		Printed by
Nov 09, 11 11:32	AGILE_data_release_note_v6.0.txt	Page 1/5
***** AGILE Public	Data Release Note v6.0 ************************************	
User's README AGILE Data Center (ADC), November	9, 2011	
The proprietary period for 19 AGIL OB 8600 to OB 10400 (from 2009-11- and the data are now public and av. http://www.asdc.asi.it/mmia/ for t	E Cycle-3 Observation Blocks (OBs), from 30 to 2010-10-15), has currently expired, ailable from the ASDC Multimission Archive webpage he AGILE Mission.	
AGILE Cycle-3 data were obtained w and they have been processed with the new observing mode. Delivered were processed with the OB pipelin	ith AGILE observing a large portion of the sky in spinning mode, the latest available software and calibrations optimized for Cycle-3 spinning data (file type: STD1) e software version 5_21_18_19.	
The new public AGILE archive now c	ontains both Cycle-3 spinning, and Cycle-1 and Cycle-2 pointing data.	
POINTING: OB 4900 Õ 8400 During the previous Cycle-1 and Cy- characterized by long observations typically of 2-4 weeks duration, f The public archive for Cycle-1 and and Cycle-1 and Cycle-2 pointing d	cle-2 AGILE was operated in "pointing observing mode", called Observation Blocks, ollowing a predefined Baseline Pointing Plan. Cycle-2 OBs remains unchanged with respect to AGILE Public Data Release v5 ata are those reprocessed with the OB pipeline software version 5_19_18_17.	5.0,
SPINNING: SPINNING: OB >= 8600 On November 4, 2009, toward the en- a malfunction of the rotation whee in a spinning observing mode, i.e. sweeping the sky with an angular sp operating nominally producing data	d of Cycle2, AGILE scientific operations were reconfigured following l occurred in mid October, 2009. The satellite is currently operating with the solar panels pointing at the Sun and the instrument axis peed of ~ 0.8 degree/sec. The instrument and all the detectors are with quality equivalent to that obtained when operating in pointing mode.	
IMPORTANT: To avoid spurious artef analysed in a similar way to the ": to the following recommendations of	acts, the spinning dataset (file type: STD1) must be non-cleaned" pointing data (file type: STD0P) according n parameter values when generating maps:	
fovrad=50 albrad=90 phasecode	=2	
See explanation for the new file n	aming convention in the Data Retrieval section below.	
**************************************	***************************************	******
The AGILE observing mode was "spin AGILE Cycle-3 spinning observation each corresponding to a unique ide (OB 8600 - OB 8900) of variable du	ning" for all Cycle-3 OBs. s were structured as a series of 22 OBs with a "dummy" pointing direction, ntifing number and lasting ~15 days, except for the first four OBs ration.	
The set of GRID data which become provide the last three Cycle-3 OB (10500,	public today corresponds to the first 19 Cycle-3 observations. 10600, 10700), will be released at the end of their proprietary period on D	December 16, 2011.
Nodpoodov Novombor 00, 2011	ACILE data relaces note v6.0 tvt	
veunesuay Nuvernuer US, 2011	AGILE_UAIA_IEIEASE_IIUIE_VO.U.IXI	I/

