



## **Observation of TGFs** with AGILE: the current status

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### Outline



- TGF science: an overview
- AGILE-MCAL instrument and detection capabilities
- AGILE strength points in TGF science
- Characteristics of the AGILE TGF sample
- Comparison with the RHESSI and Fermi-GBM
- Work in progress
- Conclusions



# Terrestrial Gamma-ray Flashes (TGF)



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Gamma-ray flashes with incoming direction compatible with the Earth surface.

Few millisecond typical duration; hard spectrum (up to tens MeV)

Discovered by BATSE (Fishman et al., Science, 1994) and observed by RHESSI up to 20 MeV (Smith et al., Science, 2005)

Clearly associated to lightning discharges during thunderstorms by means of correlation with VLF sferic waves detection on ground (Inan et al., GRL, 1996; Cummer et al., GRL, 2005)

>2008: Observed by AGILE and Fermi-GBM

2009: AGILE reports energy up to 40MeV (Marisaldi et al., JGR 2010)

Geophysical phenomena observed from space by instruments designed for gamma-ray astrophysics

Challenging detection: timing and energy range are key issues



#### **1994: BATSE discovery of TGF**



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Fishman et al., Science, 1994

~ 70 TGF detected on 9 life-Years typically 100 counts/TGF

Main limitations:

- On-Board Trigger Logic performances (shorter timescale 64ms)

- Large statistics BUT only 4 energy bins for time-tagged events



#### 2005: RHESSI detection up to 20 MeV





TGF Distribution with lighting frequency per km<sup>2</sup> per Year

#### Smith et al., Science, 2005



Contiuous time-tagged event list NO ON-BOARD TRIGGER LOGIC 10– 20 TGF per month Typically 20-30 counts/TGF ~800 TGFs reported in the 1<sup>st</sup> RHESSI TGF catalog (Grefenstette et al., JGR, 2009)



## **RREA Production model**

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Dwyer and Smith, GRL (2005) Carlson, Lehtinen and Inan (2007)

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Relativistic Runaway Electron Avalanche (RREA) with relativistic feedback (Dwyer 2008)

Bremsstrahlung + Compton scattering

RHESSI cumulative spectrum is compatible with a production altitude of 15-21 km (just above tropical thunderstorms)

Still hint for individual spectral variability: differences in production altitudes or viewing angle?

BATSE events seem produced at higher altitude (two different populations?) but discrepancy is reduced if dead-time effects are properly accounted for (Grefenstette et al., 2008; Ostgaard et al., 2008)

#### The AGILE Mini-Calorimeter (MCAL)





30 CsI(Tl) bars with Photodiode readout, like these 1400 cm<sup>2</sup> geometrical area ~300 cm<sup>2</sup> effective area @ 1 MeV 330 keV – 100 MeV energy range 14% energy resolution FWHM @ 1.3 MeV 2 μs timing accuracy in photon-by-photon mode Clever, fully-programmable trigger logic on time scales from 8s to 16ms, 1ms and 300μs

> Labanti et al., NIM A (2009): instrument paper Fuschino et al., NIM A (2008): trigger logic Marisaldi et al., A&A (2008): GRB detections Marisaldi et al., JGR (2010): TGF detections



8th AGILE WS 28 Apr. '10



# Why AGILE is good for TGF science?



- MCAL energy range is extended up to 100 MeV: probing the high energy tail of the TGF spectrum
- Efficient trigger at ms and sub-ms time scale (the TGF time scale): not biased toward brightest events
- segmented independent detectors: low dead time and pile-up



# Why AGILE is good for TGF science?



- photon-by-photon data download for triggered events with 2µs time resolution
- <100µs absolute timing accuracy: mandatory for sferics correlation
- AGILE orbit at 2.5° inclination is optimal for mapping the equatorial region, where most of the events take place, with unprecedented exposure



### MCAL TGF detection rate



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#### 153 good candidates between June '08 – Mar. '10



34 TGFs Published in M. Marisaldi et al., J. Geoph. Res., 115, A00E13, 2010.

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After entering Spinning mode



## The AGILE TGF sample





#### Average properties:

Number of counts = 17.3 +/- 6.4

Duration = (1.7 +/- 0.9) ms

Energy = (4.0 +/- 1.7) MeV

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## **TGFs and Lightnings**



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LIS-OTD High Resolution Full Climatology available at http://thunder.msfc.nasa.gov/data/ Good match between AGILE TGF pattern and lightning map

