

Molecular clouds and gamma ray emission: Young TeV γ -ray SNRs

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Y.F.

The origin of cosmic ray protons

Hadronic vs. leptonic (electron's inverse Compton etc.)

Hadronic $p + p \Rightarrow \pi^0 \Rightarrow 2\gamma$ seems promising

e.g., energy spectrum by AGILE, Fermi, HESS etc. of W44

Two young TeV γ -ray SNRs, 1600-3000yrs, non-thermal X rays:

RXJ1713.7-3946, brightest HESS source

RXJ0852.0-4622 [Vela Jr.]

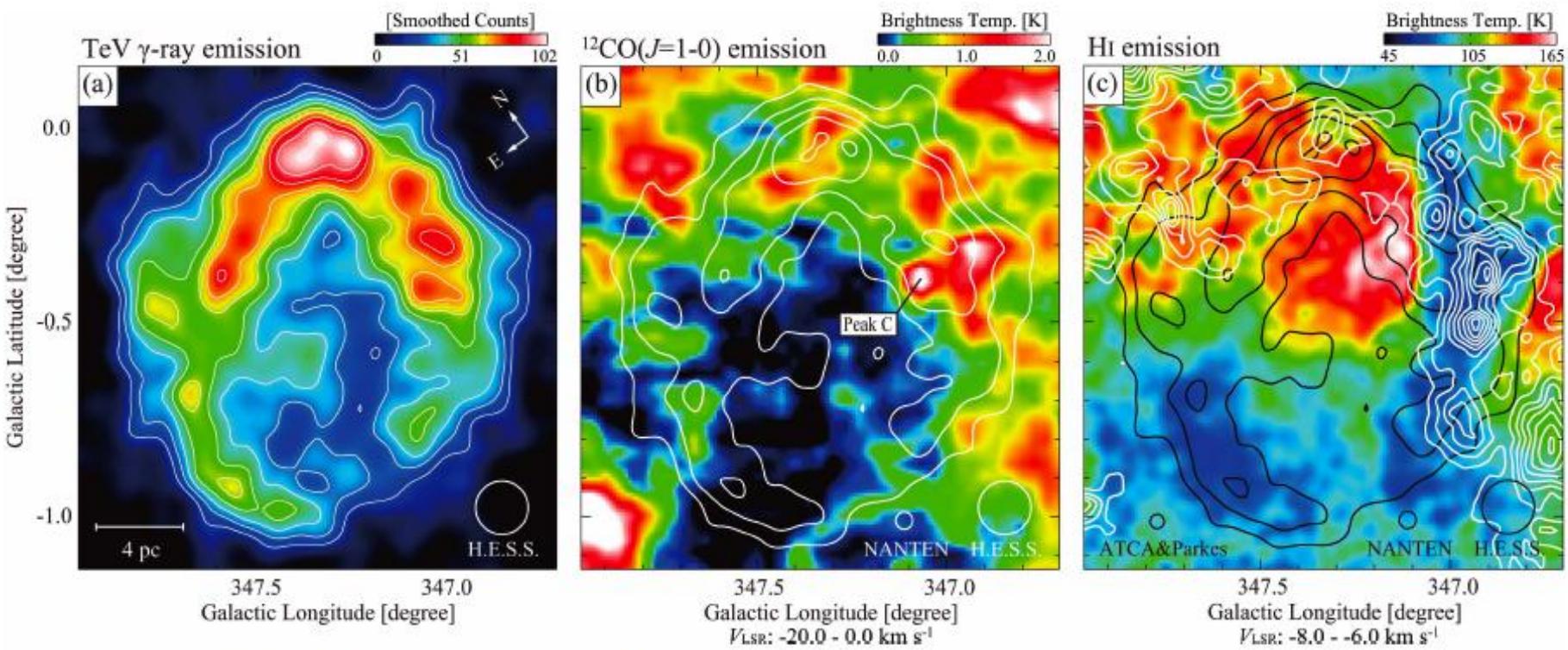
If hadronic, target ISM protons correspond to γ rays.

If leptonic, non-thermal X rays correspond to γ rays.

Collaborators

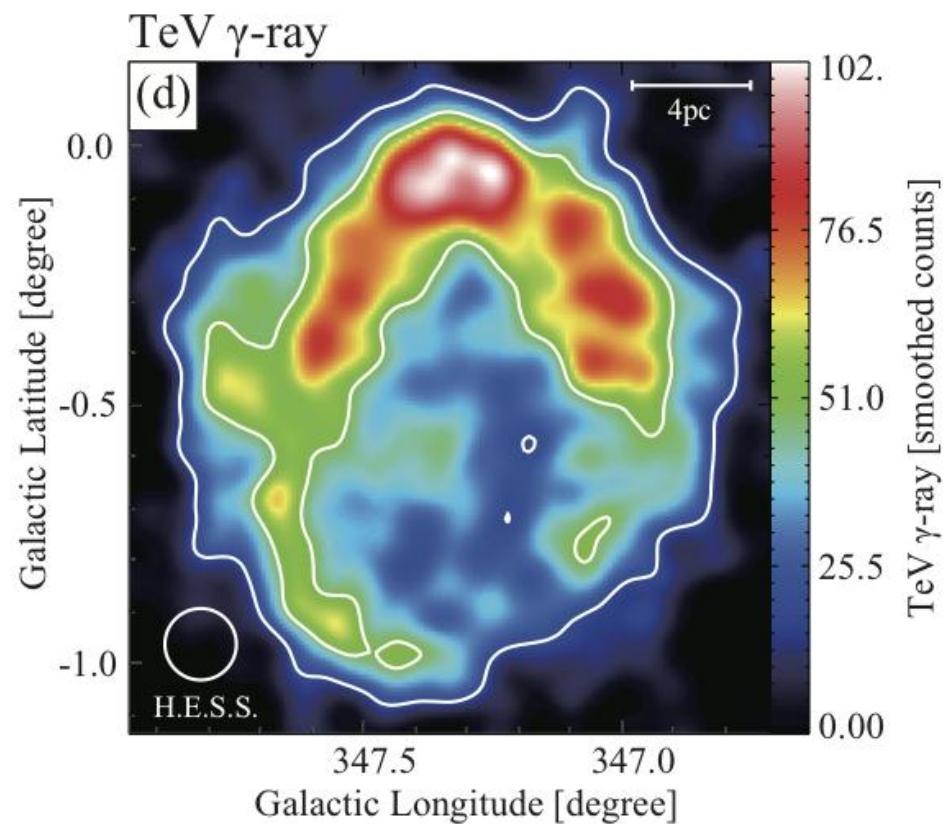
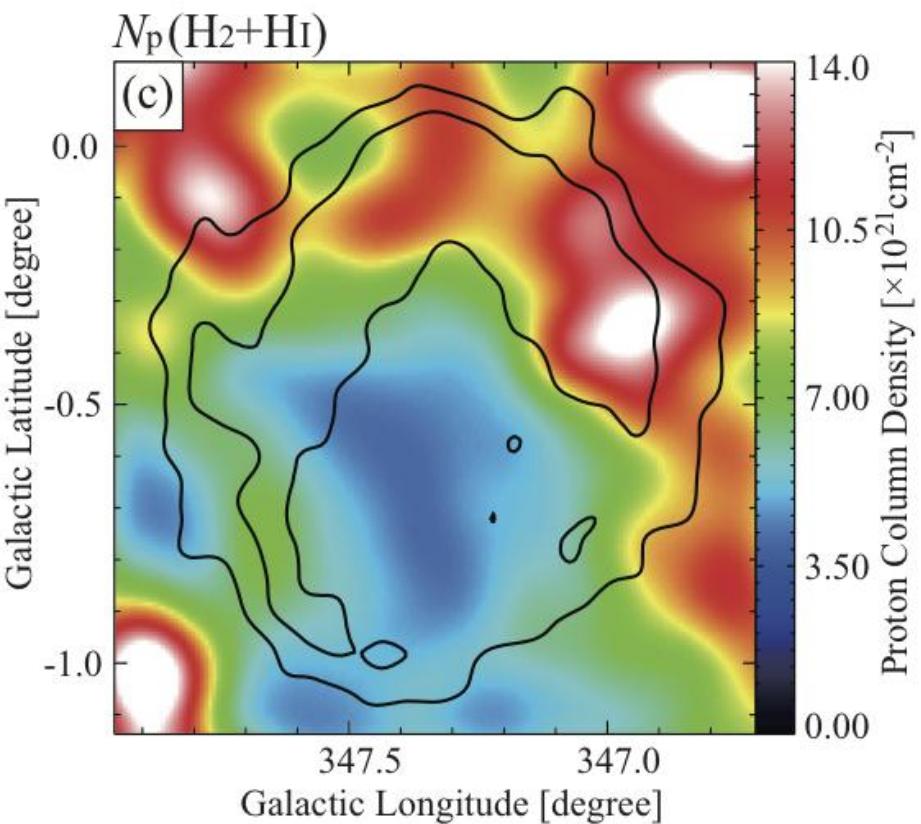
- AGILE team: M. Tavani, A. Giuliani +
- HESS team: F. Aharonian, G. Rowell +
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TeV γ -ray SNR RXJ1713

