

# Alerts and APPs

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on behalf of AGILE Team

11th AGILE Science Workshop May 16-17, 2013, ASI Headquarters, Rome

### **ALERTS?**

- Transient discoveries from gamma-ray sources during the AGILE observations is possible given the large Field of View and the AGILE sensitivity
- A quick reaction time enable us to follow the evolution of a flaring event
- The search for γ-ray transients (Galactic and extra-Galactic) detectable on timescales of 1-2 days is one of the main daily activities performed by the AGILE Team

## Two independent pipelines

- The Science Alert System (SAS)
   pipeline running at INAF/IASF
   Bologna (IASFBO)
- The Quick-Look Scientific (QLS)
   pipeline running at the AGILE Data
   Center (ADC)



**Automated scientific analysis systems** 

**Cross check of scientific results** 

## QLS@ADC

## SAS@IASFBO

It performs an accurate data processing

It performs a raw data processing

It generates daily REPORTS

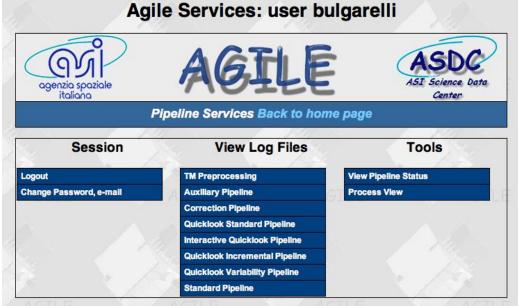
It generates ALERTS immediatly



It works with a <u>continuous</u>
<u>integration</u> of data: the data are
analyzed for each orbit

## QLS@ADC

From: no reply@asdc.asi.it



			roc. 03/11/2012 n		D:5623	4) ##				
() High Latitude: -1st selection  xsnr>2.2&sqrtts>2.9  ##										
AGILE NAME NAME	LII	BII	StartDate	EndDate	MLF1	ux Xim	SNR sqr	t(TS) Other nam	e 1AGL NAM	IE BZCAT
(d	(E-8ph/cm^2s)									
<del>-</del>	+		+	+	+	+	+	+	+	+
AGLJ0543-1544	219.876	-22.068	2012-11-01T11:50	11-03	305.4	2.52	3.9			
AGLJ0816+0922			2012-11-01T11:50					BZBJ0814+0856		BZBJ0814+08
			2012-11-01T11:50				3.3			BZBJ0814+08
			2012-11-01T11:50				3.3			
			2012-11-01T11:50					SDSSJ103317.94+		BZBJ1033+42
			2012-11-01T11:50					NGC6251		BZUJ1632+82
			2012-11-01T11:50				3.6			
			2012-11-01T11:50				3.1			BZQJ2202-83
AGLJ2240+7609			2012-11-01T11:50	11-03	243.2	3.00	2.9			
			exclude:xsnr>2.21							
AGLJ0537-1553	219.410	-23.465	2012-11-01T11:50	11-03	296.1	1.97	4.2			
B) Galactic Pl	ane (low	Latitud	e): -1st selectio	n lxsnr>	3&sart	ts>2.1	##			
			2012-11-01T11:50					GEMINGA	1AGLJ0634+17	
AGLJ0750-2649	243.181	-0.350	2012-11-01T11:50	11-03	159.9	3.56	3.3			
AGLJ0835-4515	263.623	-2.814	2012-11-01T11:50	11-03	874.5	8.35	11.1	VelaPSR	1AGLJ0835-45	
AGLJ0954-6123	283.289	-5.461	2012-11-01T11:50	11-03	241.1	3.15	2.0			
GLJ1317-6308	305.812	-0.406	2012-11-01T11:50	11-03	464.3	3.82	2.6			
GLJ1715-3820	348.621	0.181	2012-11-01T11:50	11-03	640.2	3.15	2.2			
GLJ1740-3155	356.853	-0.570	2012-11-01T11:50	11-03	590.1	3.15	2.2		1AGLJ1736-32	
GLJ1825-1417	17.235	-0.749	2012-11-01T11:50	11-03	323.5	5.02	3.5		1AGLJ1824-14	
GLJ1851-0009	32.781	-0.048	2012-11-01T11:50	11-03	445.9	5.48	2.0			
GLJ1853+0336	36.368	1.190	2012-11-01T11:50	11-03	470.3	5.57	2.5			
GLJ1956+3309	69.377	2.330	2012-11-01T11:50	11-03	312.2	4.06	2.3			
GLJ2019+3728			2012-11-01T11:50	11-03	223.5	4.82	3.2	PSRJ2021+3651	1AGLJ2021+36	
-2nd selecti	on Isqrt	ts>3.&ex	clude:xsnr>31							
AGLJ0703-0220	216 300	1 451	2012-11-01T11:50	11-03	166.6	2.52	3.3			

