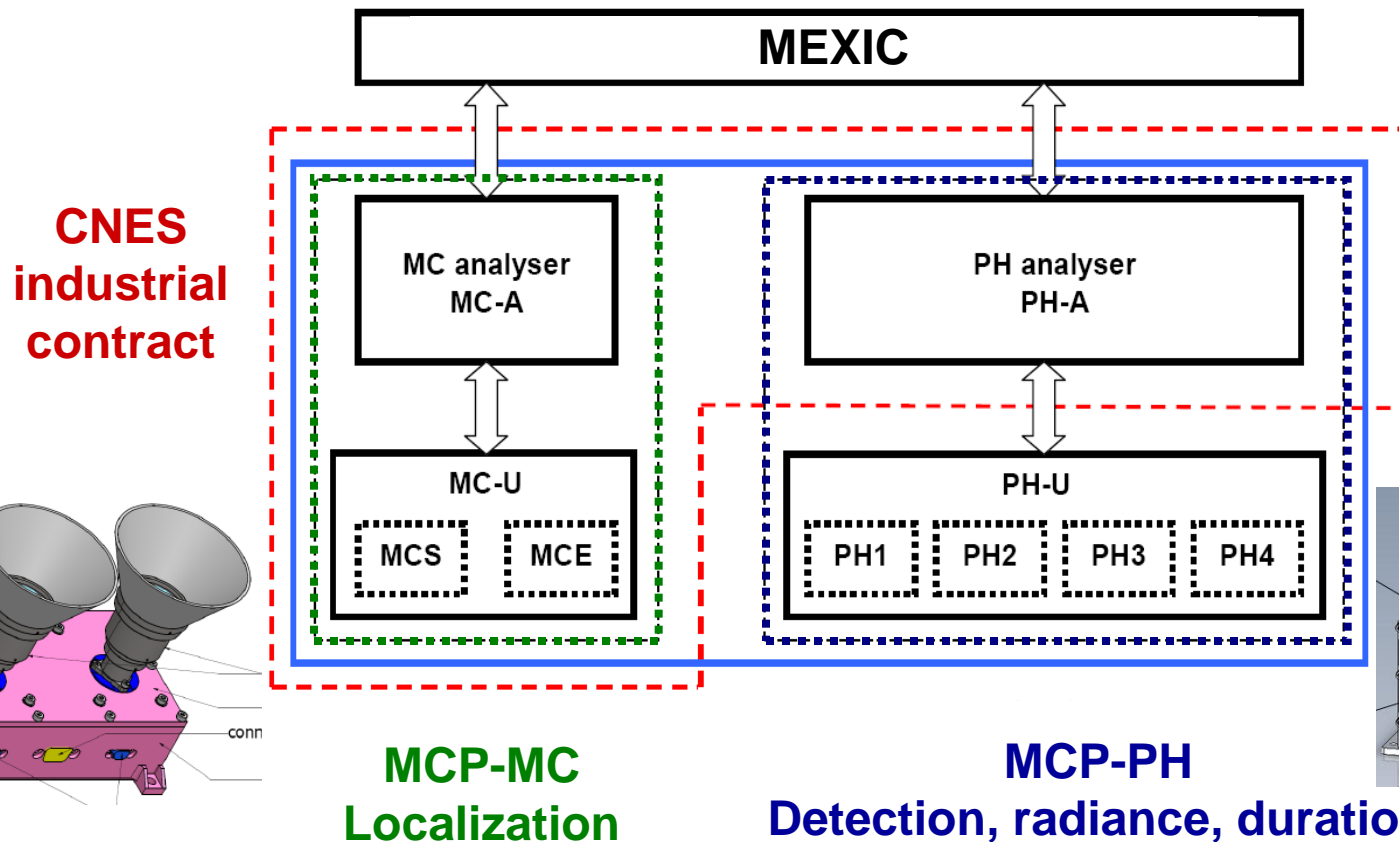
1 ms for comparison with ASIM, balloon, and ground based measurements.