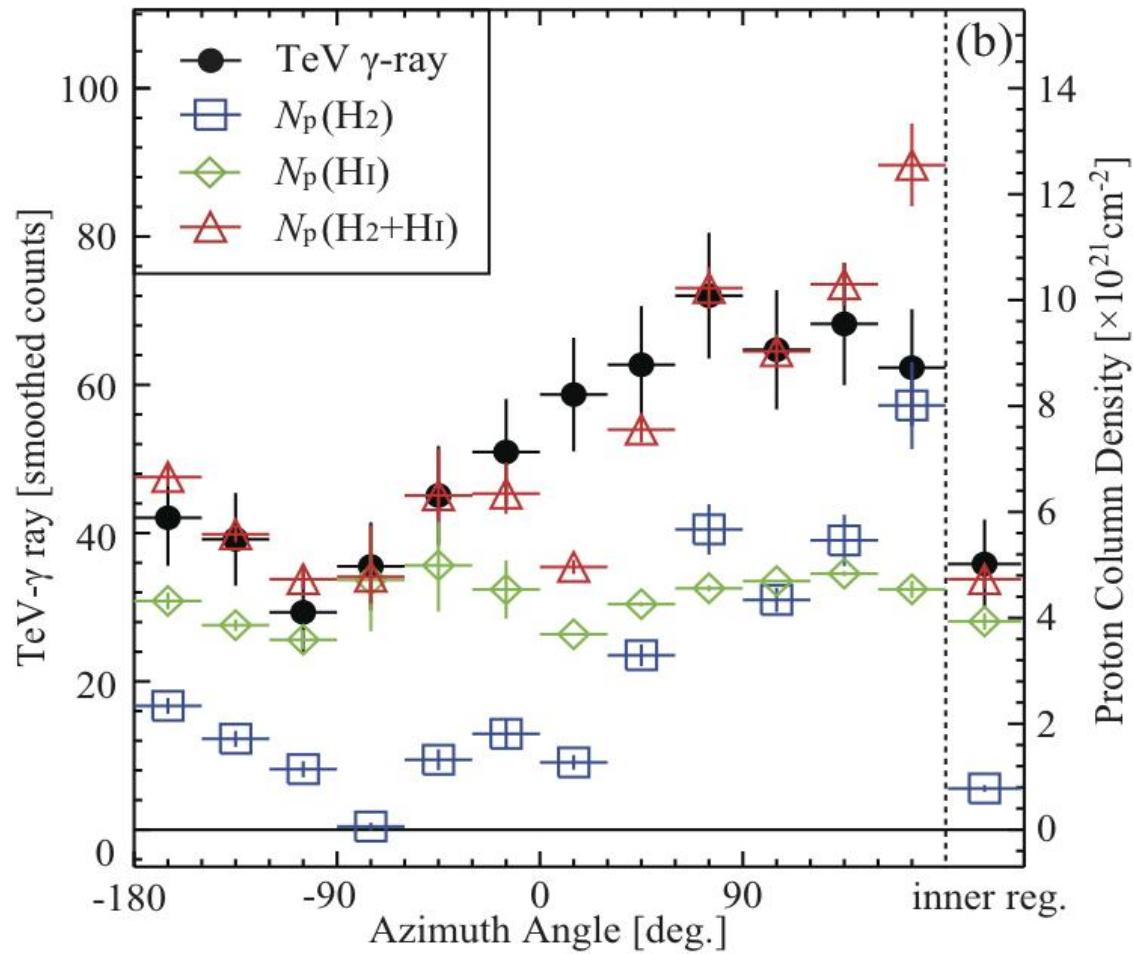


Fukui et al. 2012, ApJ, 746, 82

TeV γ -Ray SNR RXJ1713 ISM Proton Column Density Distributions



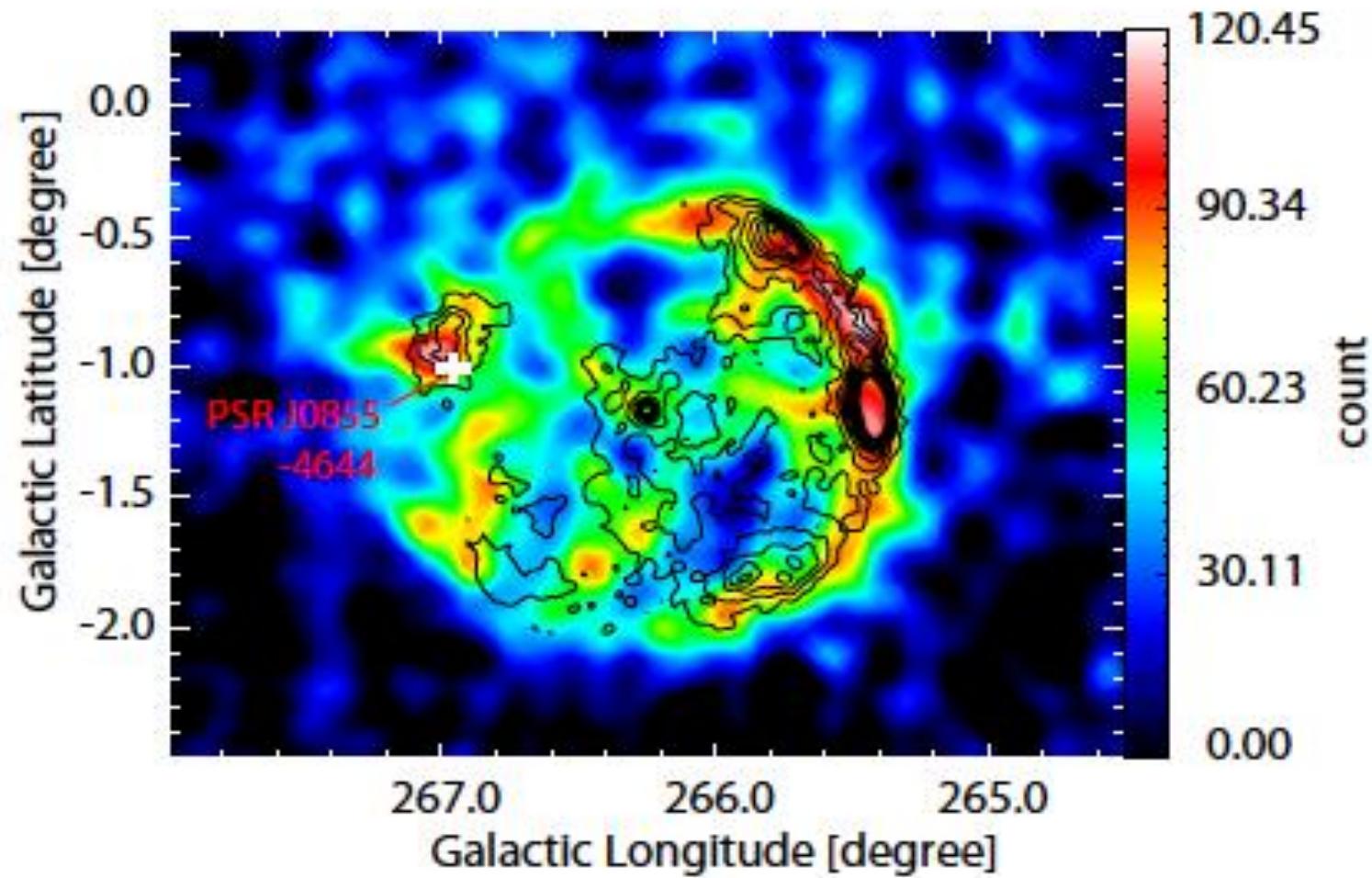
TeV γ -ray SNR RXJ1713 ISM Proton and TeV γ -ray Distributions