**AGILE** services

Daily REPORTS

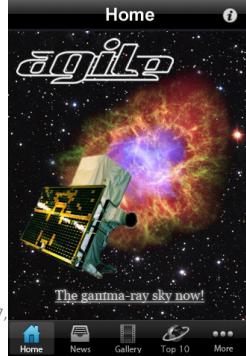
## SAS@IASFBO ALERTS & APPS

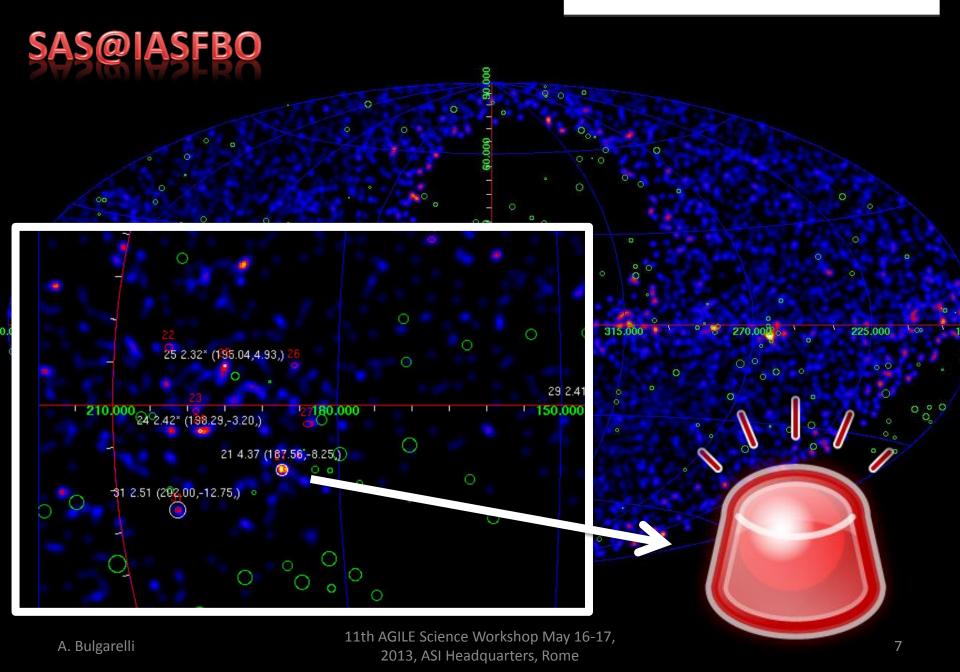
Science Alert System



To react in the shortest time

To access the results with a mobile phone





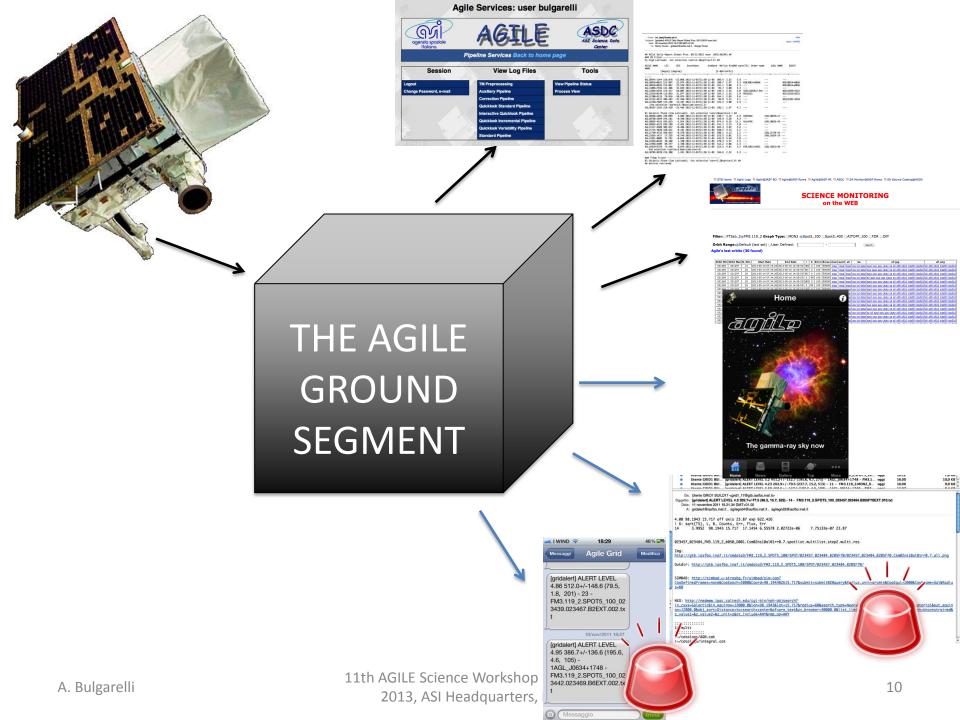


Messaggio

# within 2-3

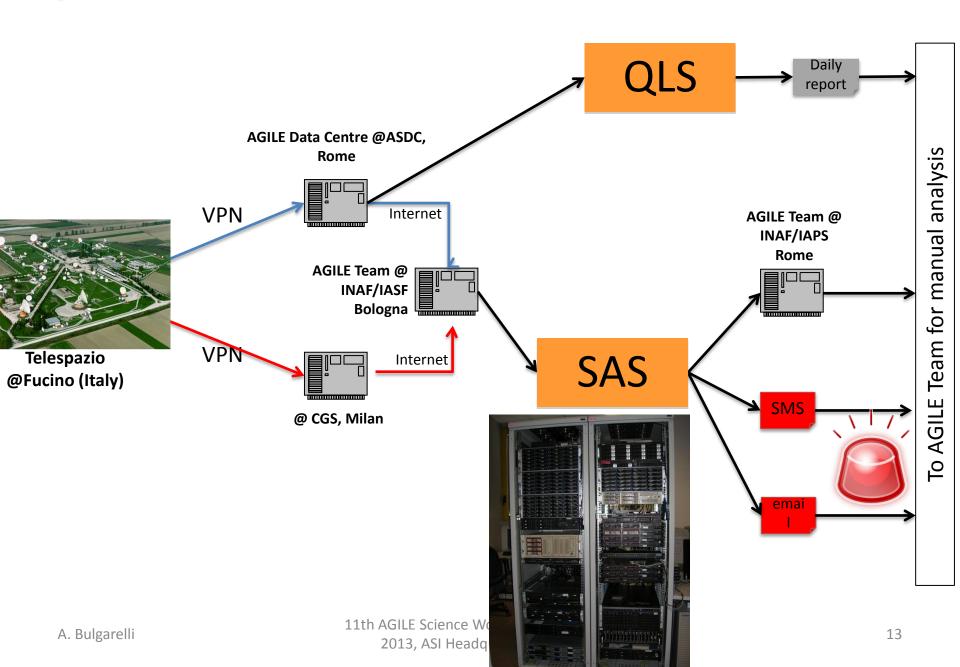