Relative accuracy:

10 μs for comparison between TARANIS experiments.

MCP: Instrument and team

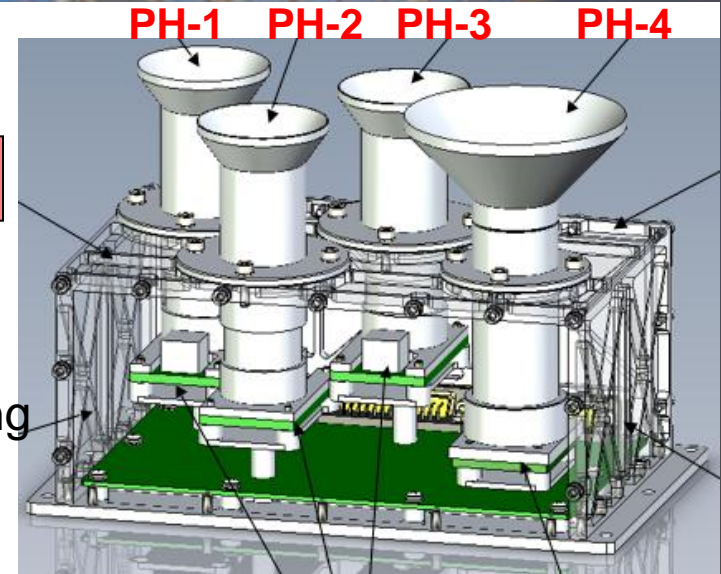
MCP: PI E. Blanc (CEA)
MCP-MC: E. Blanc and Th. Farges (CEA)
MCP-PH: Lead Co-I M. Sato (Hokkaido Univ.)



MCP- PH Units

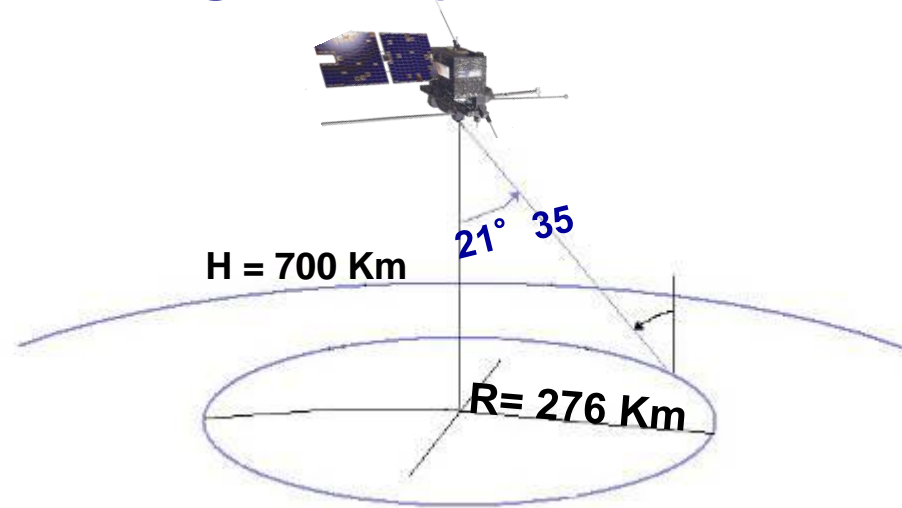
➤ Spectral bands

• PH1: UV-C	150 - 280 nm	Sprite
• PH2: UV-A	337 ± 5 nm	Sprite
• PH3: NIR	761 ± 5 nm	Sprite
• PH4: Red - NIR	600 – 900 nm	Lightning



➤ Field Of View (FOV):

PH1 – PH2 – PH3: "small" FOV ($\pm 21^\circ$ 35 $\Rightarrow R = 276$ Km)
 PH4 : "large" FOV ($\pm 43^\circ$ 4 $\Rightarrow R = 700$ Km)