γ rays
ISM protons
good correspondence

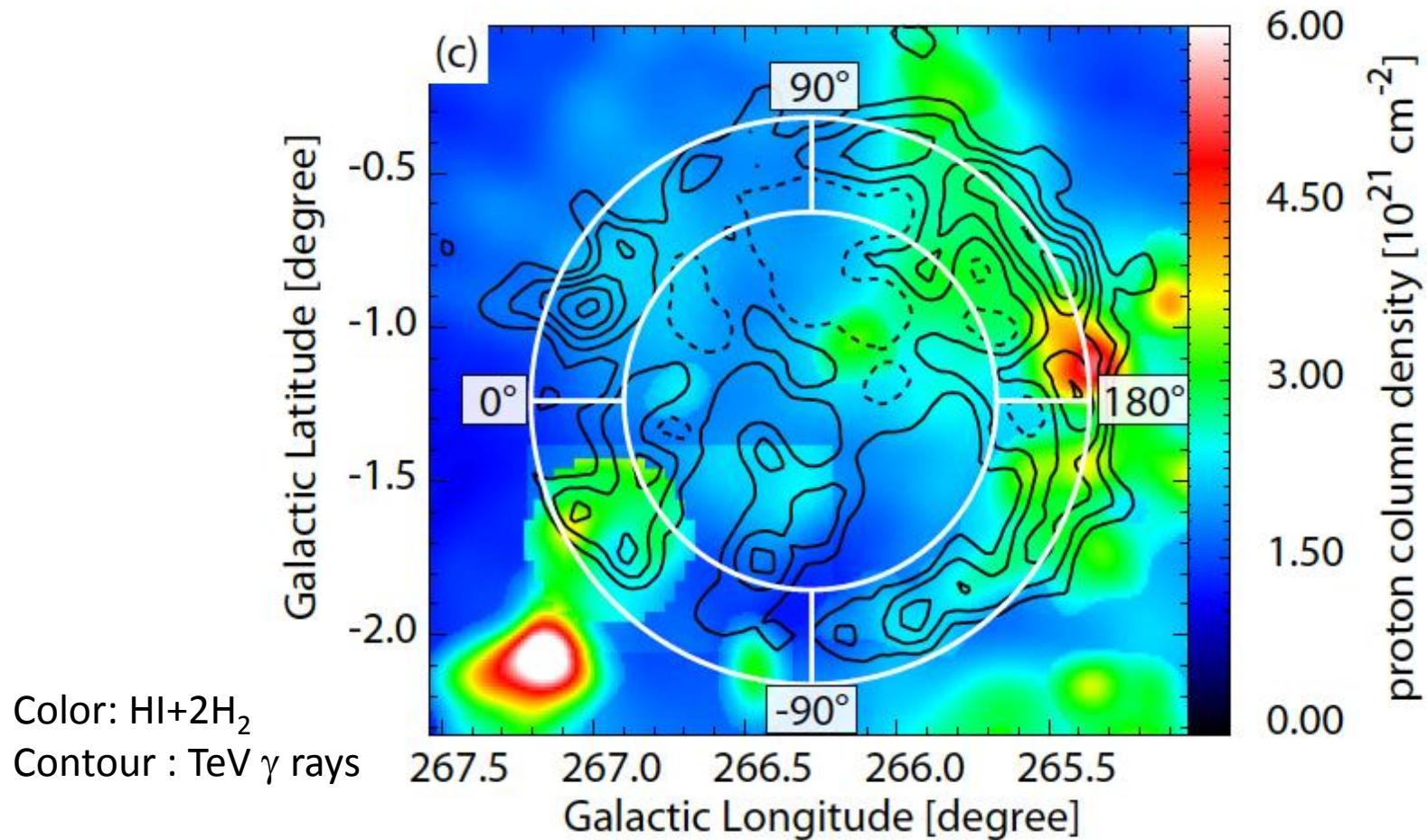
ISM protons
as targets for cosmic
ray protons

TeV γ -ray SNR RX J0852

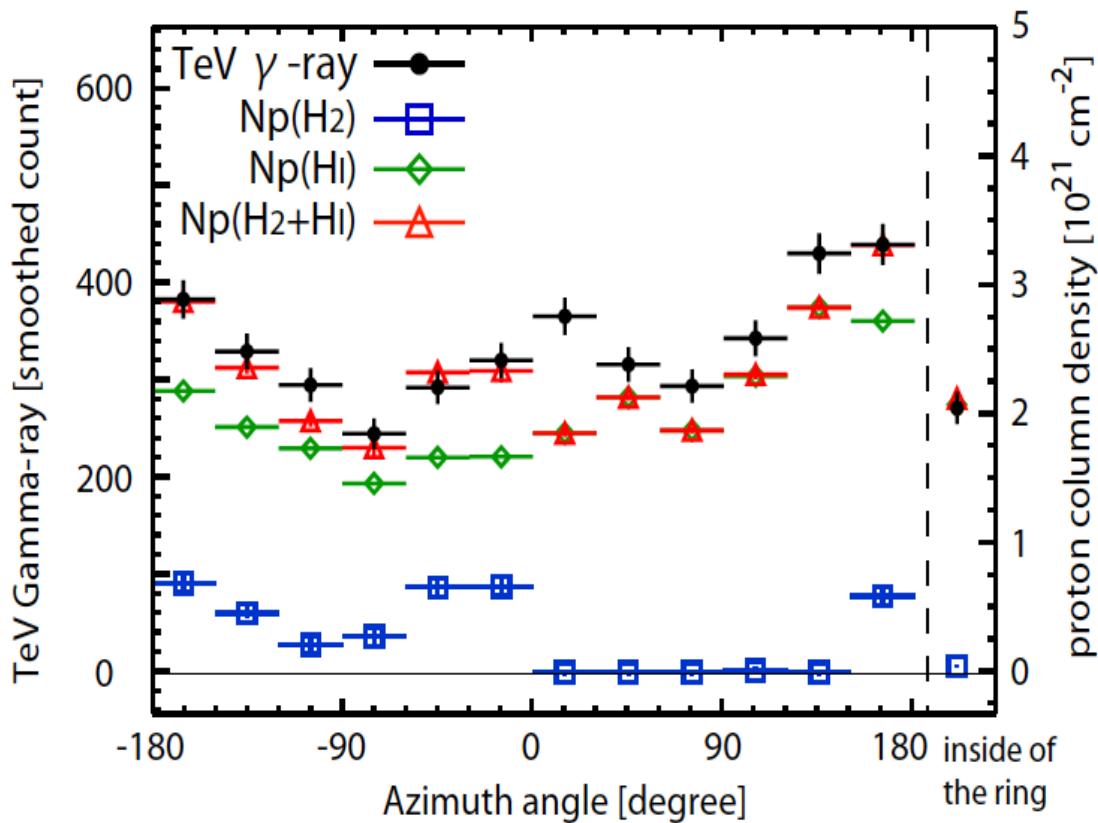


TeV γ -Ray SNR RX J0852 ISM Proton Column Density Distributions

Fukui et al. 2012, in prep.