										Printed by
No	ov 09, 11	11:32			AC	GILE_da	nta_release_note	_v6.0.txt		Page 2/5
Deli	vered Cy	cle-3 OBs ar	e listed i	below:						
#	OB #	OB Name	RA_PNT	DEC_PNT	RA_SUN	DEC_SUN	OB start date	OB end date	Mean OB Exp.	Notes
 1 2 3 4 5 6 7 8 9 10 11 12 13 14 15 16 17 18 19 The p due Notes whose	8600 8700 8800 9000 9100 9200 9300 9400 9500 9600 9700 9800 9900 10000 10100 10200 10300 10400 celestia rovide i: to solar that the	Dummy 1_1 Dummy 1_2 Dummy 1_3 Dummy 1_4 Dummy 2_1 Dummy 2_2 Dummy 2_3 Dummy 2_4 Dummy 3_1 Dummy 3_2 Dummy 3_3 Dummy 3_4 Dummy 4_1 Dummy 4_2 Dummy 4_3 Dummy 4_3 Dummy 5_1 Dummy 5_2 Dummy 5_3 Lowed is allowed is	<pre></pre>	<pre>mean point mean point osition (R 60 degree s pointing me so that</pre>	259 283.8 309.2 341.6 348.2 2 15.43 29.92 43.56 57.83 73.48 90.15 114.1 121.3 136.4 151.4 165.9 180 193.9 ing direct A_SUN, DEC s around direction	-20.33 -21.31 -16.14 -4.574 -1.802 3.405 9.475 15.6 19.48 23.39 25.76 26.25 25.71 22.98 19.02 14.38 8.307 2.478 -1.843 tion (RA_P: C_SUN), de the Sun an ns lie on b	<pre>2009-11-30 12:00:00 2009-12-20 12:00:00 2010-01-15 12:00:00 2010-02-05 12:00:00 2010-02-28 12:00:00 2010-03-15 12:00:00 2010-04-15 12:00:00 2010-04-15 12:00:00 2010-05-15 12:00:00 2010-05-15 12:00:00 2010-06-15 12:00:00 2010-06-30 12:00:00 2010-07-15 12:00:00 2010-07-15 12:00:00 2010-08-15 12:00:00 2010-08-15 12:00:00 2010-09-15 12:00:00 2010-09-15 12:00:00 2010-09-15 12:00:00</pre>	<pre>2009-12-20 12:00:00 2010-01-15 12:00:00 2010-02-05 12:00:00 2010-02-28 12:00:00 2010-03-15 12:00:00 2010-04-15 12:00:00 2010-04-15 12:00:00 2010-05-15 12:00:00 2010-05-30 12:00:00 2010-06-15 12:00:00 2010-06-30 12:00:00 2010-07-15 12:00:00 2010-08-15 12:00:00 2010-08-15 12:00:00 2010-09-15 12:00:00 2010-09-15 12:00:00 2010-09-15 12:00:00 2010-09-15 12:00:00 2010-09-15 12:00:00 2010-09-15 12:00:00 2010-10-15 12:00:00</pre>	14782732 17876769 13806594 15618616 11062229 12567574 10529106 9780278 12120621 11419530 11518075 11668613 12309031 12991280 12198474 13453435 12249951 12251493 12457466 ode. B	spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning spinning
* * * * * * * * * * * *	******* ** Data ******	**************************************	* * * * * * * * * * * * * * * * * * * *	*********	* * * * * * * * * *	*********	******************************	*****	*************	* * * * * *
The prod	query fo uces an	r the AGILE interactive	Mission d table sho	ata in the wing OBs s	ASDC Multelected ac	ti Mission ccording t	Interactive Archive h o the chosen "Search T	ttp://www.asdc.asi.it/n ype" option:	mmia/	
- se - se - se	arch by arch by arch by j	coordinates: time: the ob parameter: t	the prov servation he observ	ided source lies in t ation has	e position he specif: the select	n lies out ied time-: ted parame	side the forbidden reg range ters (OB number values	ions around the sun/an <sup>.</sup> )	ti-sun position	S
A qu resu	ery with lts in a	"Search Type complete tai	e" by "Ti ble of al	me", with a	default S GILE OB av	tart Date vailable t	and End Date values o date.			
The corr	Mean OB : esponds	Exposure col to the effec	umn (in c tive area	m2 s) in t associate	he intera d to the 1	ctive tabl FM filtere	e d events.			
The file	"Public a	access" link ch OB:	makes it	possible	to downloa	ad the fol	lowing			
* * *	SPINNING	data files	* * *							
- th ne ob	e spacec eded for servatio	raft auxilia the data an n within eac	ry (LOG) alysis an h OB, are	files (ag- d covering grouped u	<tstart>_; approximander the o</tstart>	STD1.LOG.g ately one- directory	z) day of STD1_LOG:			