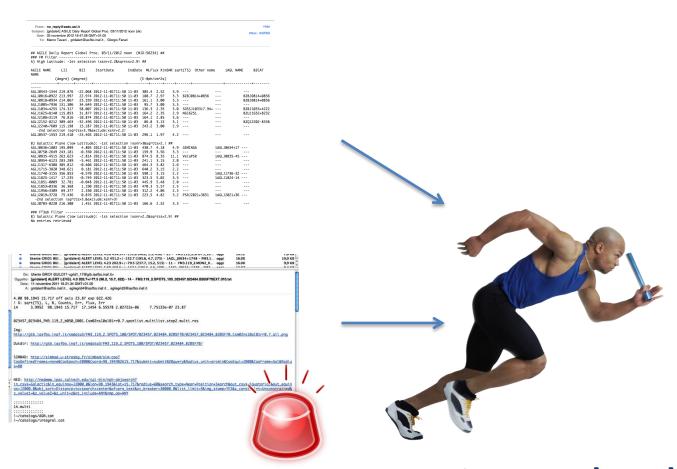
# THE FASTEST GAMMA-RAY GROUND SEGMENT OF THE WORLD













A <u>manual analysis</u> starts when an alert is generated above a well defined threshold (with  $V(Ts) \ge 4.5$ )



2013, ASI Headquarters, Rome

The key point of the overall system: the AGILE Team can start with the manual analysis within 2-3 hours (SAS@IASFBO)