TeV γ -ray SNR RX J0852 ISM Proton and TeV γ -ray Distributions

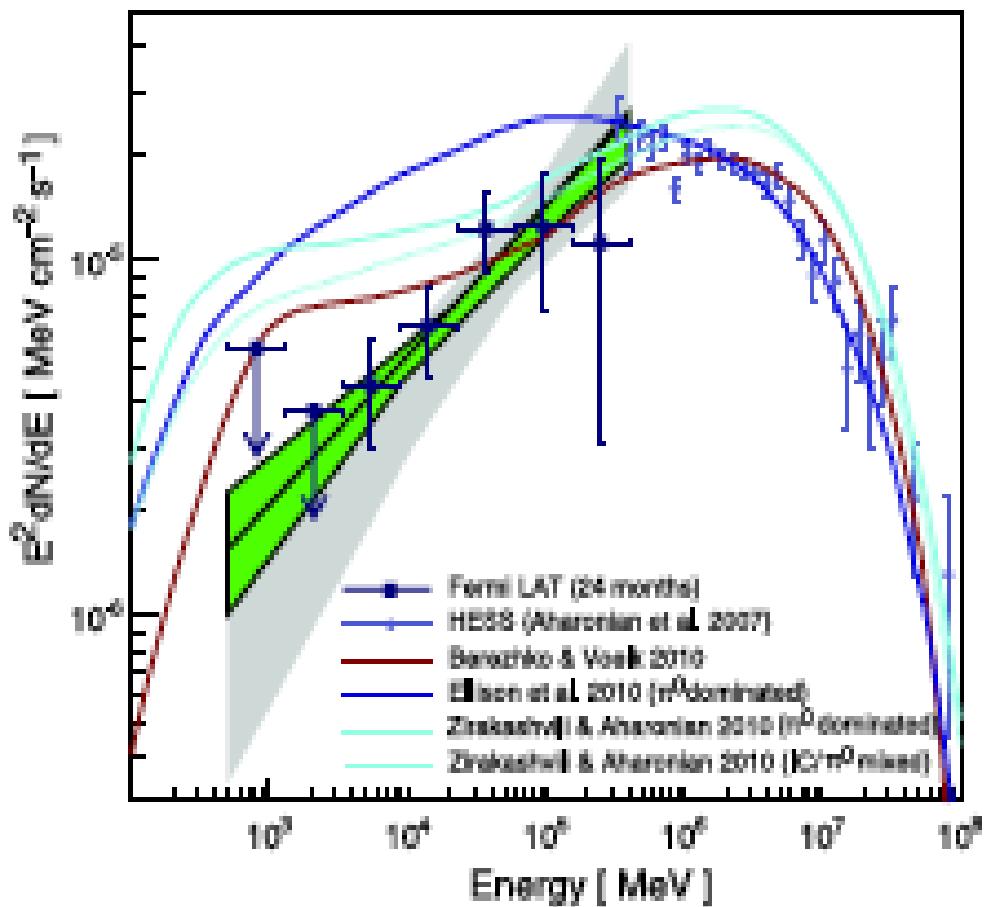


γ rays
ISM protons
good correspondence

ISM protons
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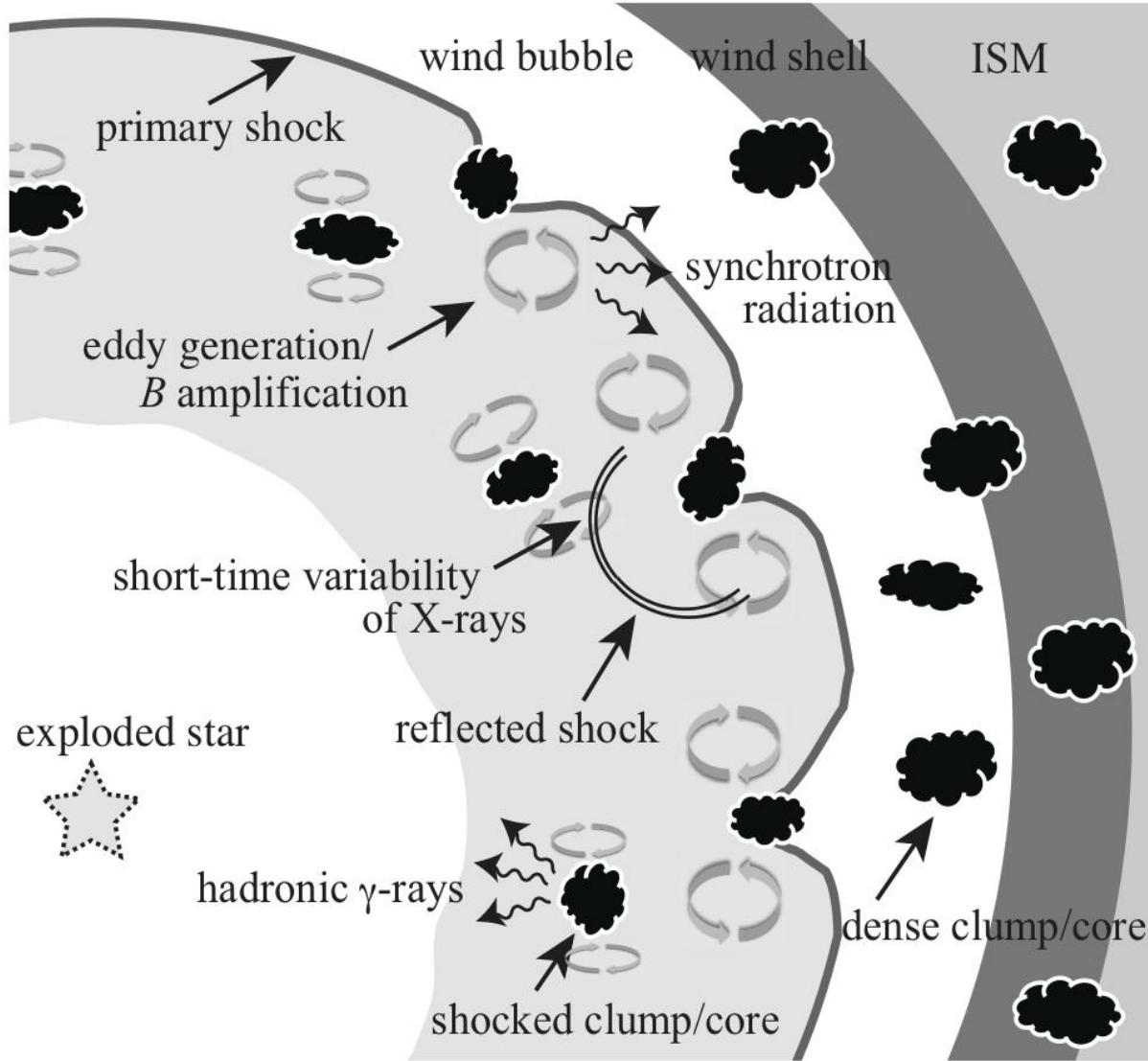
γ -ray spectrum of RX J1713

