

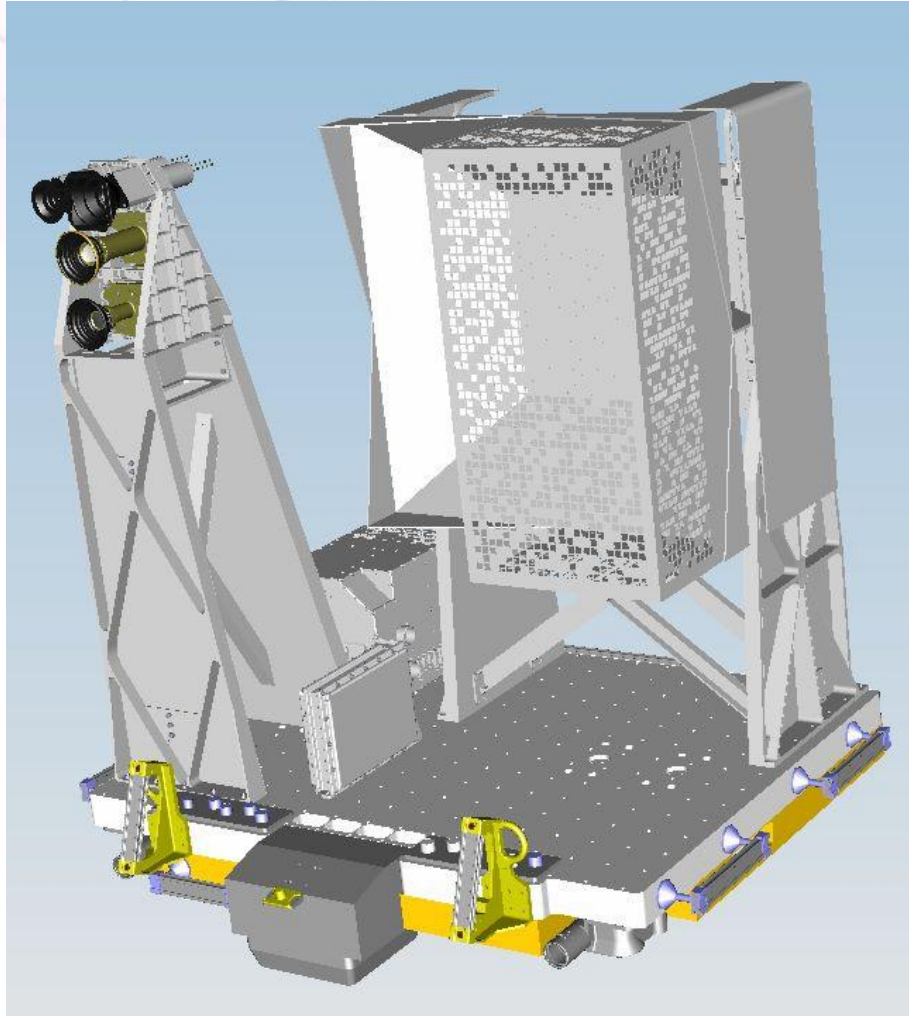


ASIM – challenges and possibilities

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On behalf of the ASIM team
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ASIM - payload



MXGS

Modular X- and Gamma-ray Sensor:

X- and gamma rays,
2 layers

Coded mask – full FOV

MMIA:

Modular Multi-spectral Imaging Array

Nadir - looking:

2 photometers

2 cameras

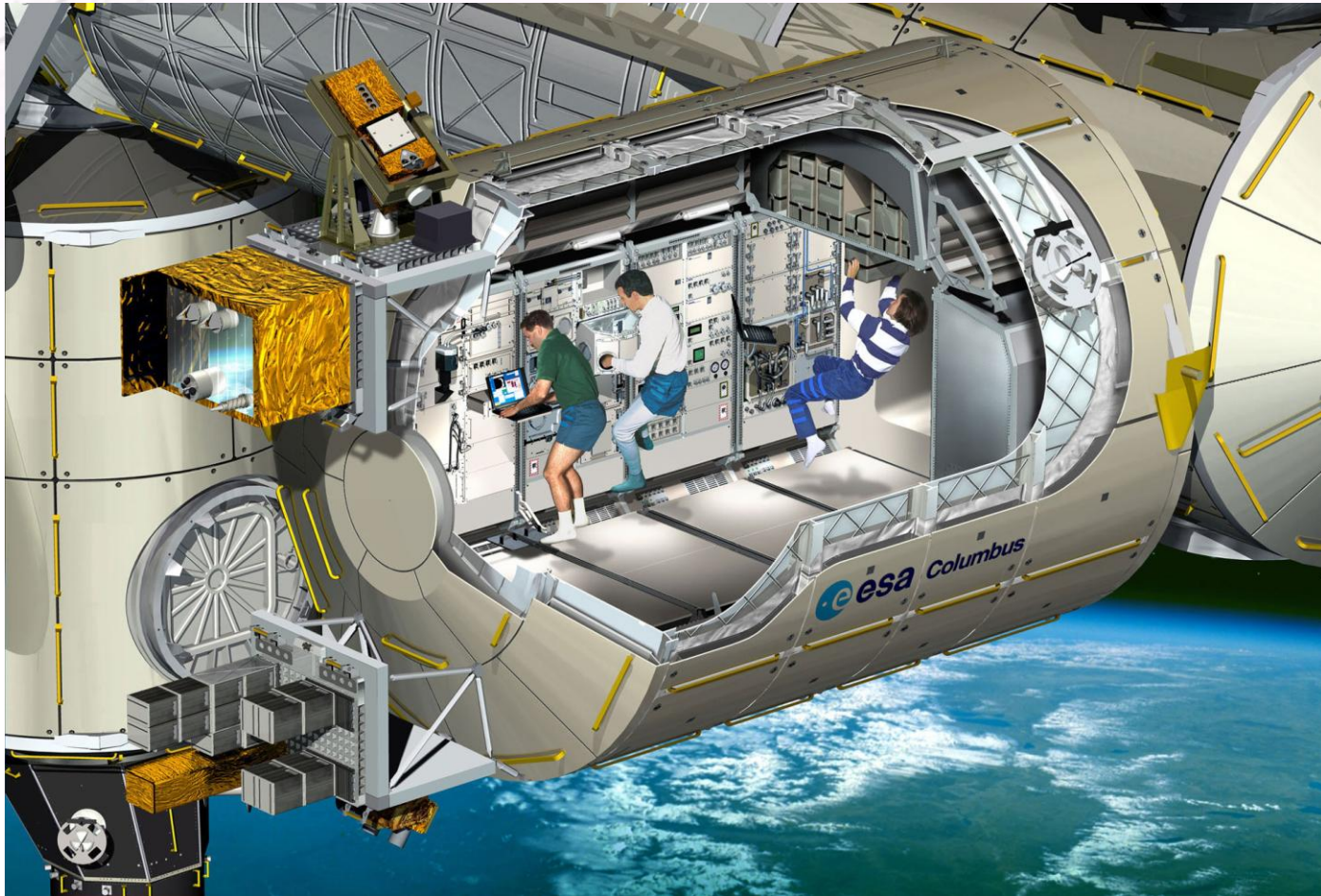
If funding available:

One more MMIA

Limb-looking

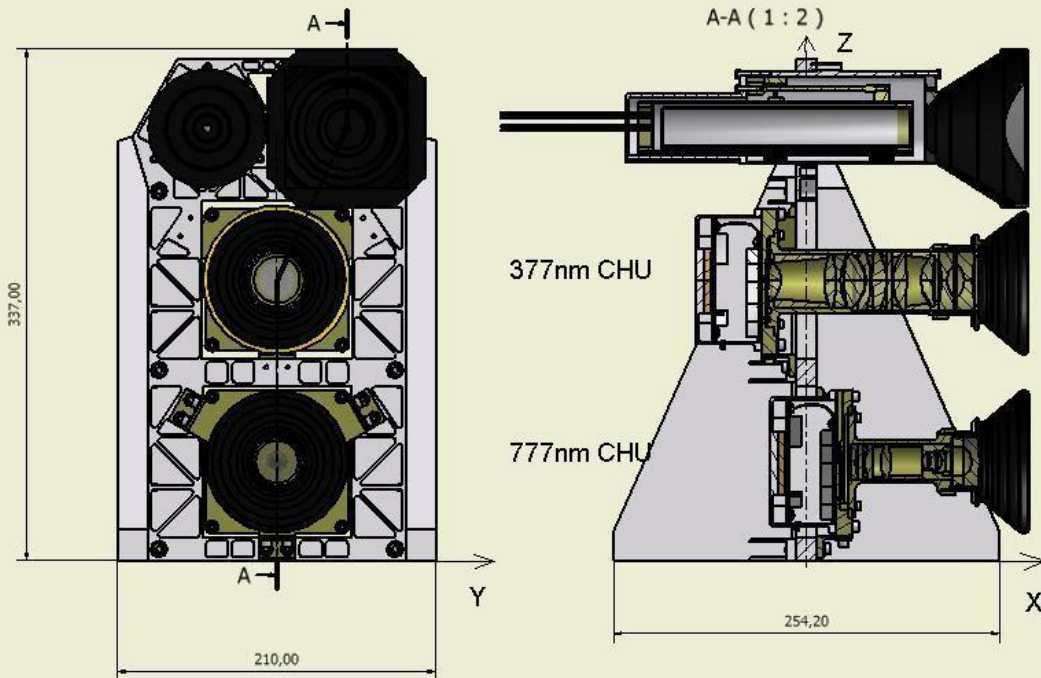


Columbus on ISS

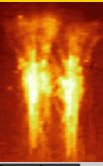




MMIA



- Photometers – 10 microsec (10 kHz)
- Cameras – 40 ms
- 337.0 nm (N₂ – 2PG)
Sprites (absorbed by O₂)
- 777.4 nm (O₁)
Lightning



“Textbook Sprite”

2004-07-18/21:30:15.267

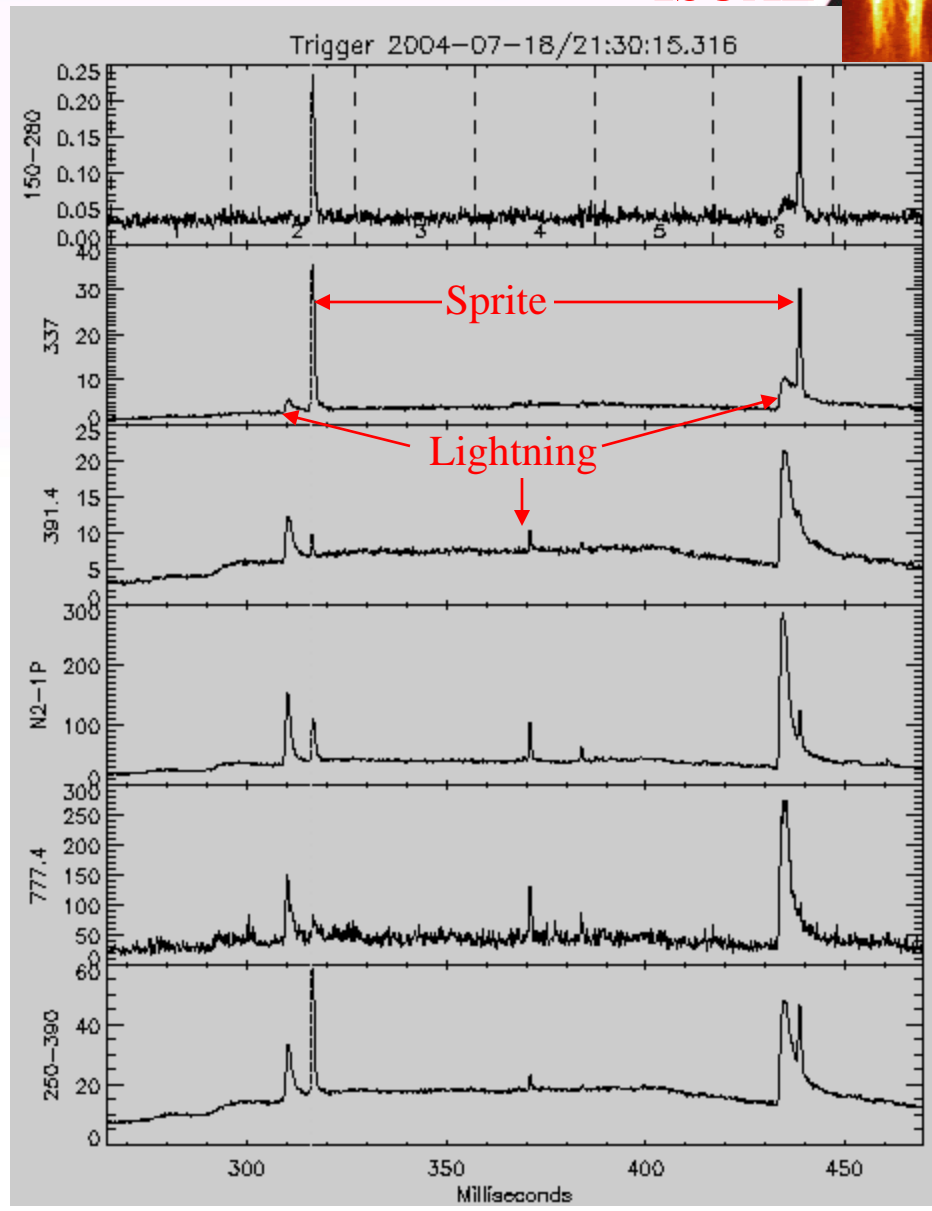
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2004-07-18/21:30:15.327