During this "quick-look" analysis the data are consolidated by a more accurate data processing (QLS@ADC)

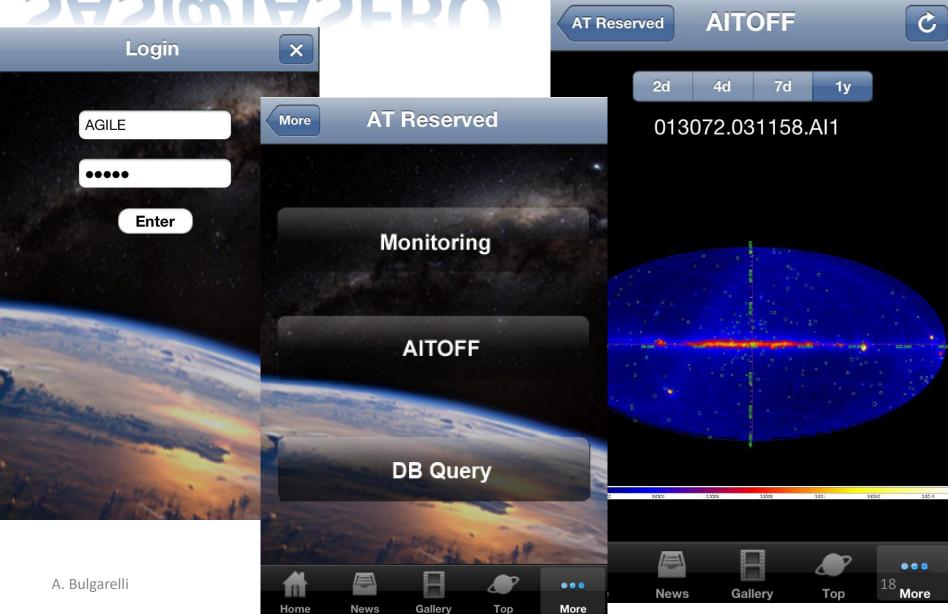


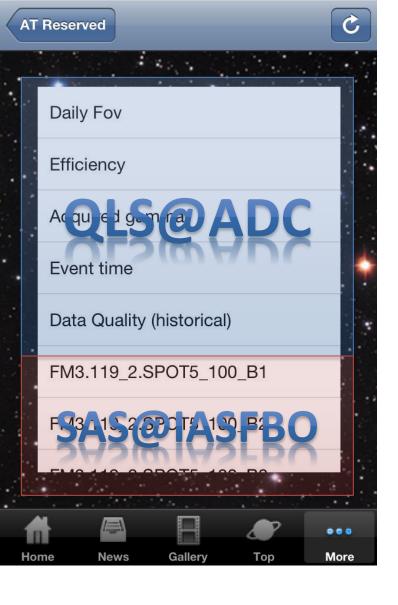
Both pipelines work with a <u>common goal</u> → to produce scientific results in the shortest possible time and with the best quality of data