Abdo et al. 2011



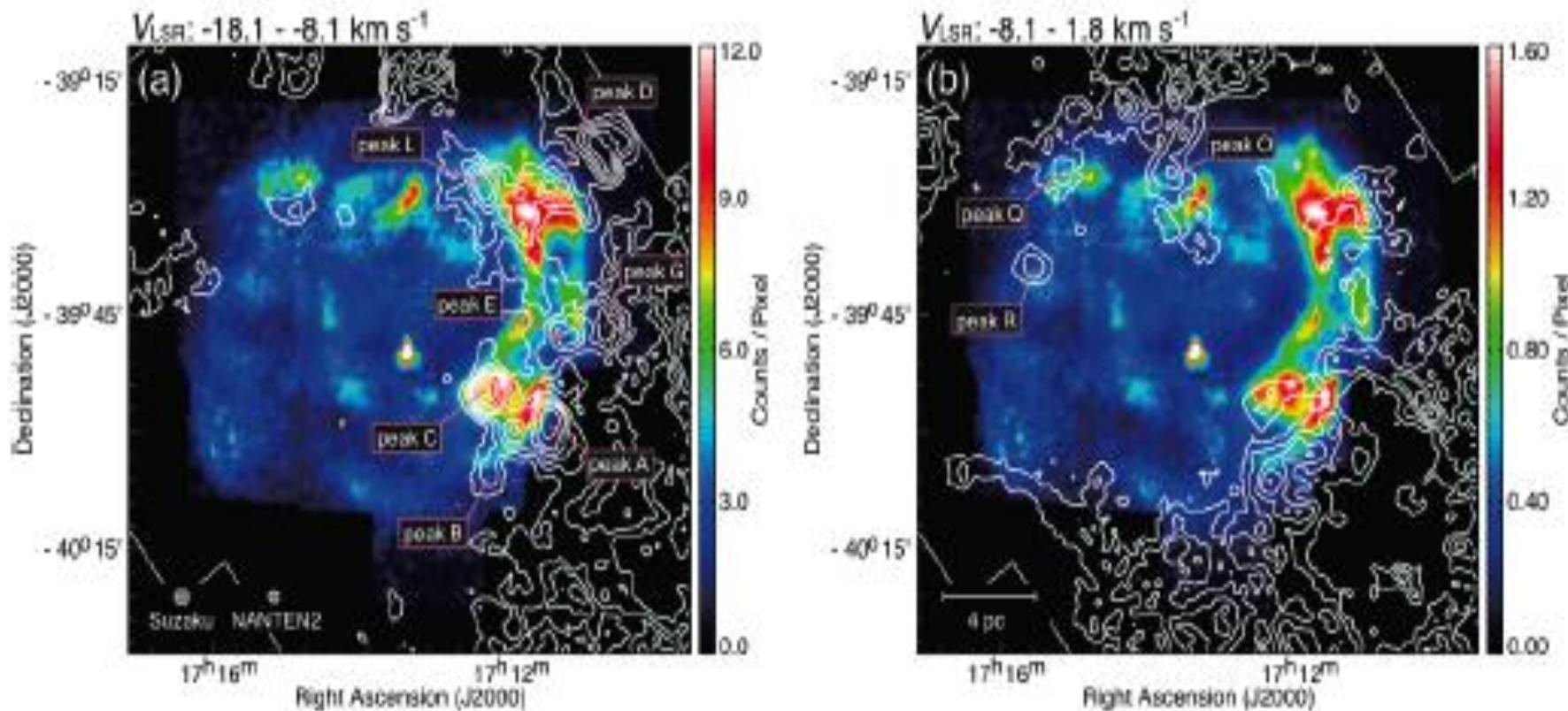
The hard spectrum is not unique to the leptonic scenario

The hard spectrum is explained by energy dependent penetration of CR protons into dense molecular gas.



Inoue, Yamazaki, Inutsuka, Fukui 2012, ApJ, 744, 71

Suzaku X-rays (1–5 keV) and Molecular Clouds



Color Image: *Suzaku* XIS 0+2+3 1–5 keV count map

Contours: $^{12}\text{CO}(J=2-1)$ integrated intensity (a) $V_{\text{LSR}}: -18.1 \text{--} -8.1 \text{ km s}^{-1}$ (b) $V_{\text{LSR}}: -8.1 \text{--} 1.8 \text{ km s}^{-1}$



18 arcsec のガウシアンでスムージング済 (this work についてはスムーズ前のデータからの動径分布の計算 + エラー評価が可能)

RX J1713.7–3946 Azimuthal Plot (TeV gamma-rays and X-rays)

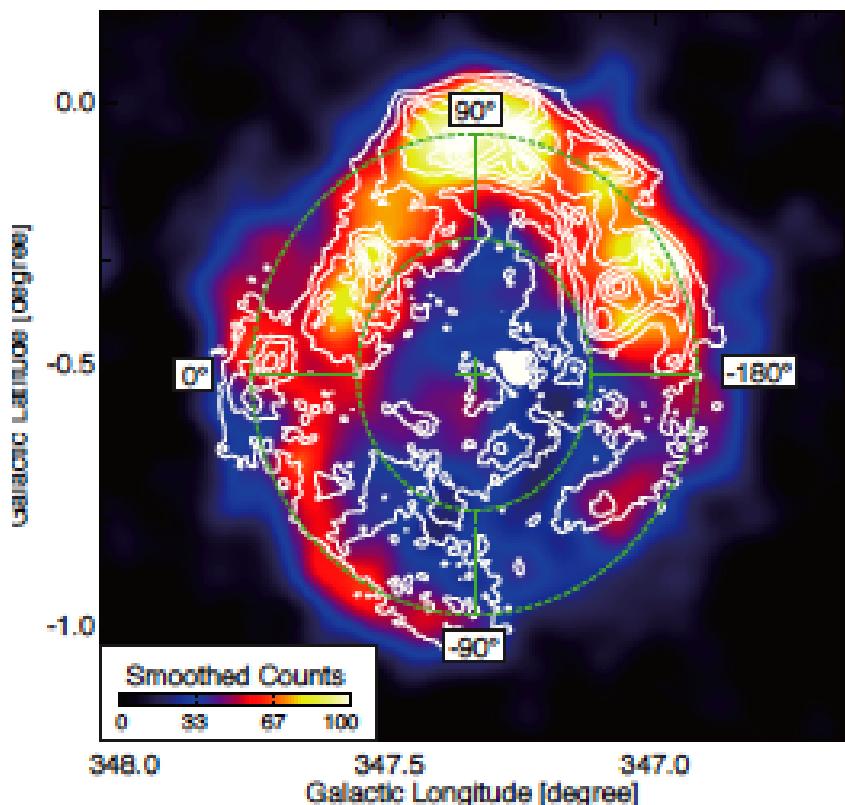
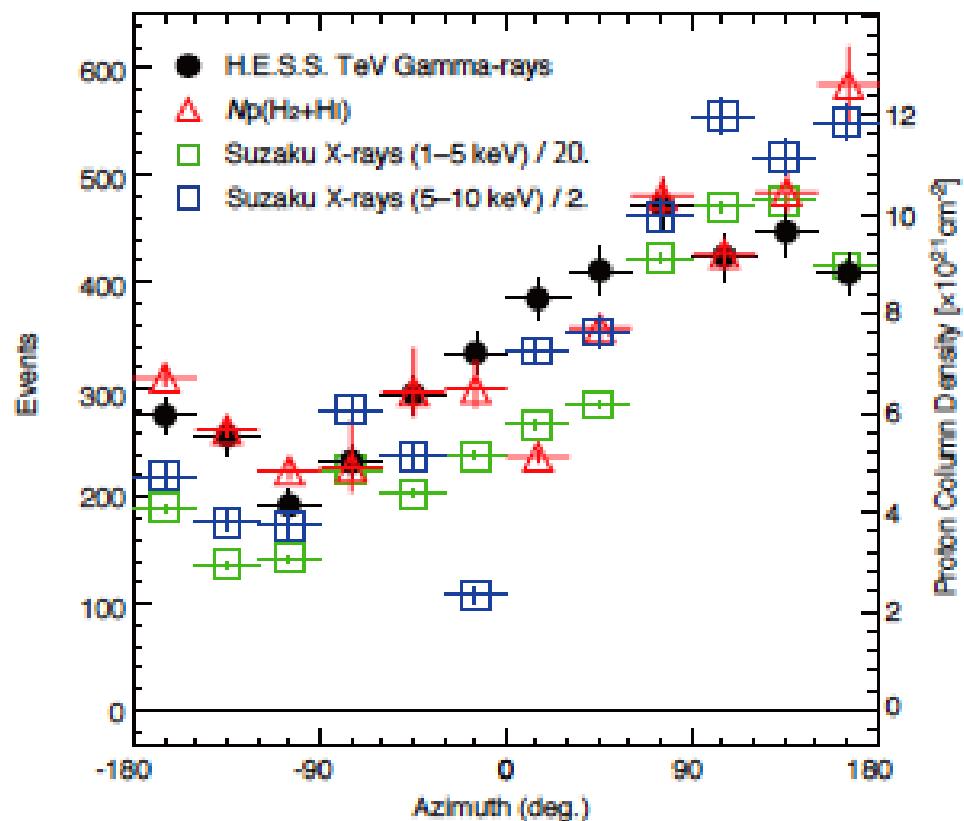