XGRE instrument requirements

■ Speed

- ♦ High count rate per cm^2 ($10^4 - 10^5 \text{ cm}^{-2} \text{ s}^{-1}$)
 - Dead time $< 300 \text{ ns}$ (10^6 s^{-1})
- ♦ Datation
 - triggers with 300 ns relative time accuracy
 - events (TGFs) with 1 μs relative time accuracy

■ Detection

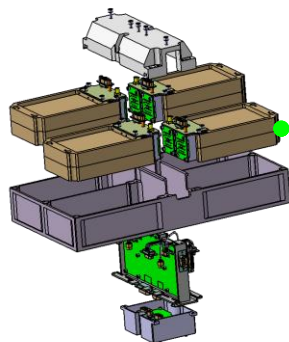
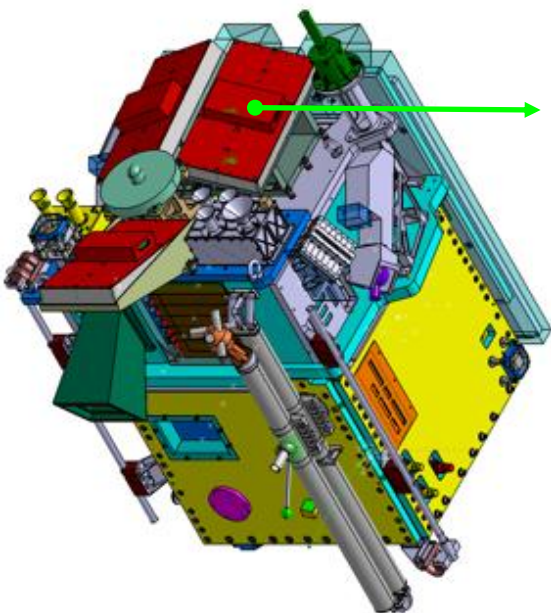
- ♦ Separation between γ -rays and electrons
- ♦ γ -rays
 - Energy range
 - 20 keV-12 MeV
 - 30 % accuracy at 20 keV
 - 9 % accuracy at 511 keV
- ♦ Electrons
 - Energy range
 - $<1 \text{ MeV} - >10 \text{ MeV}$
- ♦ minimize risk of missing TGF
 - Storage capacity of 200 000 photons
 - burst and survey algorithms

■ Localization

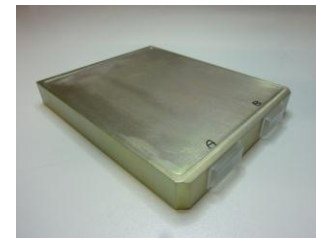
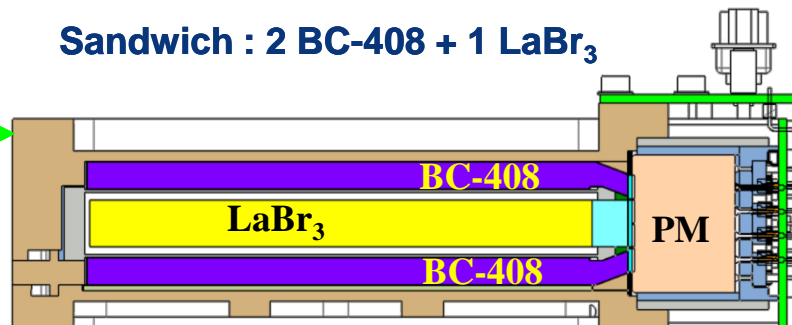
- ♦ Ascending or descending particle motion
- ♦ Zenithal and azimuthal direction ($\sim 30^\circ$ resolution)

XGRE (PIs : P-L Blelly (IRAP) et F. Lebrun (APC))

Three sensors on the platform oriented towards Earth + 1 toward space.