2004-07-18/21:30:15.357

2004-07-18/21:30:15.387

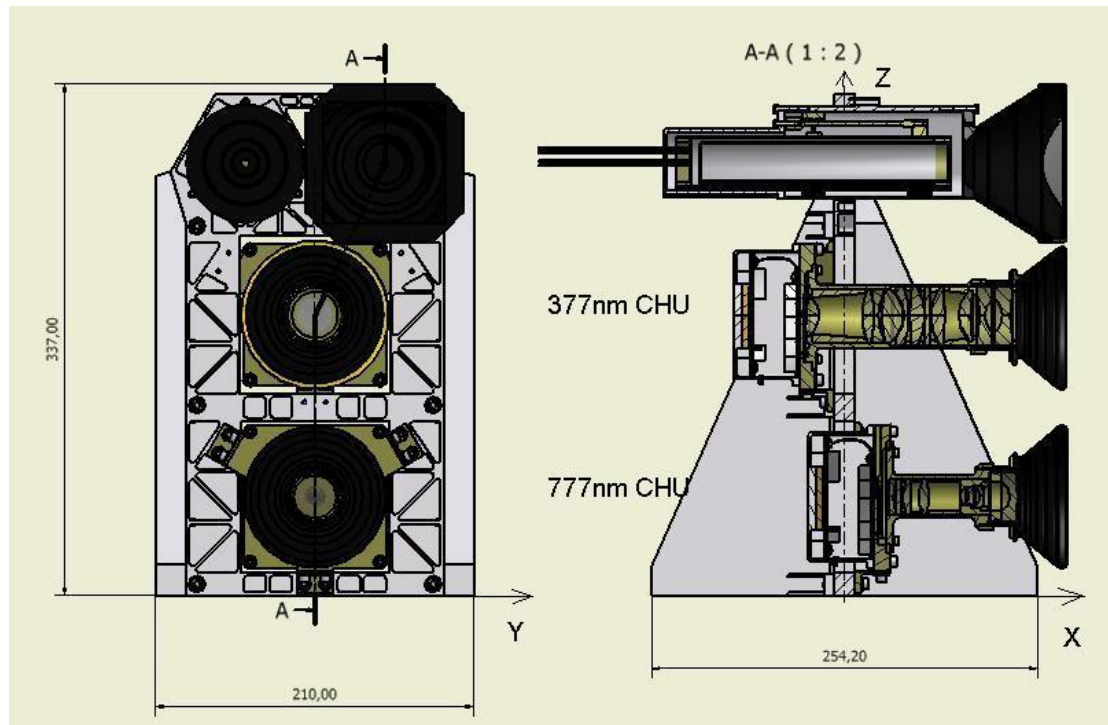
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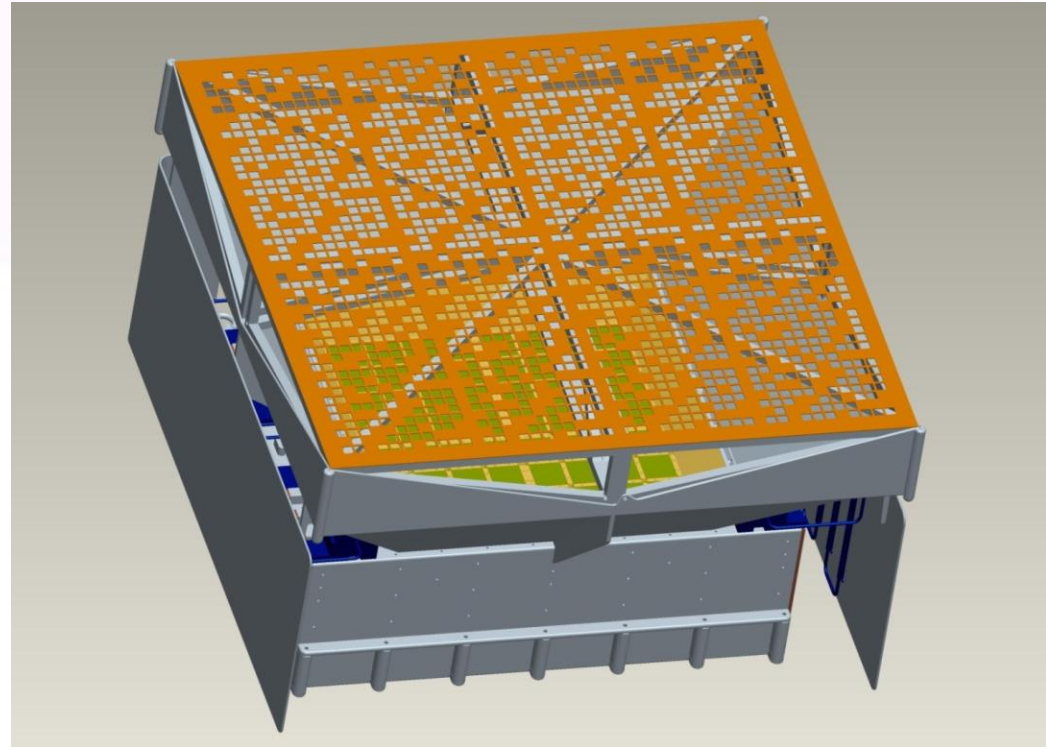
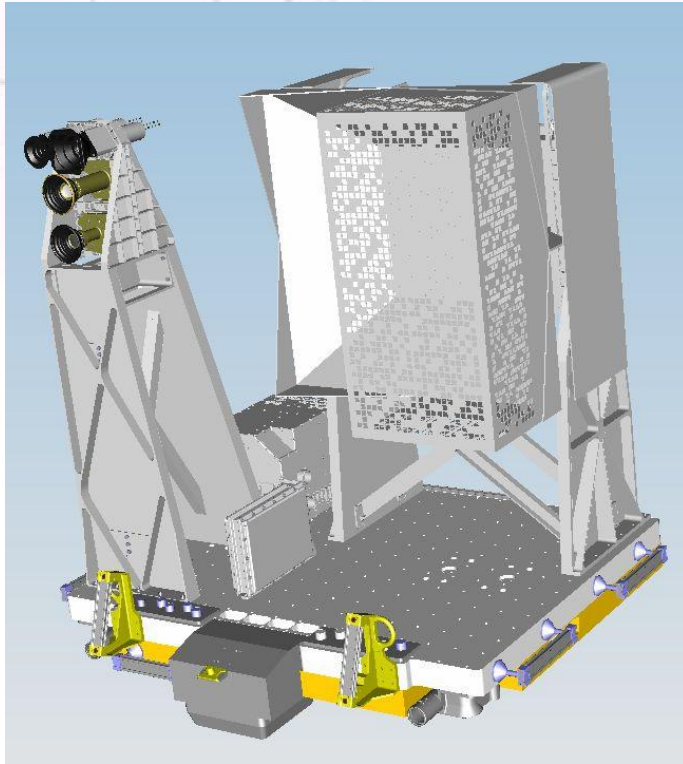
MMIA

- Camera – position of lightning/sprite
- Photometer – time at 10 microsec





ASIM - MXGS

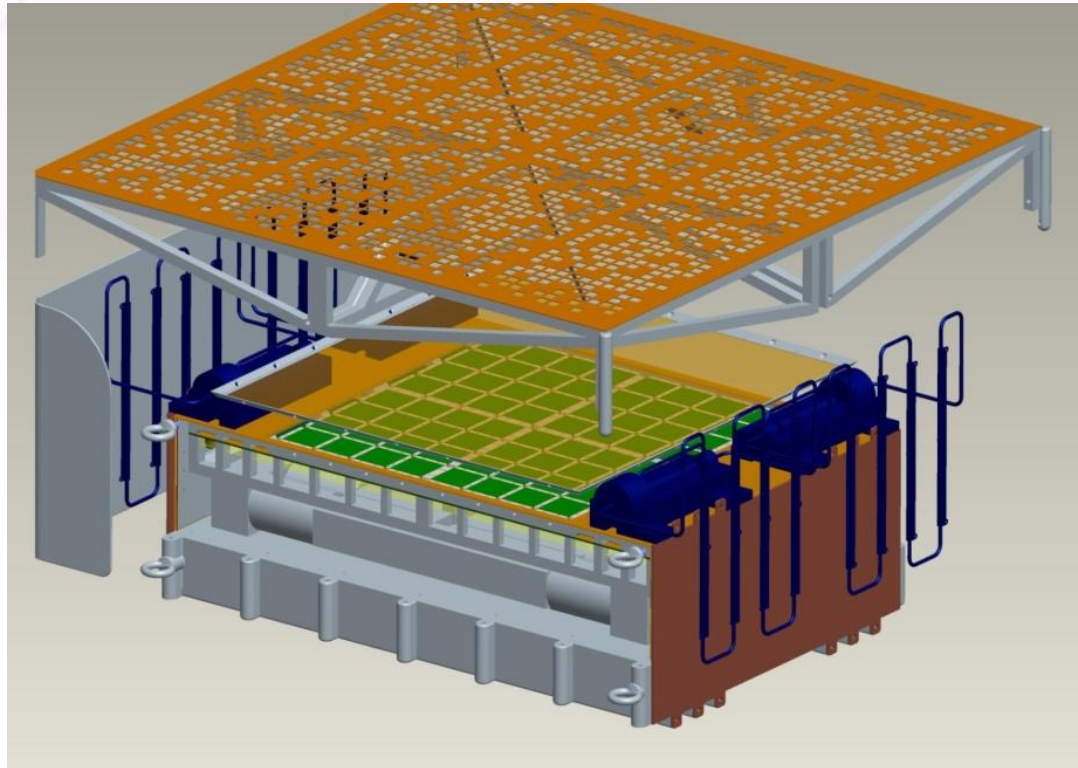


Coded mask – full FOV

- 1- Imaging (15-150 keV) – 1024 cm² - CZT
2. Spectral analysis (15 keV-20 MeV) CZT+BGO