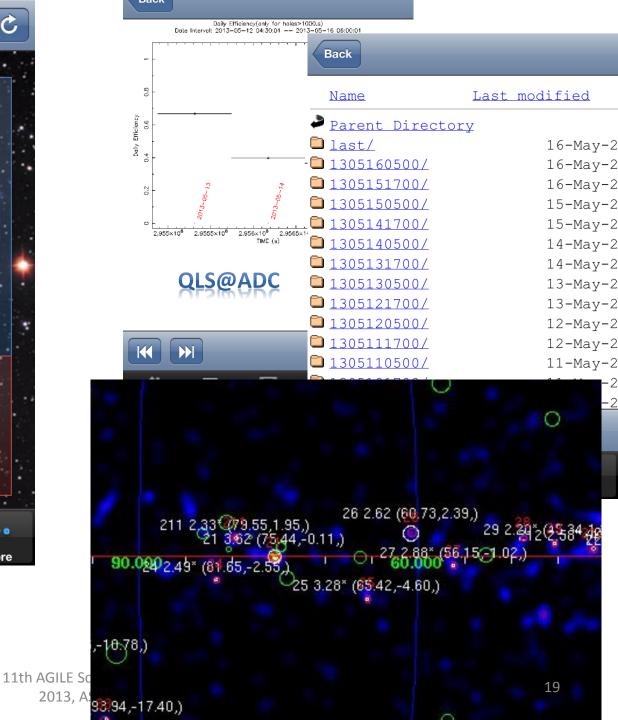


## SAS@IASFBQ









## SAS@IASFBO

#### For the first time

- An App is integrated into a scientific ground segment
- The App is used for scientific activities

#### In addition

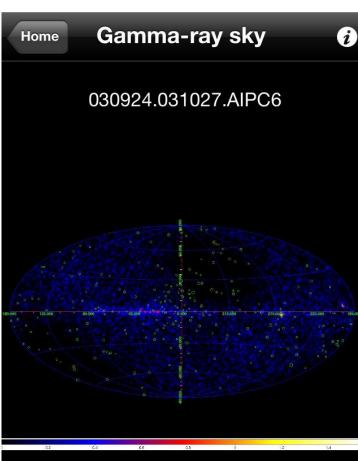
- Some step of the monitoring activities become 'mobile'





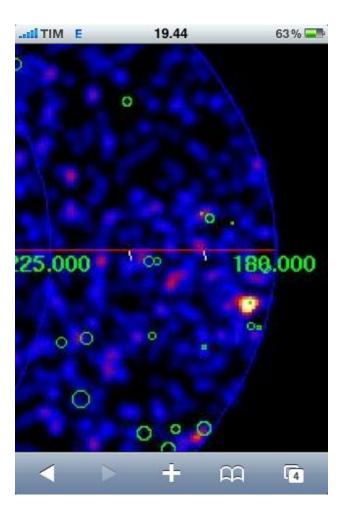
## SAS@IASFBO Put all together!







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#### "AGILE: from Black Holes to the Earth"

AGILE is a satellite of the Italian Space
Agency (ASI) dedicated to unveiling the most
exciting mysteries of our Universe. Black
holes, neutron stars, stellar explosions, very
far galaxies from deep space and other
exotic cosmic sources emit very

radiation (gamma-rays) that A detect in orbit.

Launched in 2007, the satellite Serving the Universe with m

# GTB Home 

Agile Logs 
Agile@IASF BO 
Agile@IASF Roma 
Agile@IASF MI 
ASDC 

SA Monitor@IASF Roma 
SA Source Catalog



SCIENCE MONITORING on the WEB

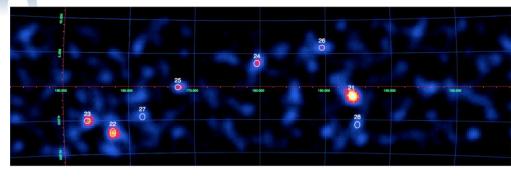
## not only APPS

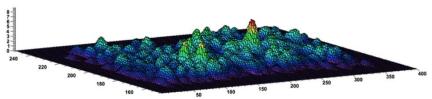
11th AGILE Science Workshop Ma 2013, ASI Headquarters, Ro