Sandwich : 2 BC-408 + 1 LaBr₃



XGRE Instrument

- 3 sensors
- total detection surface ~ 900 cm²

XGRE Sensor

- 4 detection units
- AD converters
- 12 bits (LaBr₃)
- 10 bits (BC408)

LaBr₃ (photons)

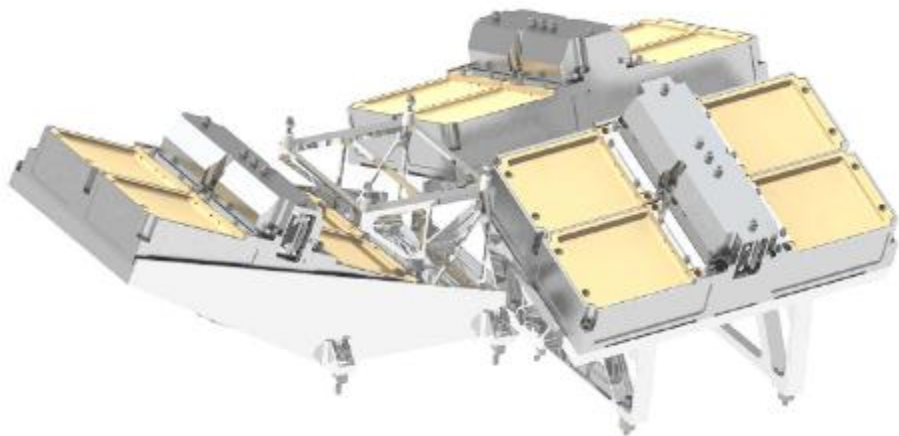
- Rapid (dead time < 300 ns)
- good linearity
- spectral resolution

BC-408 (electrons)

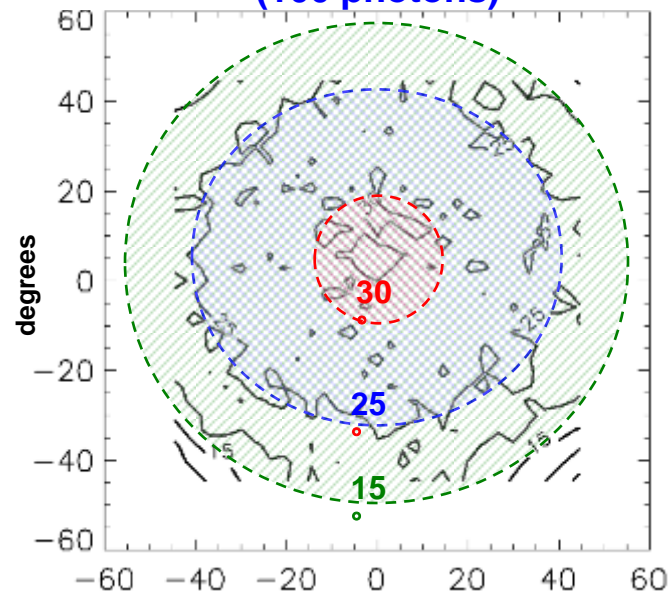
XGRE (PIs : P-L Blelly (IRAP) et F. Lebrun (APC))

TGF localization

- three planes ($\sim 20^\circ$ inclination)

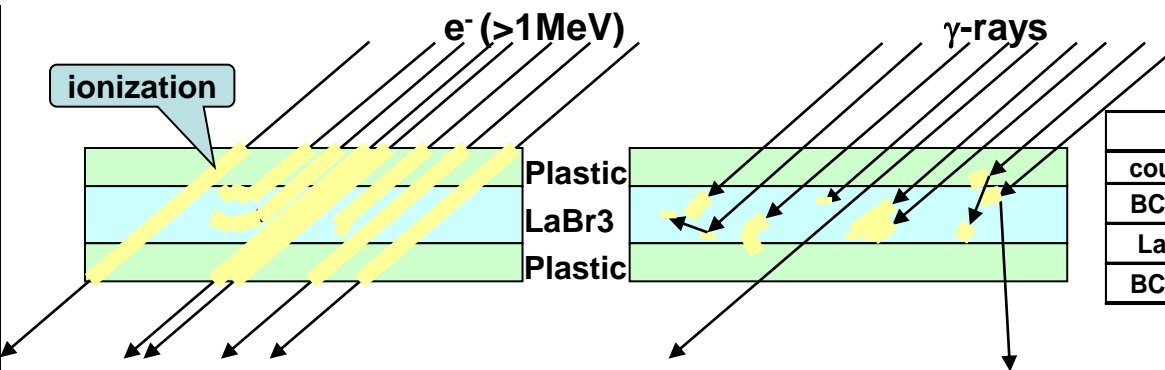


accuracy on the arrival direction
(100 photons)