Image: H.E.S.S. TeV gamma-rays (Aharonian et al. 2007)
 Contours: Suzaku X-rays (1–5 keV)

Azimuthal plot (Fukui et al. 2012)
 center: (l, b) = (347.34 deg., -0.52 deg.)
 outer ring radius: semi-major 0.46 deg., semi-minor 0.42 deg



RX J1713.7-3946 Radial Profiles

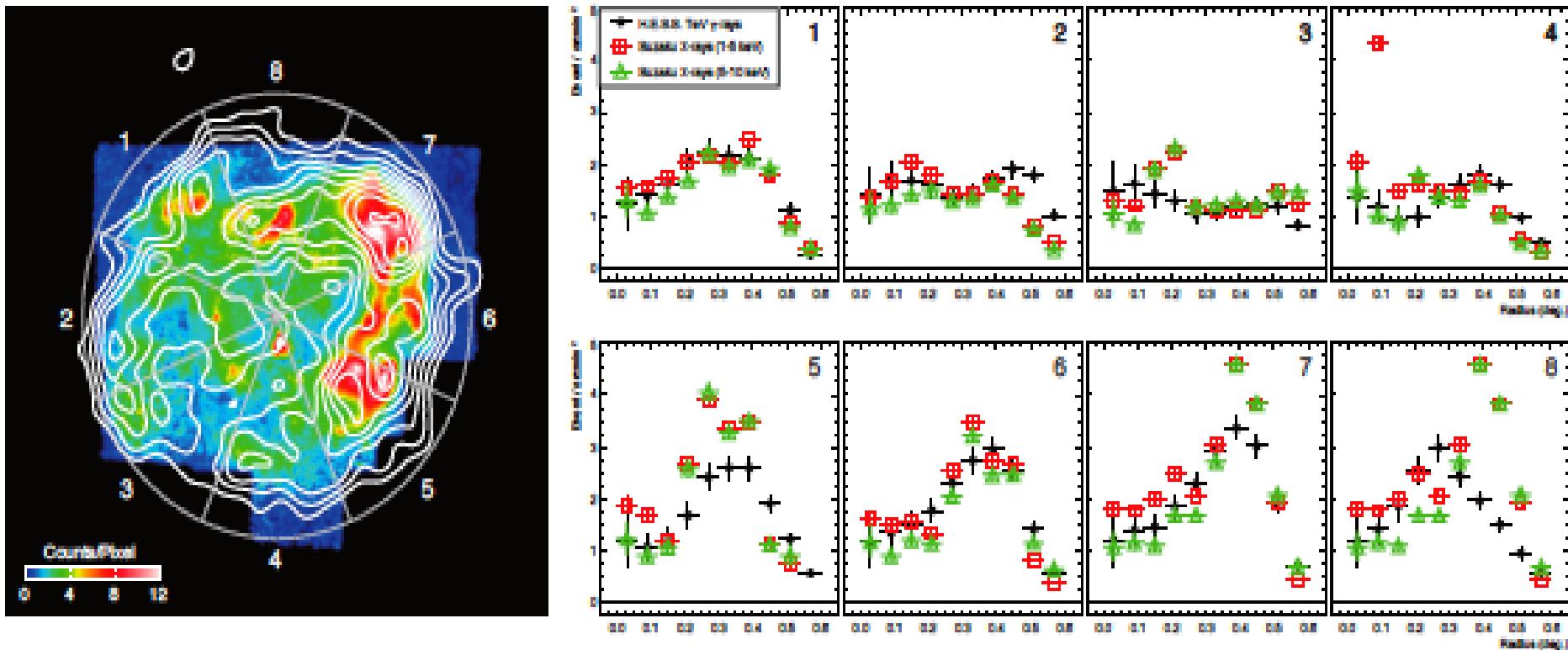
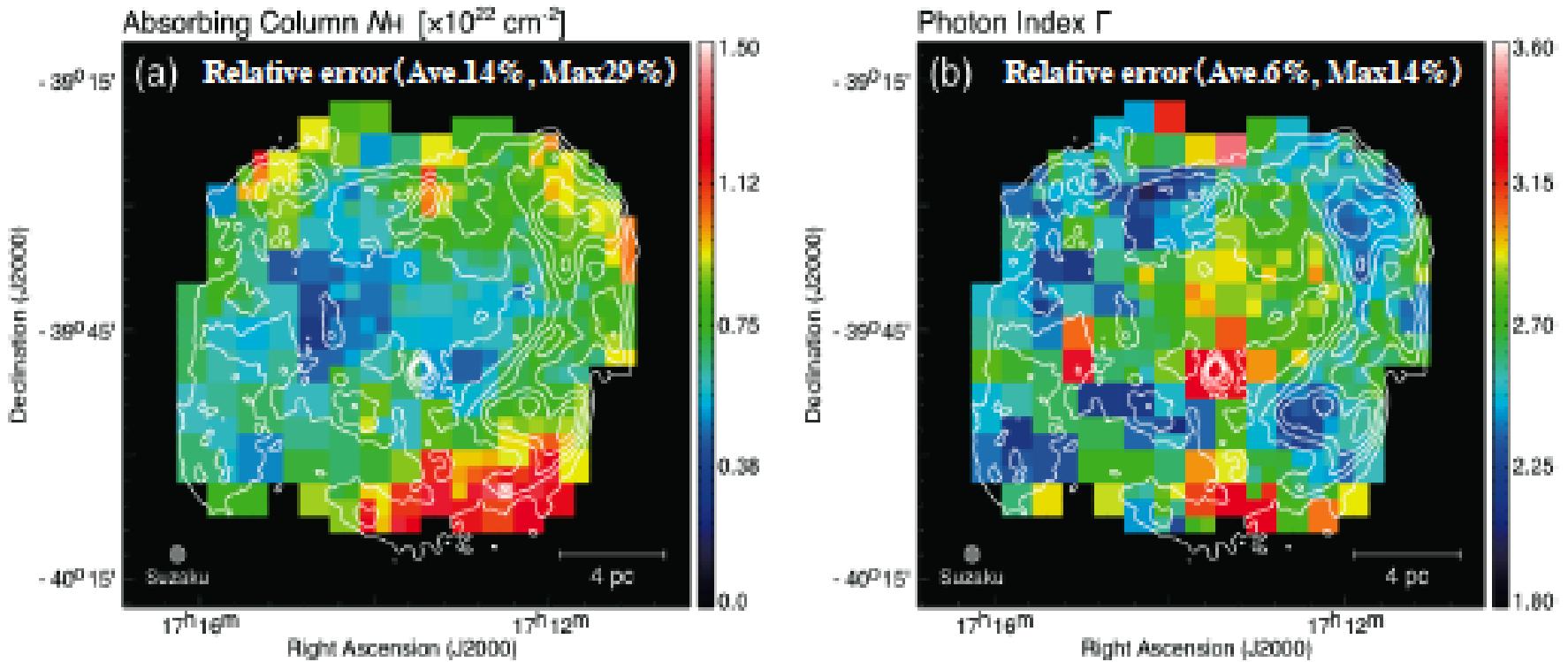