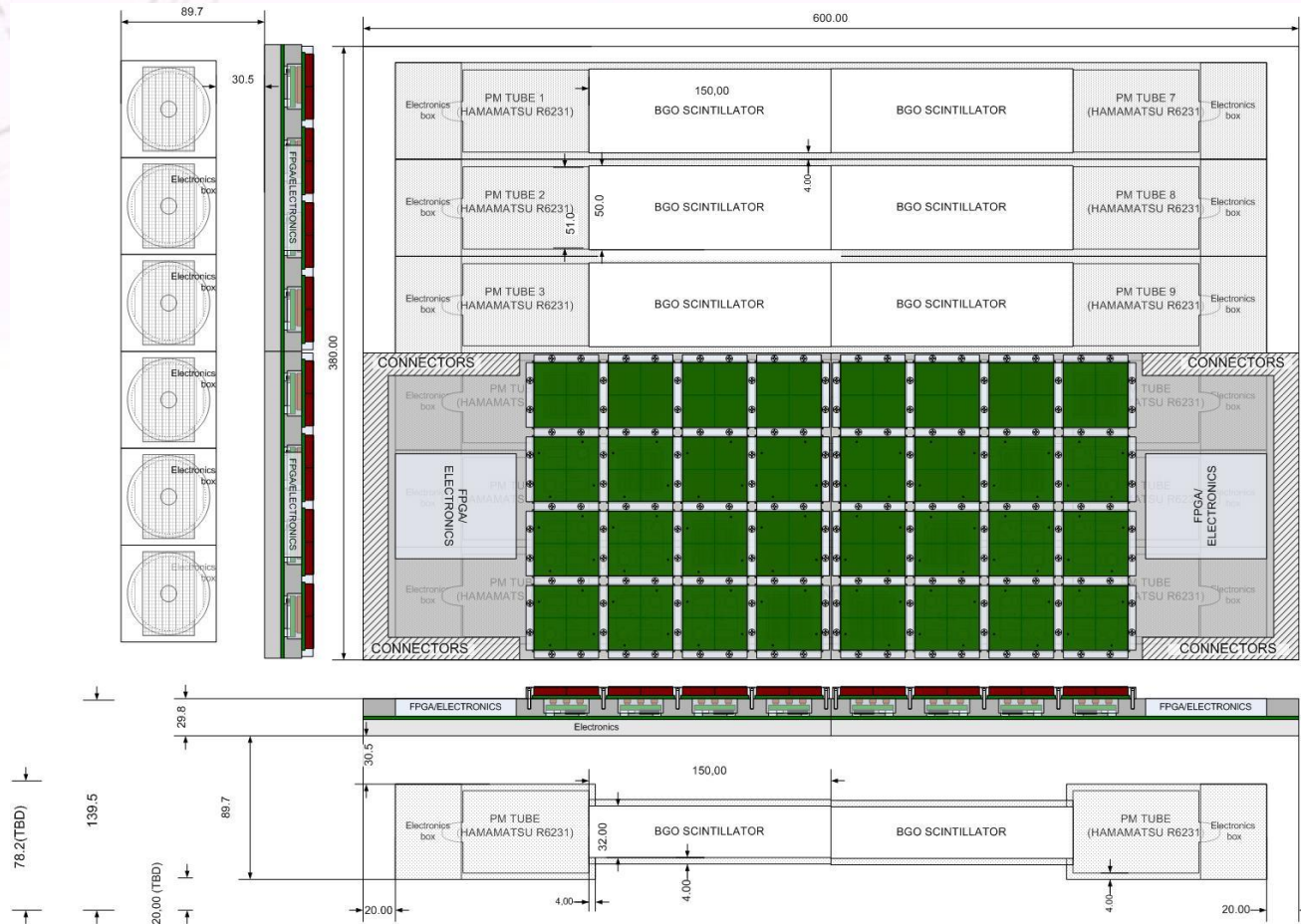
MXGS – exploded view



1024 cm² CZT – 5 mm thick – 16384 pixels
900 cm² BGO – 32 mm thick



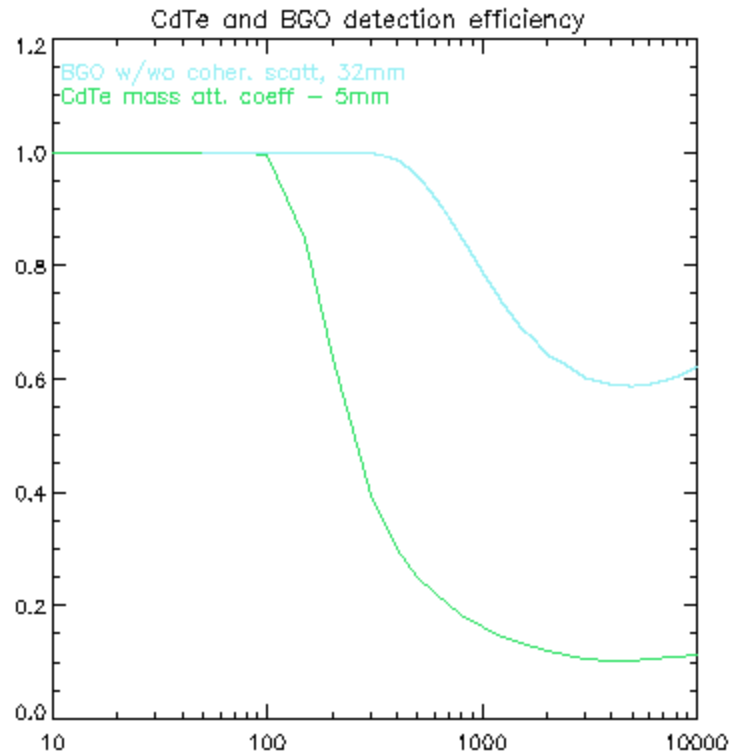
MXGS – two layers



- 1- Imaging (15-150 keV) – Pixelated CZT layer
2. Spectral analysis (15 keV-20 MeV) – CZT + BGO



RHESSI TGF if they were detected by ASIM

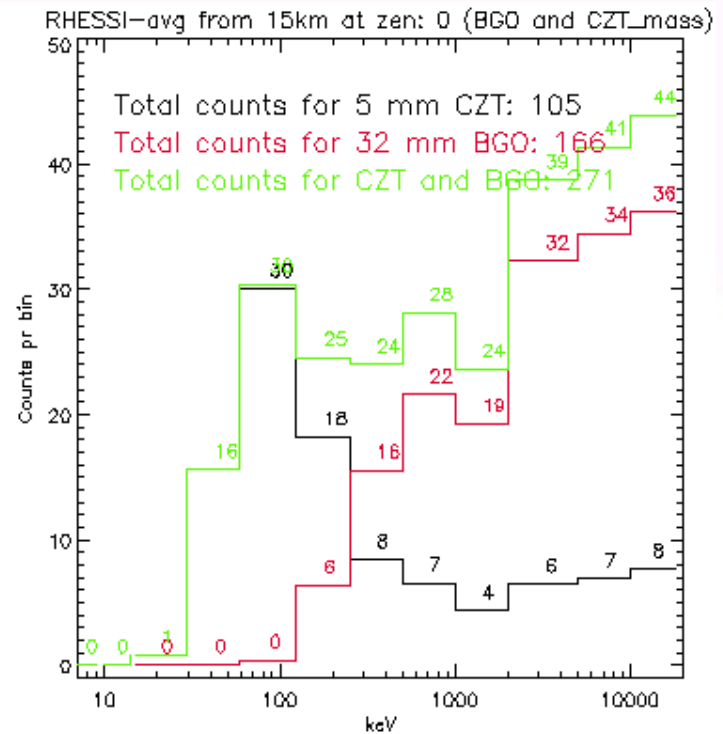
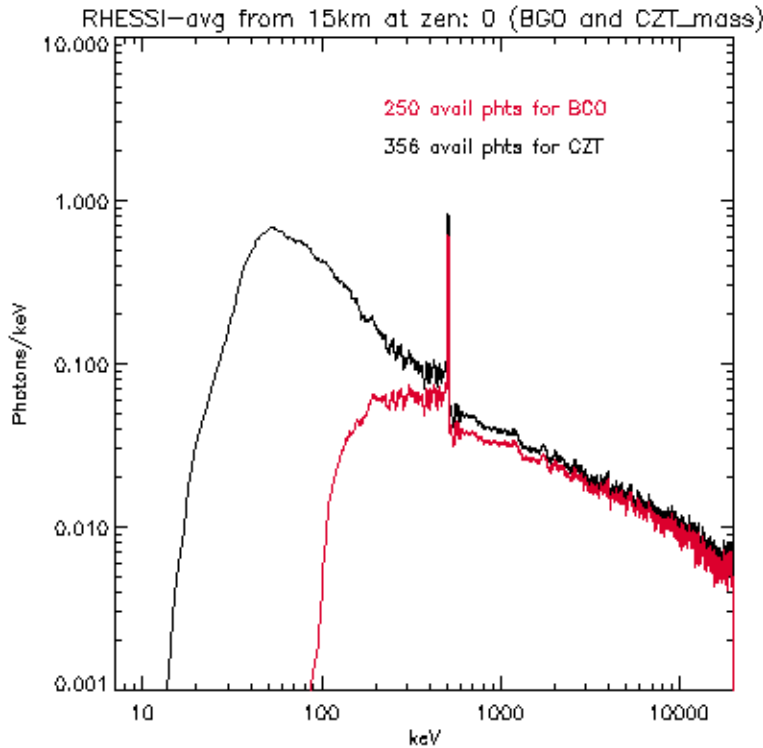


1/r²: ISS at 350 km
 RHESSI at 600 km

RHESSI: 250 cm²
ASIM: 1024 cm²



RHESSI TGF if they were detected by ASIM



RHESSI: average 27 cnts
ASIM: average 270 cnts



ASIM Q1: How common are TGFs?