Separation between electrons and photons

- Coincidence/anti-coincidence entre LaBr3 et BC408



couche	électrons			photons	
	< 1 MeV	1 - 8 MeV	> 8 MeV	<200 keV	>200 keV
BC-408	E	~ 1 MeV	~ 1 MeV	0	0
LaBr3	0	~ E-1 MeV	~ 8 MeV	E	≤ E
BC-408	0	0	~ E-9 MeV	0	0

IDEE (PI : J-A Sauvaud (IRAP))

■ 2 spectrometers

- ♦ nadir
- ♦ zenith

■ Energy range

- ♦ 60 keV – 5 MeV

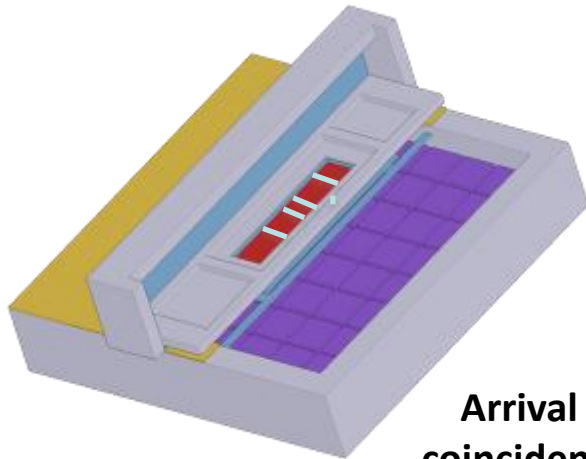
$\Delta E = 16 \text{ keV}$ [60 keV – 500 keV]

$\Delta E = 60 \text{ keV}$ [500 keV – 5 MeV]

■ 8 angles

Field of view = $150 \times 40^\circ$ [60 keV – 500 keV]

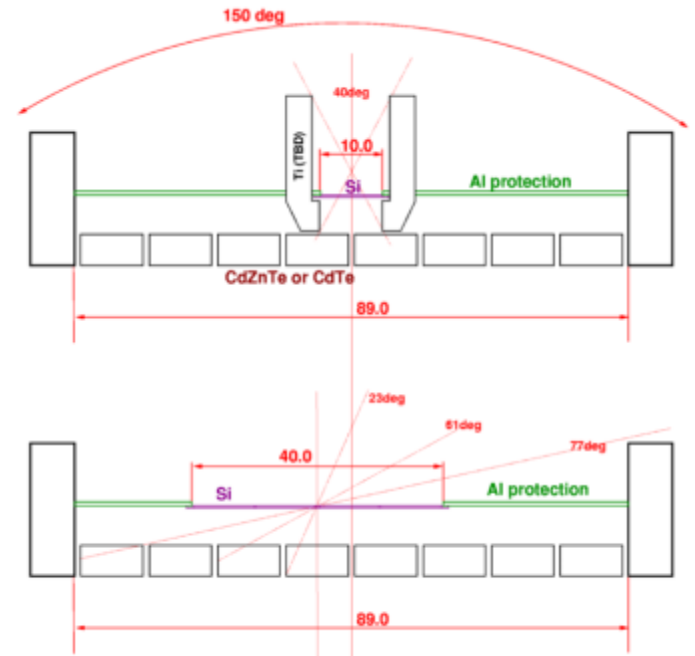
Field of view = $150 \times 150^\circ$ [500 keV – 5 MeV]