Image: Suzaku XIS 0+2+3 count map (1-5 keV, square root scale).

Contours: TeV gamma-rays (lowest; 20 smoothed counts, interval; 10 smoothed counts)

Absorbing Column & Photon Index



Color Image: (a) Absorbing column N_{H} [$\times 10^{22} \text{ cm}^{-2}$], (b) Photon index Γ

Contours: *Suzaku XIS 0+2+3 1-5* (square root scale)

RX J0852.0–4622 Azimuthal Plot (TeV gamma-rays and X-rays)

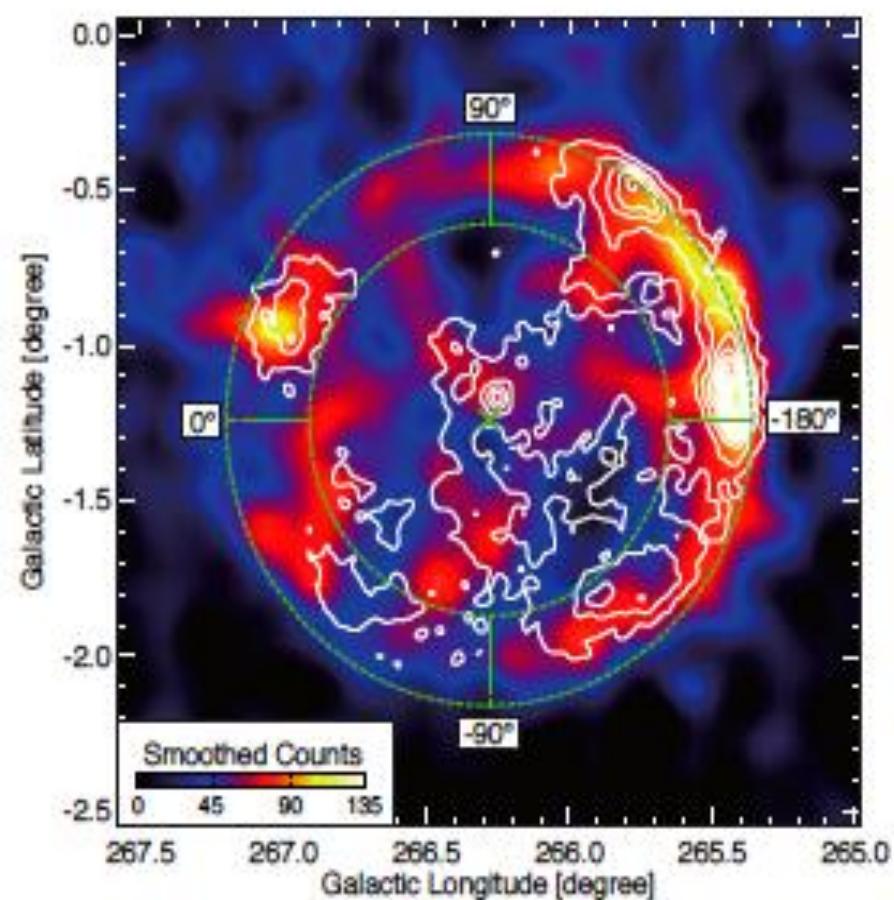
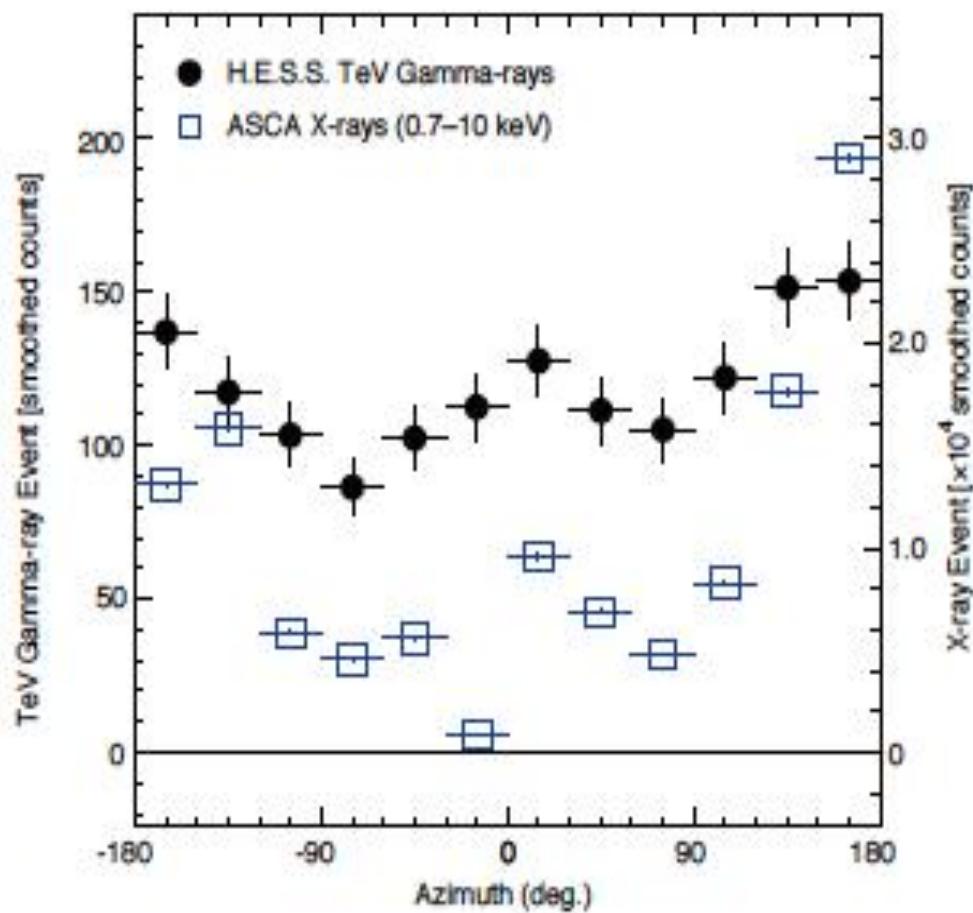


Image: H.E.S.S. TeV gamma-rays (Aharonian et al. 2007)
Contours: ASCA X-rays (0.7–10 keV; Tunemi et al. 2000)

Azimuthal plot (Fukui et al. 2012 in prep.)
center: $(l, b) = (266.28 \text{ deg.}, -1.24 \text{ deg.})$
radius: 0.63 deg. and 0.92 deg