### SAS@IASFBO

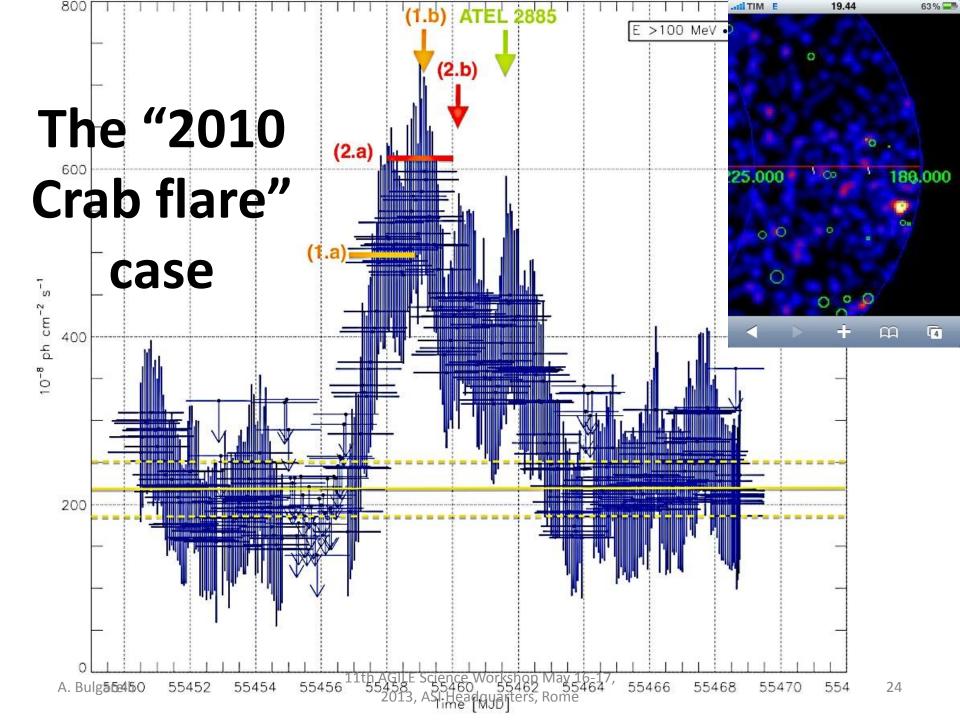
#### **Technicalities**

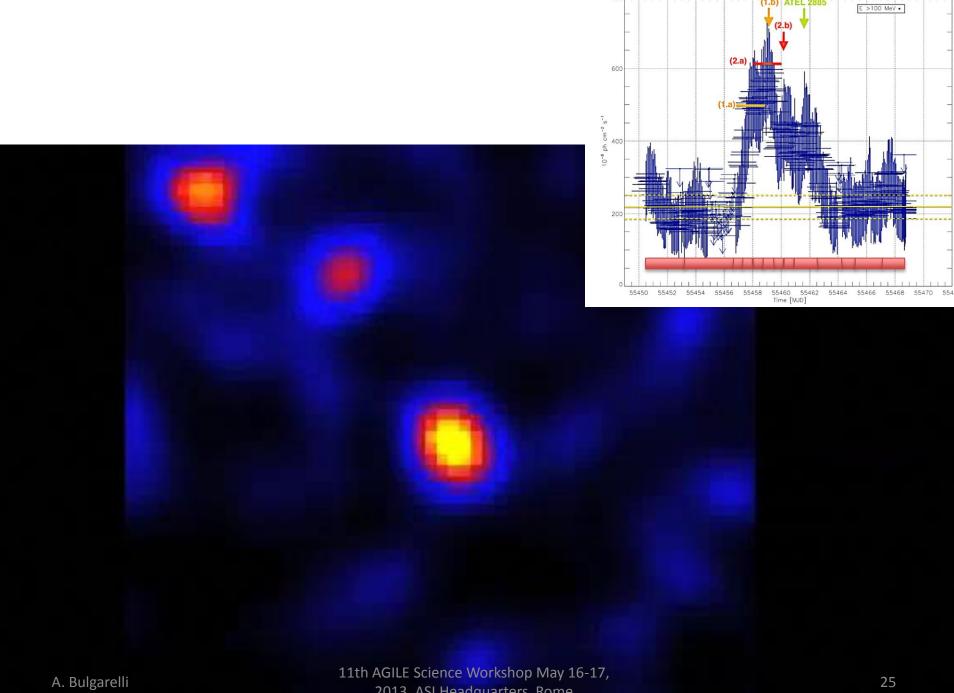
- Scripting in Ruby
- About 64 cpu cores for overall processing
- The system works with different <u>integration</u> <u>time windows:</u> (1-2-7 days)
- A binned maximum likelihood estimator (developed @IASFMI) is used as analysis tool