■ Silicium, 5 cells

■ CdTe, 64 cells

Arrival direction derived from the coincidence between Si and CdTe cells



■ Triaxial search coil magnetometer

■ 2 mono band antennas

- ◆ 10Hz-20kHz frequency range

■ 1 double-band antenna

- ◆ 10Hz-20kHz and 10kHz-1MHz frequency range

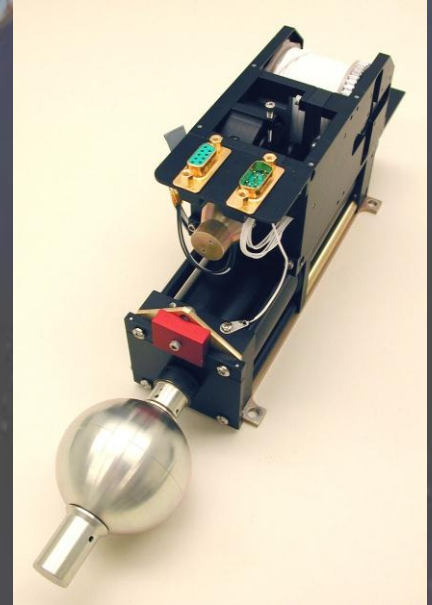


■ Two sphere sensors on booms

- ◆ 1 component of the electric field
- ◆ from DC to 3.3 MHz
- ◆ ULF frequency range: 0 – 64 Hz
- ◆ VLF frequency range from few Hz to 20 kHz,
 - frequency resolution: 94 Hz

■ Langmuir probe

- ◆ Current due to the ion/electron thermal plasma
- ◆ Plasma density range: 10^2 - 10^7 cm⁻³
- ◆ frequency: 128 Hz

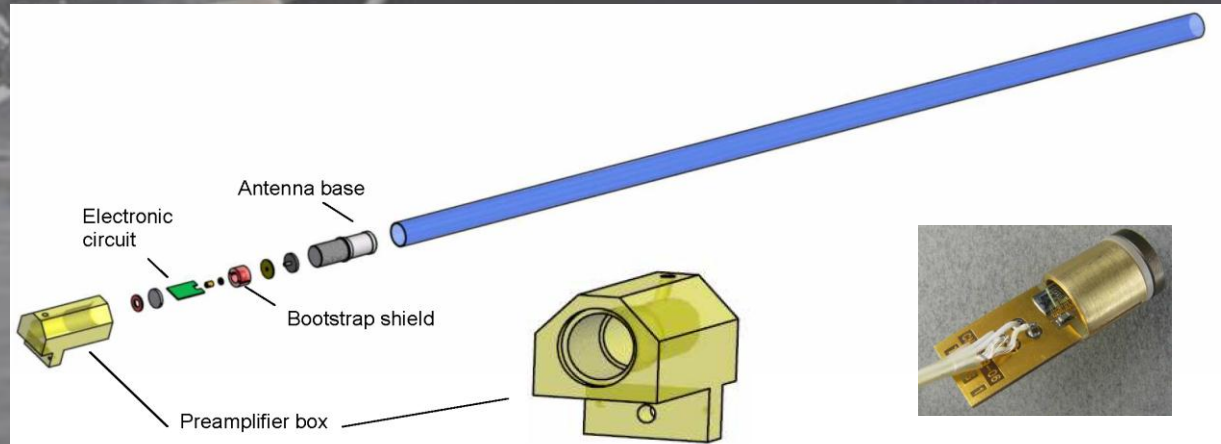


■ Double-wire Hertz dipole

- ◆ aligned with a satellite diagonal
- ◆ on the opposite side of the solar panel

■ 1 component of the electric field

- ◆ from 100kHz to 30MHz



TARANIS STATUS

		Data access					
		Mission PI	Instrument PI	Instrument Lead Co-I	Instrument Co-I	Guest Investigator	Public
Data & Products	Raw Survey data	Yes	instrument	instrument PI agreement required	No	No	No
	Raw Event data	Yes	instrument	instrument PI agreement required	No	No	No
	Calibrated Survey data	Yes	Yes	Yes	instrument	CST agreement required	No
	Calibrated Event data	Yes	Yes	Yes	instrument	CST agreement required	No
	Quickview Survey	Yes	Yes	Yes	Quicklook + instrument	CST agreement required	Quicklook only
	Quickview Event	Yes	Yes	Yes	Quicklook + instrument	CST agreement required	No
	Plot Survey data	Yes	Yes	Yes	instrument	CST agreement required	No
	Plot Event data	Yes	Yes	Yes	instrument	CST agreement required	No
	Auxiliary data	Yes	Yes	Yes	Yes	Yes	No