BATSE: 10 TGF pr year

RHESSI: 100 TGF pr year - average 27 cnts

ASIM: 1000 TGF pr year – average 270 cnts

If TGF follow a power law distribution:

$$F(n) \sim n^{-\lambda}$$

$\lambda=2.0$: 4/24 hour: 1000 TGFs

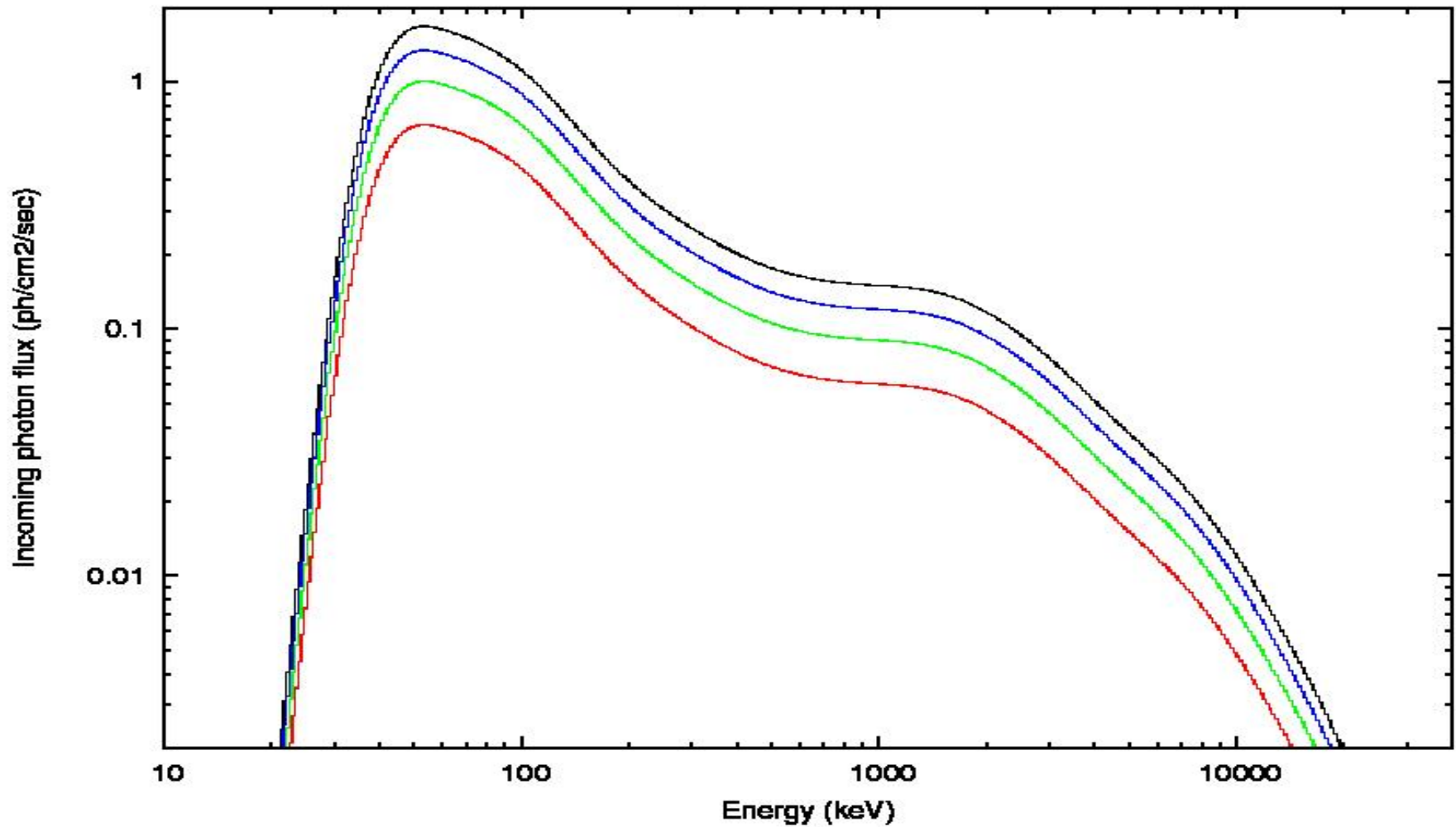
$\lambda=2.5$: 9/24 hour: 2000 TGFs

This is one of the main questions answered by ASIM.



IMAGING

RHESSI calibrated TGF spectra for 500, 400, 300, 200 photons over energy range 20-30000 keV



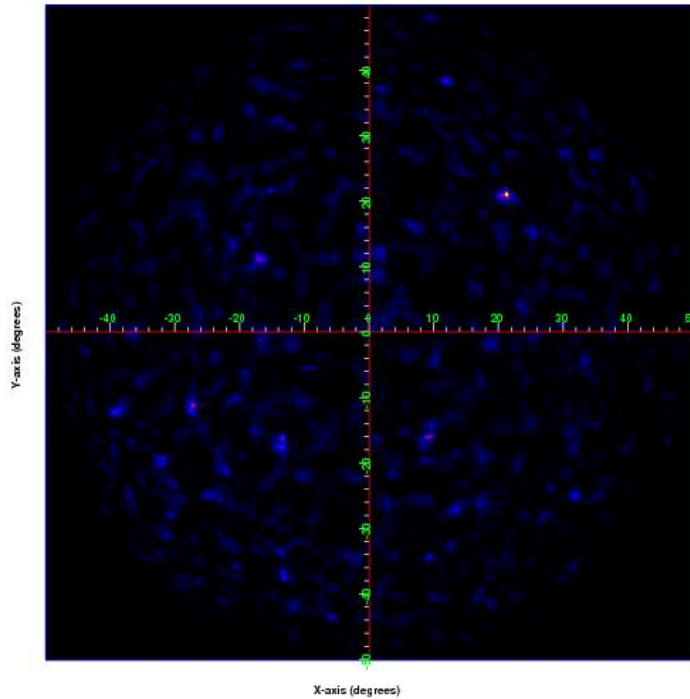


Images for 400 incoming TGF photons on 1x1 and 2x2 cm² mask elements – 30 deg off-axis

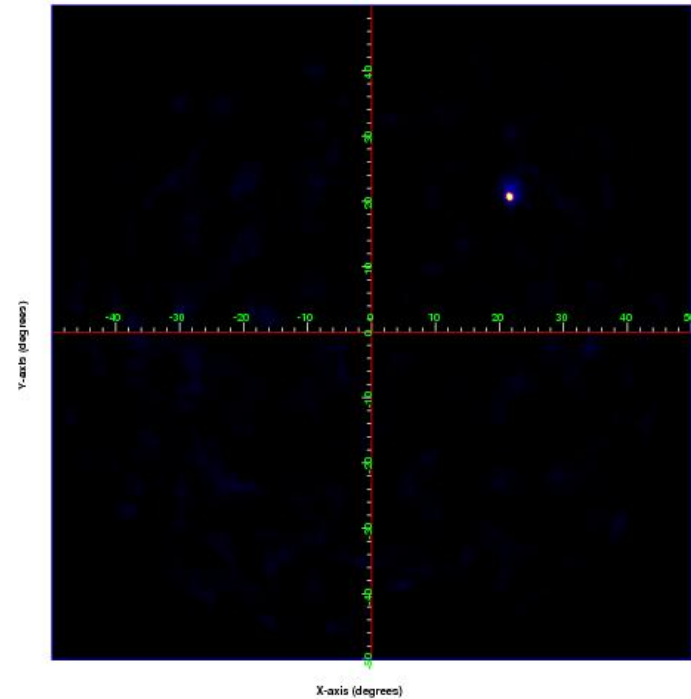
1x1 cm

2x2 cm

Backprojection image of TGF 30 deg off-axis with 400 incoming photons to 1x1 cm² mask elements



Backprojection image of TGF 30 deg off-axis with 400 incoming photons to 2x2 cm² mask elements