- The selection of candidate flares is performed with
  - transient blind search methods ("spotfinder")
  - transient search from a list of known sources





#### Outside GCN IAUCs Other MacOS: Dashboard Widget Follow ATel on Twitter

ATel Community Site

The Astronomer's Telegram

Post a New Telegram I Search I Information Telegram Index Obtain Credential To Post I RSS Feeds I Email Settings

Present Time: 27 Mar 2012; 09:36 UT



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#### AGILE detection of enhanced gamma-ray emission from the Crab Nebula region

ATel #2855; M. Tavani (INAF/IASF Roma), E. Striani (Univ. Tor Vergata), A. Bulgarelli (INAF/IASF Bologna), F. Gianotti, M. Trifoglio (INAF/IASF Bologna), C. Pittori, F. Verrecchia (ASDC), A. Argan, A. Trois, G. De Paris, V. Vittorini, F. D'Ammando, S. Sabatini, G. Piano, E. Costa, I. Donnarumma, M. Feroci, L. Pacciani, E. Del Monte, F. Lazzarotto, P. Soffitta, Y. Evangelista, I. Lapshov (INAF-IASF-Rm), A. Chen, A. Giuliani(INAF-IASF-Milano), M. Marisaldi, G. Di Cocco, C. Labanti, F. Fuschino, M. Galli (INAF/IASF Bologna), P. Caraveo, S. Mereghetti, F. Perotti (INAF/IASF Milano), G. Pucella, M. Rapisarda (ENEA-Roma), S. Vercellone (IASF-Pa), A. Pellizzoni, M. Pilia (INAF/OA-Cagliari), G. Barbiellini, F. Longo (INFN Trieste), P. Picozza, A. Morselli (INFN and Univ. Tor Vergata), M. Prest (Università dell'Insubria), P. Lipari, D. Zanello (INFN Roma-1), P. W. Cattaneo, A. Rappoldi (INFN Pavia), P. Giommi, P. Santolamazza, F. Lucarelli, S. Colafrancesco (ASDC), L. Salotti (ASI) on 22 Sep 2010: 14:45 UT

> Distributed as an Instant Email Notice Transients Credential Certification: Marco Tavani (tavani@iasf-roma.inaf.it)

Subjects: Pulsar

Referred to by ATel #: 2856, 2858, 2861, 2866, 2867, 2868, 2872, 2879, 2882, 2889, 2893, 2903, 2921, 2967, 2968, 2994, 3058, 3276, 3283

AGILE is detecting an increased gamma-ray flux from a source positionally consistent with the Crab Nebula.

Integrating during the period 2010-09-19 00:10 UT to 2010-09-21 00:10 UT the AGILE-GRID detected enhanced gamma-ray emission above 100 MeV from a source at Galactic coordinates (1,b) = (184.6, -6.0) + -0.4 (stat.) + -0.1 (syst.) deg, and flux F > 500 e-8 ph/cm2/sec above 100 MeV, corresponding to an excess with significance above 4.4 sigma with respect to the average flux from the Crab nebula (F = (220 +/- 15)e-8 ph/cm^2/sec, Pittori et al., 2009, A&A, 506, 1563).

We strongly encourage multifrequency observations of the Crab Nebula region.