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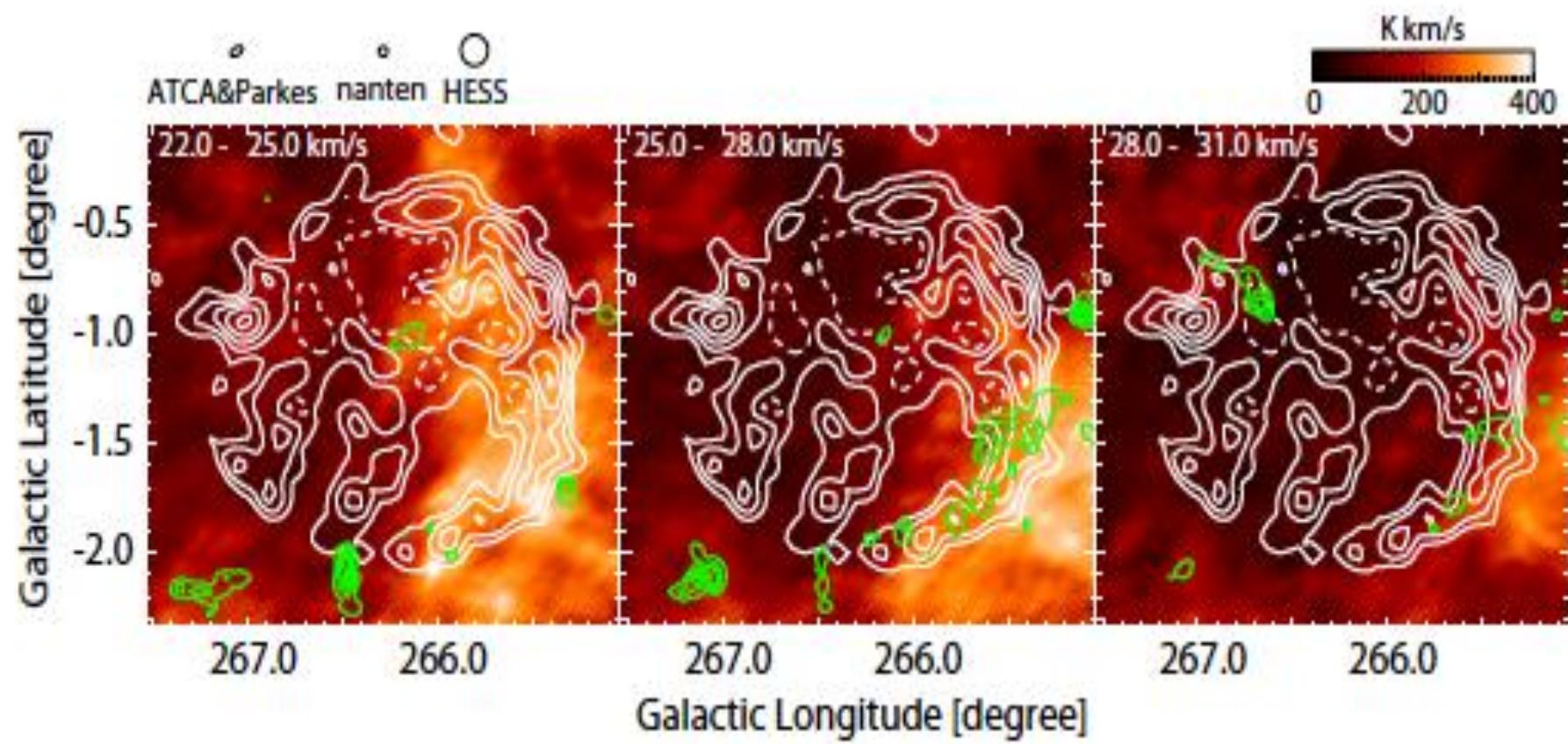
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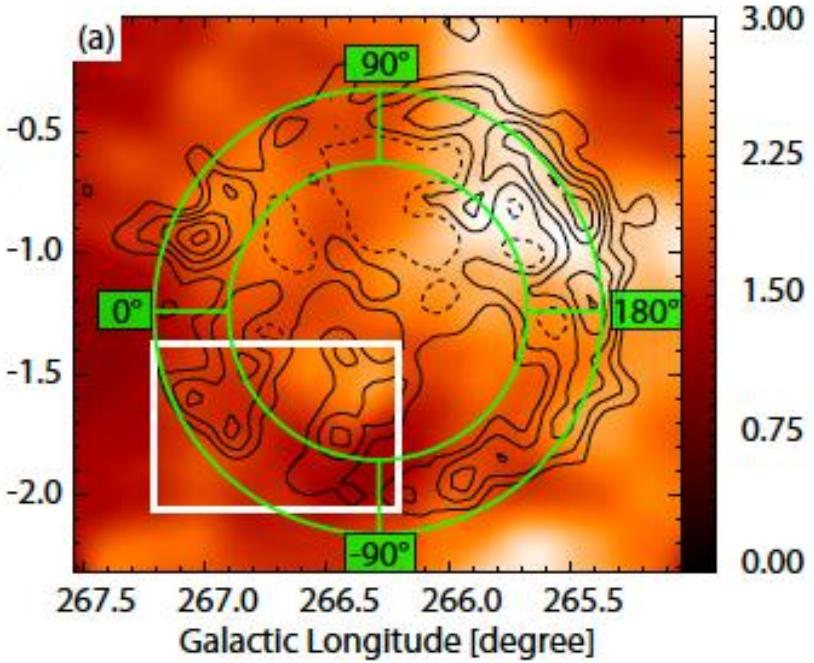
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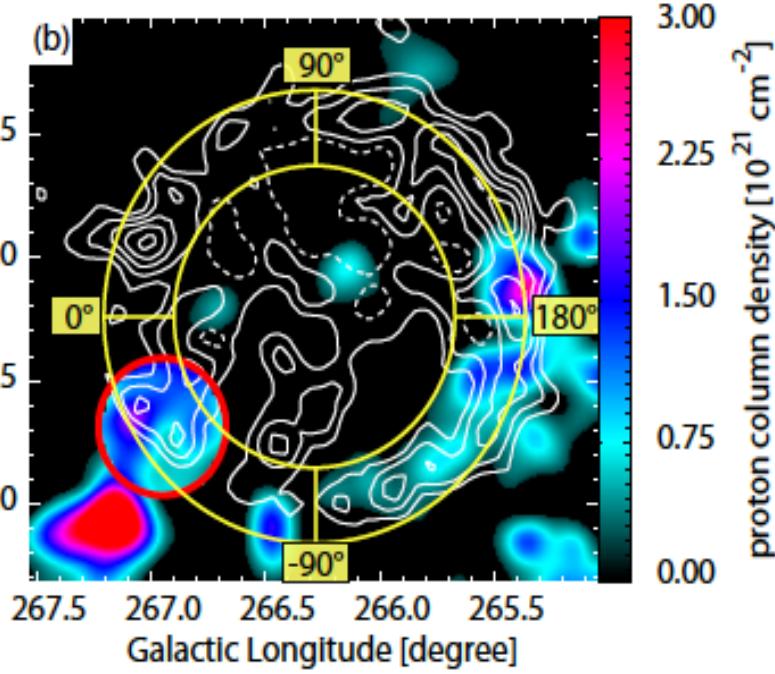
Galactic Latitude [degree]



267.5 267.0 266.5 266.0 265.5
Galactic Longitude [degree]

3.00
2.25
1.50
0.75
0.00

Galactic Latitude [degree]



267.5 267.0 266.5 266.0 265.5
Galactic Longitude [degree]

3.00
2.25
1.50
0.75
0.00