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TGF production <~300km close to sub-satellite point, Cummer et al., GRL (2005)



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## AGILE vs RHESSI: longitude and local time



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1<sup>st</sup> RHESSI TGF catalog Grefenstette et al., JGR, (2009)

selected RHESSI TGFs in a +/- 2.5° latitude belt (like AGILE orbit)  $T_0 < 1^{st}$ Jan. 2006: 84 TGFs

Longitude and local time distributions are compatible

> double peaked feature on South East Asia

sharp cut on western Africa

late afternoon occurrence peak



### AGILE vs Fermi-GBM



#### Fermi TGF #7

#### AGILE TGF 11026-1

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#### Fermi:

+ larger effective area and lower threshold: more statistics on single events

- trigger on time >=16ms: less events, brightness bias (AGILE triggers on >=290µs)

from A. Von Kienlin, presentation at the 7<sup>th</sup> AGILE WS

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#### AGILE vs Fermi-GBM



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(more since Dec.'09)





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All range

<0.7 MeV

0.7-1.4

.4-2.8

>2.8 MeV

#### Trigger 11026-1 in details INAF Light curve **Position distribution** Energy vs time Plane X total position distribution Plane 7 total position distributio Energy vs time 40 MeV T0 = 2009-06-12 11:01:59 UT 171889319.034 TT Energy (MeV) Orbit: 11026 Energy range: 0 - 2 Selected bars: 7FFF 7FFF 200 MCAL Team - INAF/IASF-B ۸ 10 0.03 0.035 Plane X 0.00 - 0.70 MeV Plane Z. 0.00 - 0.70 MeV 0.00 - 0.70 MeV -5 0 5 10<sup>-1</sup> Plane X 0.70 - 1.40 MeV lane Z 0.70 - 1.40 MeV 0.014 0.0145 0.015 0.0155 0.016 0.0165 0.017 0.0175 0.018 0.0185 0.70 - 1.40 MeV Bar address vs time bars vs time g 0.01 0.015 0.02 0.025 0.03 0.035 t - 10 (s) ++ 0-20-15 -10 -5 0 5 Position (cm) 0 -0.005 0.005 6 8 10 bar Plane Z 1.40 - 2.80 MeV Plane X 1.40 - 2.80 MeV 15 1.40 - 2.80 MeV Position (cm) 4 0 5 0 0.5 0 -0.005 0.025 0.03 0.035 0.015 t - t0 (s) 6 8 10 -10 -5 0 5 Position (cm Plane X 2.80 - 1433.60 MeV Plane Z. 2.80 - 1433.60 MeV 2.80 - 1433.60 MeV

0-20-15 -10 -5 0 5 10 Position (cm) 0.014 0.0145 0.015 0.0155 0.016 0.0165 0.017 0.0175 0.018 0.0185

t-t ini (s)

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0.005

0.01

0.015 t - t0 (s) 0.02 0.025 0.03 0.035



011026-1 spectrum is compatible with cumulative spectrum

Need more bright events

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0 Parameter: Pholndex 0.5

2002

-0.5

#### Increasing statistics are we looking at the tip of the iceberg? IASF Bologna



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# Work in progress



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- on-line TGF archive
  possibility of TGF alerts
  link with meteo data
- study of the TGF mechanism





Scientific analysis:

Apply calibration

Spurious trigger rejection

Standard products (light curves, event topology, intensity duration, hardness ratio)

Write trigger info to webaccessible DB

**Issue alert** 

Very fast real-time alert system:

In 20-30 minutes from telemetry download an alert to the community can be issued Suitable for real-time correlation with meteo observations Potentially useful for air navigation: add high energy information to storm-tracking monitoring

systems

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# Work in progress: INAF CONTROL INAF CONTROL INAF CONTROL INAF CONTROL INAF CONTROL INAF BOLOGNA

#### Current triggers selection geographical map here)



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- AGILE is an important instrument for TGF science:
  - the only one with energy range extended up to 100 MeV
  - <u>the only one</u> with <1ms trigger logic
  - photon-by-photon with  $\mu s$  timing
  - -~equatorial orbit
- AGILE detects ~10 TGFs / month with current selection criteria. Rate can be at least doubled with improved selections
- AGILE and RHESSI TGF samples are consistent concerning longitude, local time distribution and spectral shape
- Cumulative spectrum with significant detection above 40 MeV: challenge for production models
- A real-time monitoring and alert system can be implemented for correlation with other meteo resources