Launcher:
SOYUZ



Launch from KOUROU
by the end of 2015

TARANIS Planning

2011		2012	2013	2014	2015
Phase A & B		Phase C (Engineering model)		Phase D (Flight model)	AIT

Context: TARANIS – ASIM – GLIMS – COBRAT and Chibis-M



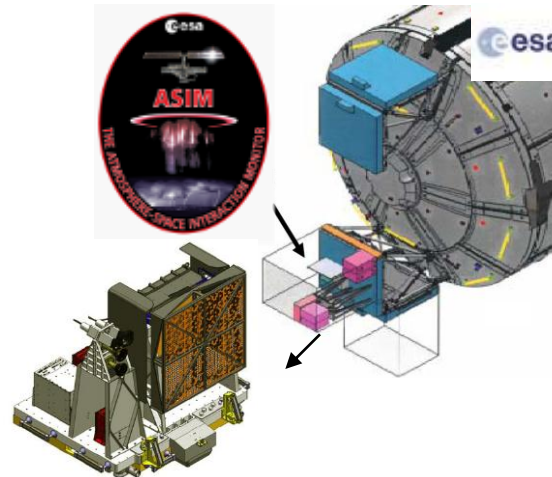
Tool for the **A**nalysis of **R**adiation
from light **N**ing and **S**prites

Launch: 2015

Polar orbit, altitude 700 km

Optical - gamma-ray – energetic e^-
& electromagnetic detectors

Nadir Obs.



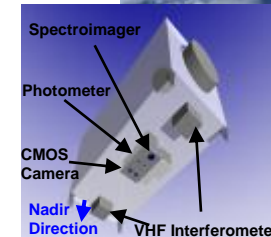
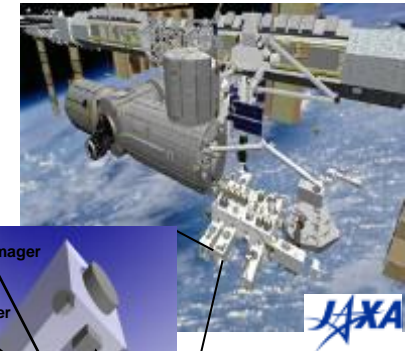
Atmosphere-**S**pace **I**nteraction **M**onitor

Launch: 2014

ISS orbit, inclination = 51°

Optical & gamma-ray detectors

Nadir Obs.



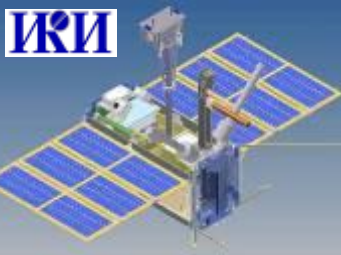
Global **L**ightning and spr**I**te **M**easurements

Launch: end of 2012

ISS orbit, inclination = 51°

Optical detectors

Nadir Obs.



Launch: 2012

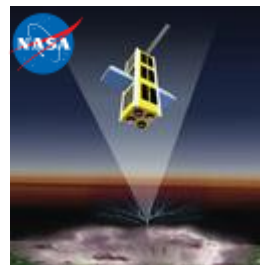
inclination = 51.6°

Altitude = 500 km

Optical - gamma-ray -
energetic e^- & EM
detectors

Nadir Obs

Chibis-M



Firefly ?

COBRAT

Coupled **O**bservation from **B**alloons **R**elated to **A**SIM and **T**ARANIS

Long duration balloon flights for the study of high energy
phenomena observed in the atmosphere above
thunderstorms and their consequences for stratospheric
chemistry