AGILE\_data\_release\_note\_v6.0.txt

## AGILE\_data\_release\_note\_v6.0.txt

* STD1_LOG ag- <tstart1>_STD1.LOG.gz ag-<tstart2>_STD1.LOG.gz</tstart2></tstart1>					
 ag- <tstartn>_STD1.LOG.gz and a single general LOG index file with suffix: ag-STD1_LOG-<tstart1>.index WARNING: LOG files have a very accurate time resolution of 100 msec and need several GB of available disk space.</tstart1></tstartn>					
- two event files including all gamma-ray events in the GRID Field of View (FoV) using two AGILE event filters:					
the standard OB event file ag <tstart>-<tstop>_STD1_FM.EVT and its index file with suffix: .index the additional OB event file ag<tstart>-<tstop>_ST1_FT3AB.EVT and its index file with suffix: .index</tstop></tstart></tstop></tstart>					
WARNING: data analysis with event files obtained with FT3AB filter may be more efficient in detecting sources with a soft energy spectrum, but there may be noisy residual artifacts at the border of the Field of View (off-axis angle> 40 deg).					
<ul> <li>three Aitoff maps, count, exposure and diffuse background, centered on the Galactic Center: ag<tstart>-<tstop>_STD1_FM.COUNTS ag<tstart>-<tstop>_STD1_FM.EXP ag<tstart>-<tstop>_STD1_FM.GAS</tstop></tstart></tstop></tstart></tstop></tstart></li> </ul>					
These maps were automatically generated with the FM filter with the following parameters:					
mdim=360.0       index=-2.1       roll_tol=360.0         mres=0.3       fovrad=50       earth_tol=5.0         lonpole=180       fovradmin=0.0       keepmono=NO         emin=100       albrad=90       phasecode=2         emax=50000       y_tol=0.5       projection=AIT         timestep=160       build=0       step=3					
To produce your own maps and run likelihood tasks please download and install the public AGILE software available at: http://agile.asdc.asi.it/public/ and follow the User Manual included.					
<ul> <li>- a GIF file showing the images of both the OB (FM) exposure and count maps.</li> <li>For illustrative purpose only, the count map image includes automatic</li> <li>candidate detections in the FoV obtained with XIMAGE software.</li> </ul>					
*** POINTING cleaned data files ***					
<ul> <li>the spacecraft cleaned auxiliary (LOG) files (ag-<tstart>_STD0P_GO.LOG.gz) needed for the data analysis and covering approximately one-day of observation within each OB, are grouped under the directory STD0P_GO_LOG:</tstart></li> </ul>					
* STD0P_G0_LOG ag- <tstart1>_STD0P_G0.LOG.gz ag-<tstart2>_STD0P_G0.LOG.gz</tstart2></tstart1>					
ag- <tstartn>_STD0P_G0.LOG.gz ag-<tstartn>_STD0P_G0.LOG.gz and a single general LOG index file with suffix: ag-STD0P_G0_LOG-<tstart1>.index WARNING: LOG files have a very accurate time resolution of 100 msec and need several GB of available disk space.</tstart1></tstartn></tstartn>					
- two cleaned event files including all gamma-ray events in the GRID Field of View (FoV) using AGILE event filters:					

AGILE\_data\_release\_note\_v6.0.txt

			Printed by
Nov 09, 11 11	1:32	AGILE_data_release_note_v6.0.txt	Page 4/5
the stand the addit	dard OB event file ional OB event fil	ag <tstart>-<tstop>_STD0P_GO_FM.EVT and its index file with suffix: .index le ag<tstart>-<tstop>_STD0P_GO_FT3AB.EVT and its index file with suffix: .index</tstop></tstart></tstop></tstart>	
WARNING: data a soft energy spe (off-axis angle	analysis with even ectrum, but there of e> 40 deg).	t files obtained with FT3AB filter may be more efficient in detecting sources with a may be noisy residual artifacts at the border of the Field of View	
- three maps, c ag <tstart>-<i ag<tstart>-<i ag<tstart>-<i< td=""><td>count, exposure and CSTOP&gt;_STDOP_GO_FM CSTOP&gt;_STDOP_GO_FM CSTOP&gt;_STDOP_GO_FM</td><td>d diffuse background, centered on the mean OB pointing position: .COUNTS .EXP .GAS</td><td></td></i<></tstart></i </tstart></i </tstart>	count, exposure and CSTOP>_STDOP_GO_FM CSTOP>_STDOP_GO_FM CSTOP>_STDOP_GO_FM	d diffuse background, centered on the mean OB pointing position: .COUNTS .EXP .GAS	
These maps were	e automatically gen	nerated with the FM filter with the following parameters:	
mdim=120.0 mres=0.25 lonpole=180 emin=100 emax=50000	<pre>index=-2.1 fovrad=70 albrad=80 y_tol=0.5 roll_tol=360.0</pre>	<pre>earth_tol=3.0 keepmono=NO phasecode=18 projection=ARC step=4</pre>	
To produce your install the pub http://agile.as and follow the	r own maps and run blic AGILE software sdc.asi.it/public/ Software User Man	likelihood tasks please download and e available at: ual included.	
- a GIF file sh For illustrat candidate det	nowing the images of tive purpose only, tections in the Fo	of both the OB (FM) exposure and count maps. the count map image includes automatic V obtained with XIMAGE software.	
*** POINTING nc	on-cleaned data fi	les ***	
- the spacecraf needed for th observation w	t auxiliary (LOG) he data analysis an within each OB, are	files (ag- <tstart>_STD0P.LOG.gz) nd covering approximately one-day of e grouped under the directory STD0P_LOG:</tstart>	
* STDOP_LOG ag- <tsi ag-<tsi< td=""><td>; TART1&gt;_STD0P.LOG.g: TART2&gt;_STD0P.LOG.g:</td><td>z z</td><td></td></tsi<></tsi 	; TART1>_STD0P.LOG.g: TART2>_STD0P.LOG.g:	z z	
and a single	CARTn>_STD0P.LOG.g: general LOG index	z file with suffix: ag-STD0P_LOG- <tstart1>.index</tstart1>	
- two event fil	es including all g	gamma-ray events in the GRID Field of View (FoV) using AGILE event filters:	
the stand the addit	dard OB event file cional OB event fil	ag <tstart>-<tstop>_STD0P_FM.EVT and its index file with suffix: .index le ag<tstart>-<tstop>_STD0P_FT3AB.EVT and its index file with suffix: .index</tstop></tstart></tstop></tstart>	
NOTE: if you ch then each file	noose to download on name in the correst	data files with the default option "Automatically unpack the data using a Java applet" sponding .index file must be changed removing the .gz suffix before running map generator	tasks.
***** On-line A ***** On-line A	Analysis Interactiv	**************************************	