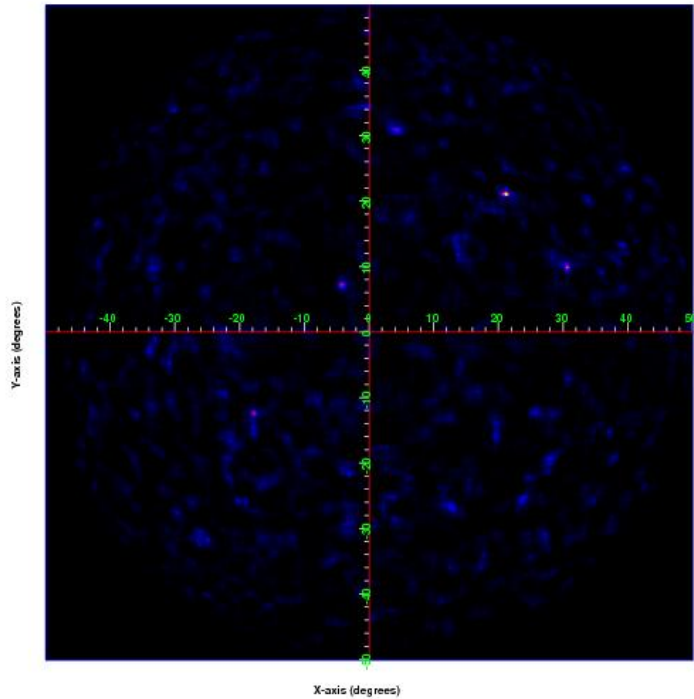


Images for 300 incoming TGF photons on 1x1 and 2x2 cm² mask elements – 30 deg off-axis

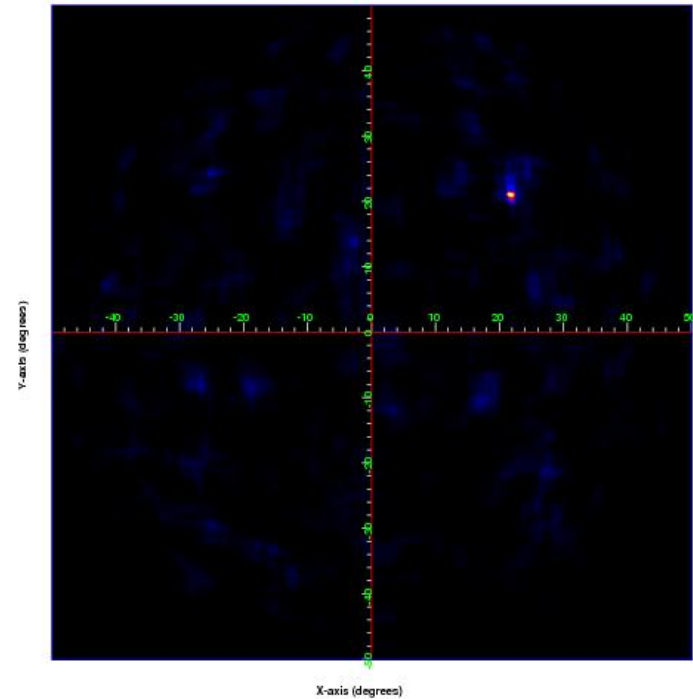
1x1 cm

2x2 cm

Backprojection image of TGF 30 deg off-axis with 300 incoming photons to 1x1 cm² mask elements



Backprojection image of TGF 30 deg off-axis with 300 incoming photons to 2x2 cm² mask elements





ASIM Q2: Are there TGFs without lightning

- ASIM first to image a TGF
- Streamers can form without subsequent return stroke.

ASIM Q3: What type of lightning?

Combined with other measurements (VLF-ELF) timing can give us information about type of lightning.

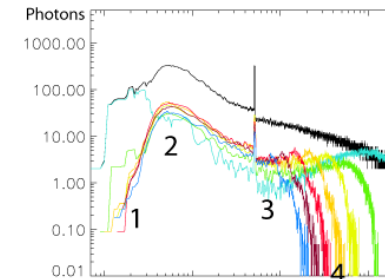


TGF spectral characteristics

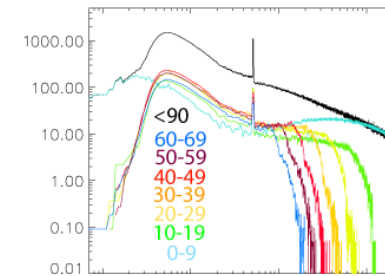
- 1) High energy cut-off moves to lower energies as escaping angles increase
- 2) Low energy cut-off moves to lower energies as TGFs are produced higher
- 3) Peak moves to lower energies
- 4) Bump around 0.7-1 MeV or flattening at 500 keV for TGFs produced deep disappears higher up

Beamed source
Discrete altitude

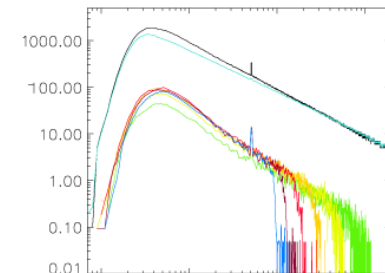
a)
15 km



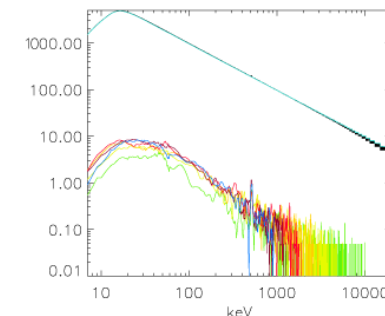
b)
20 km



c)
40 km



d)
60 km



Østgaard et al., 2008



ASIM Q4: Production altitude

Gjesteland et al., 2010

Table 1. Altitudes

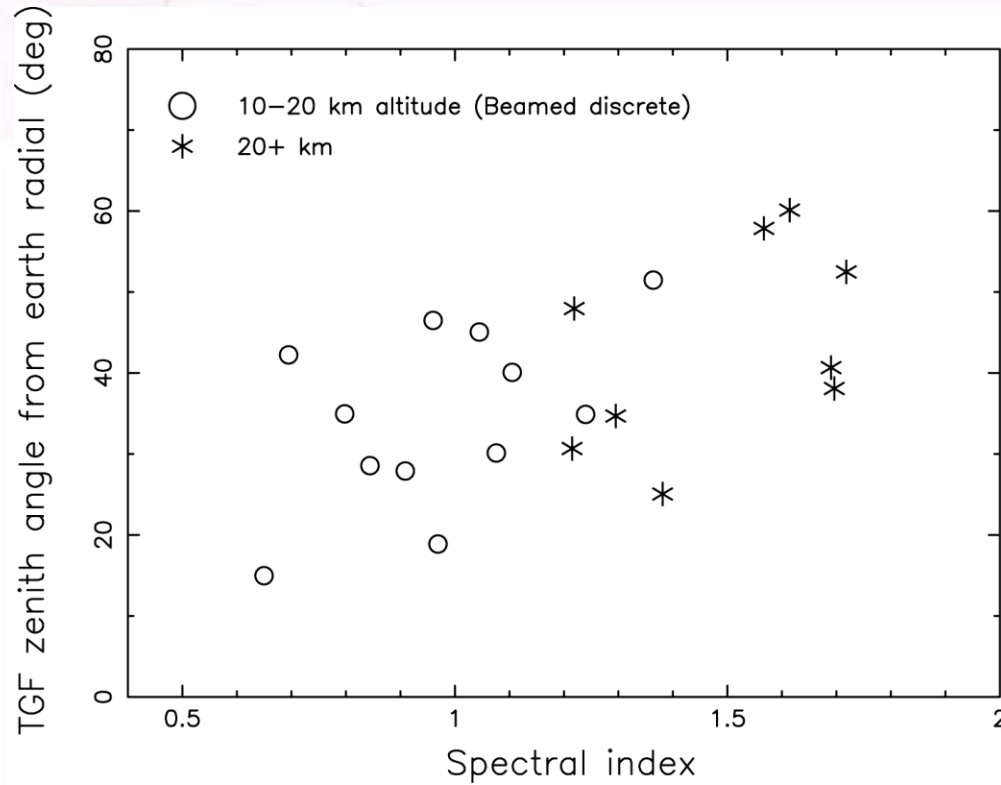
BATSE		Beamed					Cone				
TGF	TD [μ s]	\emptyset [km]	new[km]	TD [μ s]	Flux	DL[%]	\emptyset [km]	new[km]	TD [μ s]	Flux	DL[%]
2144	125 \pm 22	39	17	148 \pm 22	0.62	272	25	25	162 \pm 22	0.62	352
2370	124 \pm 18	40	16	117 \pm 21	0.36	208	41	22	96 \pm 12	0.36	230
2465	147 \pm 19	26	21	137 \pm 20	0.49	268	27	24	124 \pm 17	0.49	338
2955	145 \pm 40	39	23	118 \pm 16	0.44	216	38	26	80 \pm 21	0.51	277
5587	66 \pm 34	14	12	52 \pm 17	0.20	46	17	29	54 \pm 17	0.28	111