#### Fermi LAT confirmation of enhanced gamma-ray emission from the Crab Nebula region

ATel #2861; R. Buehler (SLAC/KIPAC), F. D'Ammando (INAF-IASF Palermo), E. Hays (NASA/GSFC) on behalf of the Fermi Large Area Telescope Collaboration on 23 Sep 2010; 17:34 UT

Distributed as an Instant Email Notice Transients Credential Certification: Rolf Buehler (buehler@slac.stanford.edu)

Subjects: >GeV, Pulsar

Referred to by ATel #: 2866, 2867, 2868, 2872, 2879, 2882, 2889, 2893, 2903, 2921, 2967, 2968, 2994, 3058, 3276, 3283

Following the detection by AGILE of increasing gamma-ray activity from a source positionally consistent with the Crab Nebula occurred from September 19 to 21 (ATel #2855), we report on the analysis of the >100 MeV emission from this region with the Large Area Telescope (LAT), one of the two instruments on the Fermi Gamma-ray Space Telescope.

Preliminary LAT analysis indicates that the gamma-ray emission (E>100 MeV) observed during this time period at the location of the Crab Nebula is (606 +/- 43) x10^-8 ph/cm2/sec, corresponding to an excess with significance >9 sigma with respect to the average flux from the Crab nebula of (286 +/- 2) x10^-8 ph/cm2/sec, estimated over all the Fermi operation period (only statistical errors are given). Ongoing Fermi observations indicate that the flare is continuing.

The flaring component has a spectral index of 2.49 +/- 0.14. Its position, Ra: 83.59 Dec: 22.05 with a 68% error radius of 0.06 deg, is coincident with the Crab Nebula.

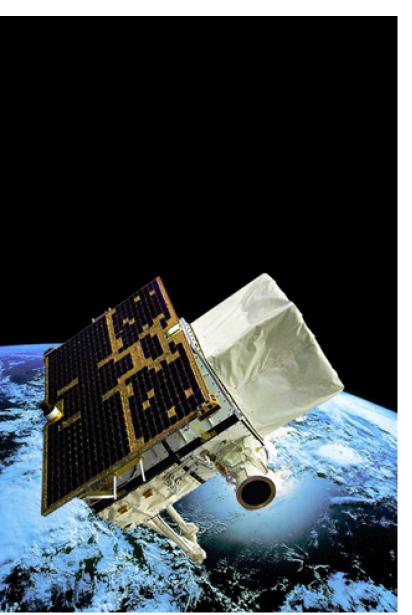
Fermi will interrupt its all-sky scanning mode between 2010-09-23 15:49:00 UT and 2010-09-30 15:49:00 UT to observe the Crab Nebula. Afterwards regular gamma-ray monitoring of this source will continue. We strongly encourage further multifrequency observations of that

For this source the Fermi LAT contact person is Rolf Buehler (buehler@stanford.edu).

The Fermi LAT is a pair conversion telescope designed to cover the energy band from 20 MeV to greater than 300 GeV. It is the product of an international collaboration between NASA and DOE in the U.S. and many scientific institutions across France, Italy, Japan and Sweden.

The first alert received from Crab with a flux level that exceed 1-sigma mean flux level was received at 2010-09-20T02:04:04 UTC.

The Crab reach its maximum flux for E>100 MeV in AGILE data between 2010-09-19T01:54:43 and 2010-09-20T23:47:51 UTC in the continuous integration procedure with 2 days integration time. The later Was generaled งาล email อกอ 2013, ASI Headquarters, Rome SMS at 2010-09-21T02:00:54 UTC.



#### **Conclusions**

- Key factors:
  - IASF Bologna pipeline that react within 2 hours
  - Two independent pipelines
  - Cross check with consolidated data.
- The effectiveness of the monitoring system is demonstrated on a daily basis and by the great number of ATel (90 in 6 years) published by AGILE Team; very important discoveries were started from this system.
- This allows an effective follow-up of flaring sources, enabling AGILE to maximize the scientific return of the mission.

# THANKS FOR YOUR ATTENTION!