Wednesday November 09, 2011

AGILE\_data\_release\_note\_v6.0.txt

		Printed by
Nov 09, 11 11:32	AGILE_data_release_note_v6.0.txt	Page 5/5
A new beta version of the ASDC interact public data fields and to perform an in The tool is in a verification phase. PI WARNING: use only as a preview of the A To perform your own scientific analysis	tive tool allows Web users to preview the AGILE nteractive preliminary analysis around a chosen sky position. lease send your feedback at the AGILE helpdesk. AGILE gamma-ray field. s, please download data and use the official public AGILE software.	
To access the preview tool click on "Or column. The interactive ASDC tool uses adapted to gamma-ray image display and	n-line Analysis" in the query output table, under the "GRID Interactive Archi the XIMAGE software package for multi-mission X-ray astronomy (v4.5.1), data analysis.	ve"
**************************************	***************************************	******
The standard AGILE event filter is the The AGILE Team recommmends the use of v gamma-ray source standard likelihood an	FM filter, optimized up to 60 deg off-axis. v3.0 EVT files produced with the FM filter for nalysis.	
EVT files produced with the FT3AB filte WARNING: data analysis with event files soft energy spectrum, but there may be (off-axis angle> 40 deg).	er are also published, but they are delivered with a warning. s obtained with FT3AB filter may be more efficient in detecting sources with noisy residual artifacts at the border of the Field of View	a
Each filter is associated with its own Be careful always to use the calibratic Calibration and diffuse model files as public software release under the direc	calibration and diffuse model files. on and diffuse model files appropriate to the chosen EVT filter type. re delivered with the ctory \$ADC/scientific analysis/data.	
* FM filter: if you use the FM-filtered you should always use as an input to	d event file with suffix: _FM scientific software tasks:	
<ul> <li>effective area files of type *FM*.sar</li> <li>energy dispersion files of type *FM*.</li> <li>point spread function files of type *FM*.ex</li> <li>flux correction files of type *FM*.ex</li> <li>diffuse model files of type *FM*.com</li> </ul>	r.gz .edp.gz FM*.psd.gz xpcorr.gz v.sky.gz	
* FT3AB filter: if you use the FT3AB-fi you should always use as an input to	iltered event file with suffix: _FT3AB scientific software tasks:	
<ul> <li>effective area files of type *FT3AB*.</li> <li>energy dispersion files of type *FT3A</li> <li>point spread funcion files of type *F</li> <li>flux correction files of type *FT3AB*.</li> <li>diffuse model files of type *FT3AB*.</li> </ul>	.sar.gz AB*.edp.gz FT3AB*.psd.gz *.expcorr.gz conv.sky.gz	
For further details please follow the state the Software User Manual.	instructions given in	
 Enjoy!		

Wednesday November 09, 2011