Results for 5 TGFs where BATSE deadtime was corrected, converge with results from others: 15-20 km

ASIM will give production altitudes for many single TGFs

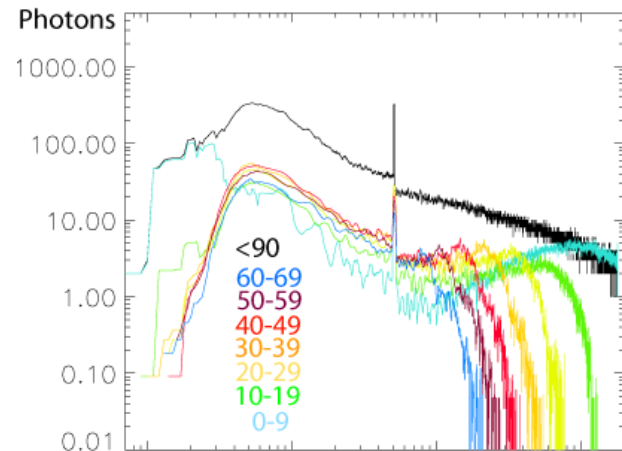


ASIM Q5: Spatial distribution of produced photons



BATSE analysis indicated beamed distribution – but this is still an open question

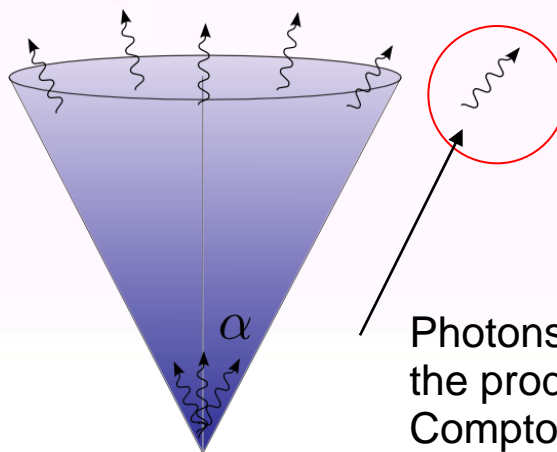
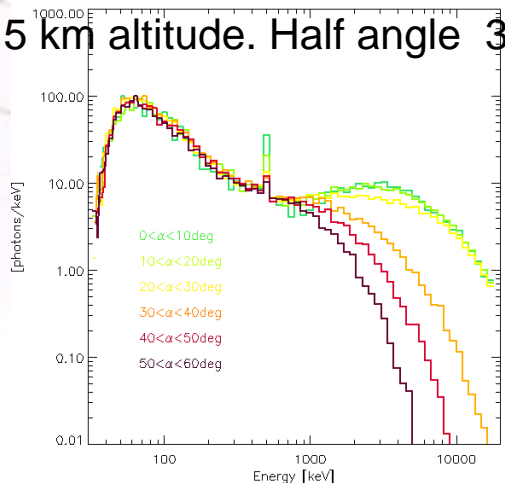
This is important for production mechanism.





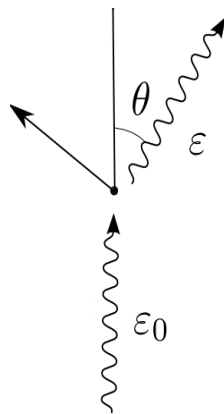
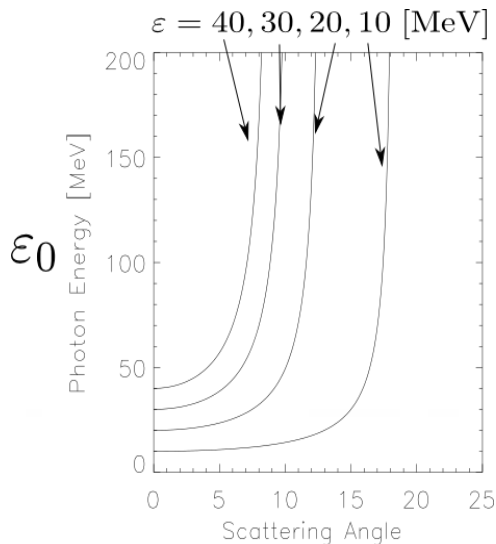
More on Spatial distribution of produced photons

15 km altitude. Half angle 30°



Photons measured outside the production cone are Compton scattered

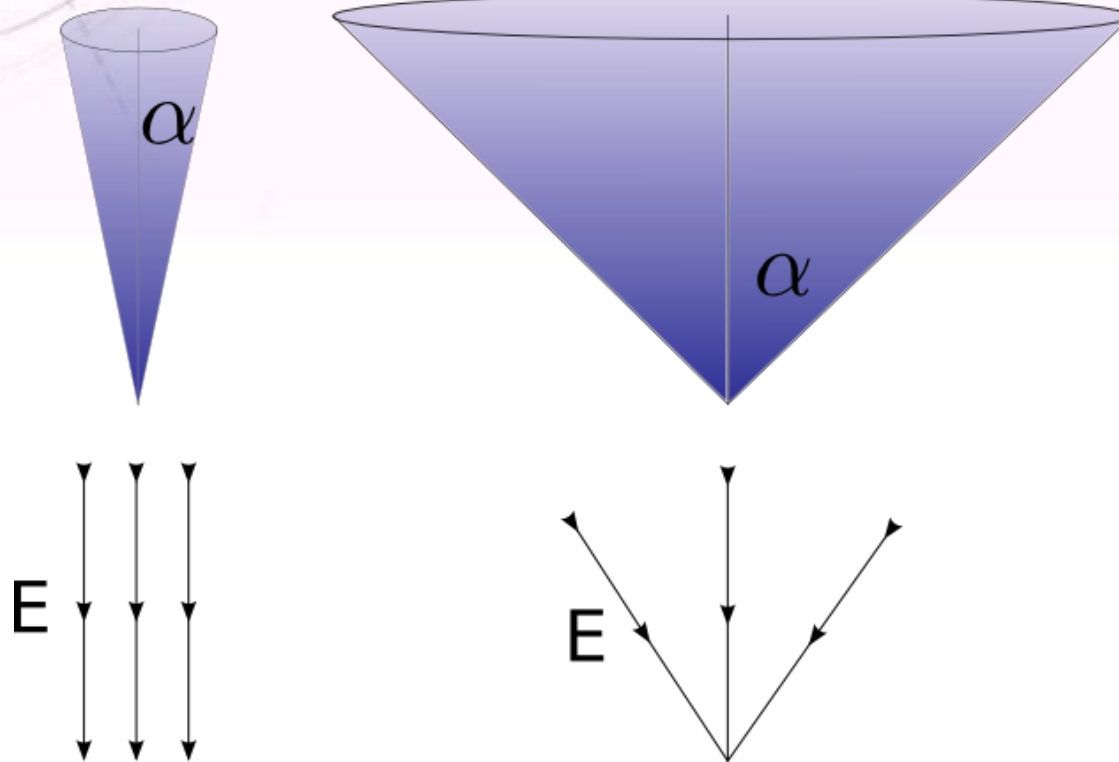
Energy loss in Compton scattering



- High energy photons ($E > 10$ MeV) measured at satellite altitudes are scattered less than 10° .
- If satellite measurements finds:
 - High energy cut off at increasing angles.
 - indicates narrow TGF production cone.
 - High energy photons measured at large angles.
 - Indicates wide TGF production cone.



Spatial distribution of produced photons and E-field



Electrons are accelerated in the direction of $-E$. The geometry of the E-field defines the production cone.

