# Gamma-ray and multi-frequency studies of Blazars

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### The electromagnetic spectrum





The electromagnetic spectrum





### **AGN : Two main categories**

1. Dominated by (mostly) thermal emission from accretion disk -

Radio quiet AGN (>~90 %)

(normal QSO powered by accretion onto a SM black hole)

1. Dominated by Non-Thermal radiation –

Jet dominated AGN (< 10%)

When  $\Theta < \Theta_{blazar}$   $\implies$  Blazar

As of today, about 3,150 blazars are known (Bzcat, edition 4.1, Massaro et al. 2012).

This number is increasing rapidly but it remains a small percentage of the over one million AGN known Flesch, E. The Million Quasars (MILLIQUAS) Catalog, Version 3.1 (22 October 2012) <u>http://quasars.org/milliquas.htm</u>









### http://tools.asdc.asi.it/SED/

## SED<sup>(t)</sup> builder V3.0

A tool to build and handle Spectral Energy Distributions, time-resolved SEDs and multi-frequency light-curves



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# A UV and X-ray spectral database of the blazars most frequently observed by Swift.

P. Giommi<sup>1,2</sup>, M. Perri<sup>1,3</sup>, M. Capalbi<sup>4</sup>, V. D'Elia<sup>1,3</sup>, F. Verrecchia, G. Tagliaferri<sup>4</sup>, G: Cusumano<sup>5</sup>, et al.<sup>1,3</sup>

2104 in preparation

## Systematic XRT (and UV+BAT) analysis of 4056 observations of 113 blazars frequently observed by Swift between 2004 and early 2014 8152 XRT orbit-based spectra Some examples

MKN421: 560 observations, 1178 orbits, 1738 XRT spectra3C454.3: 350 observations, 740 orbits, 1090 XRT spectraPKS2155-304: 126 observations, 284 orbits, 410 spectra

Results will be available via a dedicated on-line catalog (<u>www.asdc.asi.it/xrtspectra</u>) and spectral through the ASDC SED tool (tools.asdc.asi.it/SED)

### The ASDC X-ray spectral database of blazars frequently observed by Swift

Last update: 28-04-2014

Entry number	Name	Phot index Norm_pl		Alpha LP	Beta LP	Norm_lp ‡	Flux210 pl	Observation time	Orbit Orbit ‡	
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1 🗹 Select	ASDC Data Explorer BZQJ2151-3027	1.09+/-0.22	0.0023	1.25+/-0.34	-0.37+/-0.66	0.0022	2.5e-11+/-6.9e-12	2014.24134	3	
2 🗹 Select	ASDC Data Explorer BZQJ2151-3027	1.21+/-0.12	0.0021	1.16+/-0.26	0.07+/-0.38	0.0021	1.9e-11+/-2.2e-12	2014.23875	2	
3 🗹 Select	ASDC Data Explorer BZQJ2151-3027	1.03+/-0.13	0.0021	1.2+/-0.28	-0.28+/-0.42	0.0021	2.6e-11+/-5.7e-12	2014.23841	1	
4 🗹 Select	ASDC Data Explorer BZQJ2151-3027	1.11+/-0.08	0.0021	1.11+/-0.16	-0.01+/-0.24	0.0021	2.2e-11+/-2.4e-12	2014.23841	тот	
5 🗹 Select	ASDC Data Explorer PKS2004-447	1.52+/-0.25	0.0001	1.41+/-0.45	0.36+/-1.04	0.00011	5.5e-13+/-1.4e-13	2014.20553	тот	
6 🗹 Select	ASDC Data Explorer PKS2004-447	1.42+/-0.43	0.000091	1.59+/-0.57	-0.67+/-0.58	0.000086	5.8e-13+/-2.5e-13	2014.20245	тот	
7 🗹 Select	ASDC Data Explorer PG1553+113	2.46+/-0.11	0.007	2.51+/-0.11	-0.34+/-0.29	0.0067	9.3e-12+/-8.2e-13	2014.1766	тот	
8 🗹 Select	ASDC Data Explorer PG1553+113	2.46+/-0.11	0.007	2.51+/-0.11	-0.34+/-0.29	0.0067	9.3e-12+/-1.2e-12	2014.1766	1	
9 🗹 Select	ASDC Data Explorer PKS1222+216	2.31+/-0.26	0.0015	2.38+/-0.29	-0.42+/-0.9	0.0014	2.5e-12+/-9.8e-13	2014.17531	3	
10 Select	ASDC Data Explorer PKS1222+216	2.08+/-0.35	0.0015	2.17+/-0.32	-0.78+/-1.14	0.0014	3.5e-12+/-2e-12	2014.17496	2	
11 🗹 Select	ASDC Data Explorer PKS1222+216	2.22+/-0.37	0.0015	2.38+/-0.41	-0.74+/-1.27	0.0014	2.8e-12+/-1.8e-12	2014.17456	1	
12 🗹 Select	ASDC Data Explorer PKS1222+216	2.17+/-0.19	0.0015	2.37+/-0.17	-1.08+/-0.46	0.0013	2.9e-12+/-9.3e-13	2014.17456	тот	
13 Select	ASDC Data Explorer PKS1424+240	2.28+/-0.18	0.0021	2.25+/-0.2	0.27+/-0.74	0.0021	3.5e-12+/-6.5e-13	2014.17388	тот	
14 🗹 Select	ASDC Data Explorer PKS1424+240	2.28+/-0.18	0.0021	2.25+/-0.2	0.27+/-0.74	0.0021	3.5e-12+/-1.2e-12	2014.17388	1	















3C454.3 cross-correlation of light-curves in different energy bands





3C454.3



### 1WHSP: an IR-based sample of ~1,000 VHE $\gamma$ -ray blazar candidates

B. Arsioli<sup>1,2</sup>, B. Fraga<sup>1,2</sup>, P. Giommi<sup>3</sup>, P. Padovani<sup>4,5</sup>, and M. Marrese<sup>3</sup>

Submitted to A&A



# Looking for HSPs in WISE: how to find a needle in a haystack!



# Looking for HSPs in WISE: how to find a needle in a haystack!



Monthly Notices of the ROYAL ASTRONOMICAL SOCIETY

Mon. Not. R. Astron. Soc. 420, 2899-2911 (2012)

### doi:10.1111/j.1365-2966.2011.20044.

# A simplified view of blazars: clearing the fog around long-standing selection effects

P. Giommi,<sup>1\*</sup> P. Padovani,<sup>2</sup> G. Polenta,<sup>1,3</sup> S. Turriziani,<sup>1</sup> V. D'Elia<sup>1,3</sup> and S. Piranomonte<sup>3</sup>

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LETTERS

Mon. Not. R. Astron. Soc. 422, L48–L52 (2012)



doi:10.1111/j.1745-3933.2012.01234.x

### The discovery of high-power high synchrotron peak blazars

#### P. Padovani,<sup>1\*</sup> P. Giommi<sup>2</sup> and A. Rau<sup>3</sup>

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MNRAS **431**, 1914–1922 (2013) Advance Access publication 2013 March 14

#### doi:10.1093/mnras/stt305

### A simplified view of blazars: the $\gamma$ -ray case

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