Superposition of TGFs cannot tell - it could be superposition of narrow beams in different directions



Questions - summary

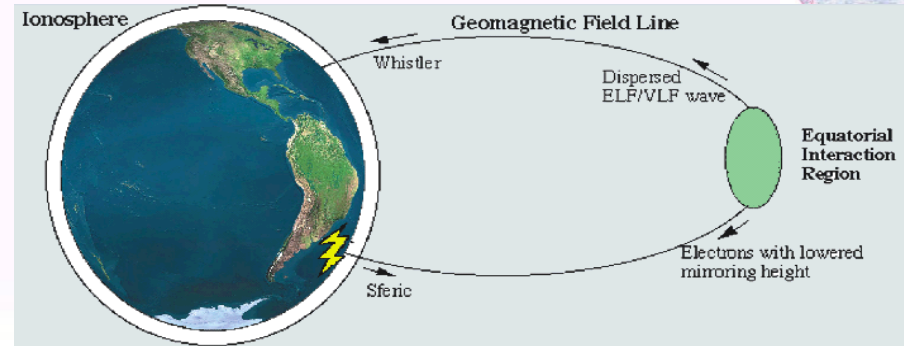
1. How common are TGFs
 - Give hint towards type of lightning
2. Are there TGFs without lightning
 - Streamer??
3. What type of lightning
 - Combine with other measurements
4. Production altitude
 - Another hint towards type of lightning
5. Spatial distribution of produced photons
 - Geometry of electric field generating bremsstrahlung



Secondary Science

Whistler-induced electron precipitation

- Lightning induced precipitation (LEP)– cyclotron resonance
- MXGS will see the soft LEP

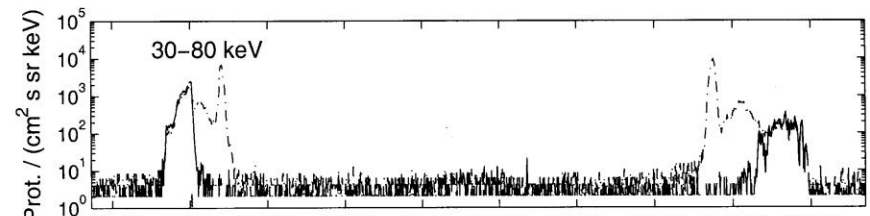
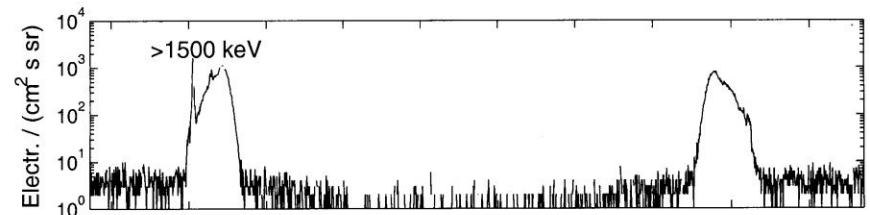
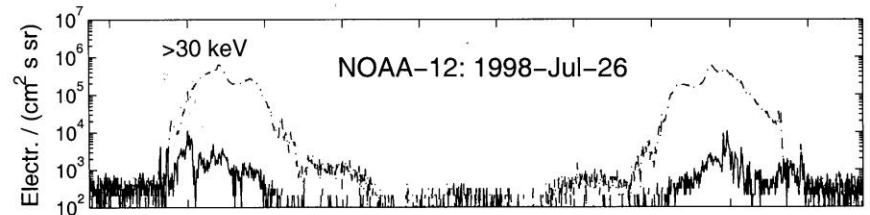


Relativistic Electron Precipitation

- Continuous and burst
- MXGS will see the continuous (500 times more intense than TGFs)

Aurora

- MXGS will see this, of course.



UT	20701	121202	121702	122202	122702	123202	123702	124202	124702	125201
GLon	113.9	99.2	92.4	87.7	83.7	79.8	75.3	69.2	57.1	6.4
GLat	-70.9	-54.2	-36.9	-19.4	-1.7	15.9	33.5	50.9	67.8	81.0
ILat	-85.4	-68.6	-51.4	-33.9	-17.6	15.0	30.4	47.8	63.9	78.6
MLT	1815	1743	1737	1734	1730	1726	1720	1712	1654	1549



Future plans

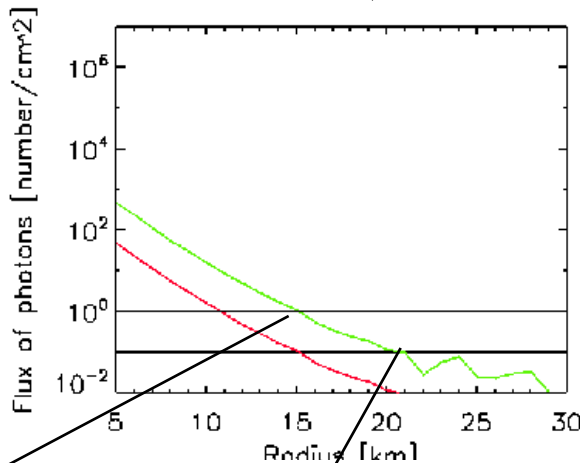
- Balloons – French COBRAT
 - Aircraft – several options if money is found
-
- X- and gammaray detector – 1 microsec res
 - Electric Field measurement – 1 microsec res
-
- Want to see if TGF is before the discharge – return stroke, which would indicate streamer/leader tip generation
 - TGF is after the discharge – favour the feed-back theory, where positrons and downgoing gamma rays contribute to avelanche process.



Aircraft - Balloon

Production: 12 km, Observed: 19 km

- Flux of photons at production is 10^{17}
- modeling and analytic solution based on RHESSI.
- Here: 250 cm^2 detector

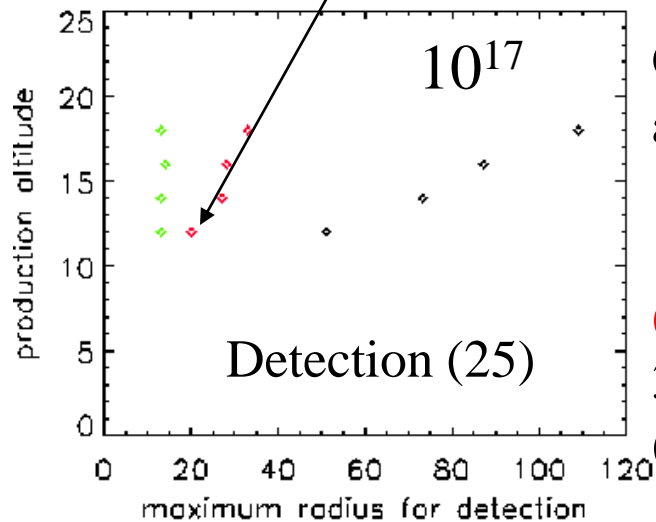
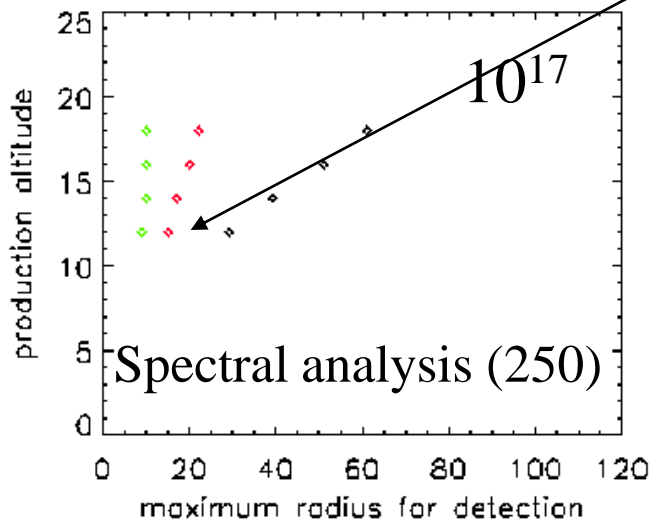


10^{17}

10^{16}

$1 \text{ cm}^{-2} - 250 \text{ phts}$

$0.1 \text{ cm}^{-2} - 25 \text{ phts}$



Observation altitude:

15 km

$19 \text{ km} - \text{aircraft}$
(radius: $15-20 \text{ km}$)

$35 \text{ km} - \text{balloon}$
(radius: $30-50 \text{ km}$)



Thanks